



Elliptic flow of light nuclei in Au+Au collisions at $\sqrt{s_{NN}}$ = 14.6, 19.6, 27, and 54.4 GeV using the STAR detector

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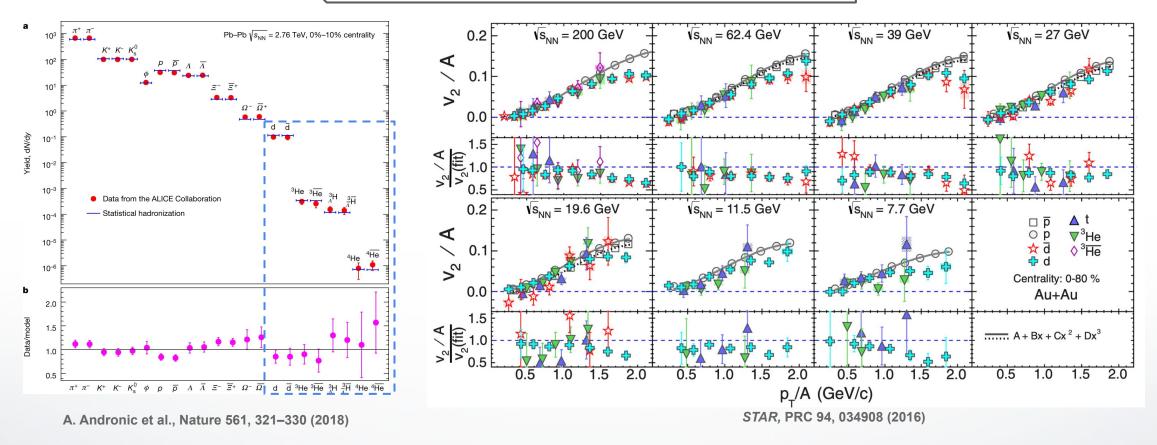


Outline

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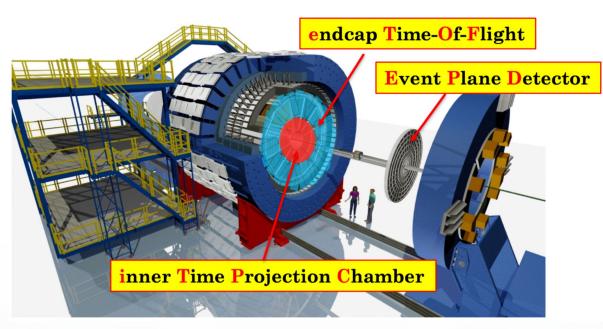
Motivation



- ★ Light nuclei production in heavy-ion collisions can be explained either by the **thermal** model or the final-state **coalescence** of nucleons
- \star v₂/A of light nuclei was observed to be close to v₂ of protons for p_T/A < 1.5 GeV/c in BES-I data
- ★ Higher statistics dataset in BES-II program will allow us to revisit and better understand the production mechanism of light nuclei



The STAR Experiment



C. Yang et al., JINST 15 C07040 (2020)

- ★ Particle identification is performed using
 - \circ dE/dx information from Time Projection Chamber (TPC)
 - \circ m^2 information from **Time of Flight (TOF)**
- ★ BES-II upgrades:
 - iTPC: Large pseudorapidity coverage ($|\eta| < 1.5$)
 - Better track and event plane resolution
- **★** Datasets
 - BES II: Au+Au collisions at $\sqrt{s_{NN}}$ = 14.6, 19.6, 27, and 54.4 GeV



Analysis details

★ The particle azimuthal distribution can be written as:

$$E\frac{d^3N}{d^3p} = \frac{1}{2\pi} \frac{d^2N}{p_T dp_T dy} \left(1 + \sum_{n=1}^{\infty} 2\nu_n \cos(n(\phi - \Psi_R))\right) \qquad v_n = \left\langle \cos[n(\phi - \psi_R)]\right\rangle$$

★ nth harmonic plane is calculated using the Q-vector:

$$Q_n \cos(n\Psi_n) = \sum_i w_i \cos(n\phi_i)$$

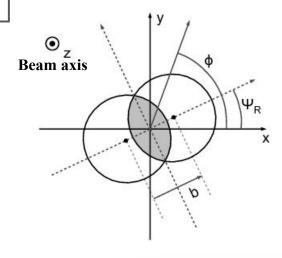
$$Q_n \sin(n\Psi_n) = \sum_i w_i \sin(n\phi_i)$$

$$\Psi_n = \left(\tan^{-1} \frac{\sum_i w_i \sin(n\phi_i)}{\sum_i w_i \cos(n\phi_i)}\right) / n$$

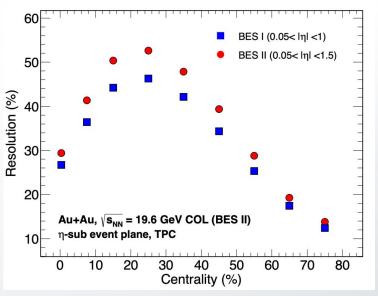
To suppress non-flow contributions

- \star η-sub event plane is used
- \star η -gap of 0.1 is taken between two subevents





CMS, PRC 87 014902 (2013)



We observe an improvement of resolution by ~10% from BES I owing to higher <u>TPC</u> acceptance and <u>track resolution</u>

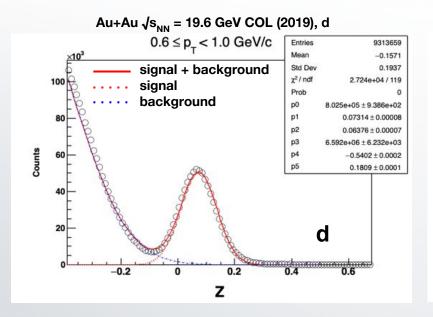


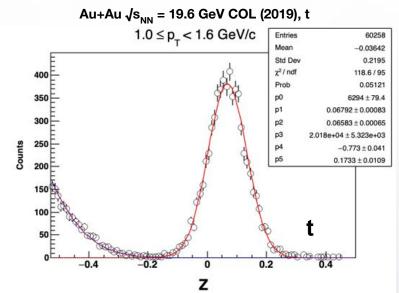
Particle identification

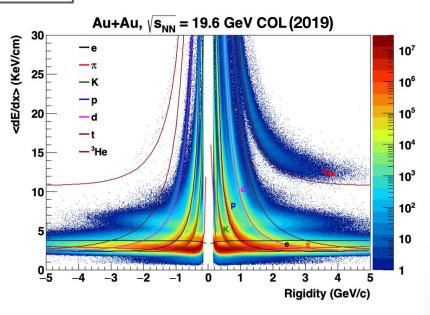
Particles are identified using dE/dx information from TPC in the range $|\eta| \le 1.0$

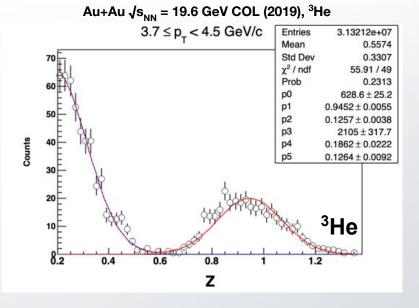
$$z_i = \ln \left(\frac{\langle dE/dx \rangle_{measured}}{\langle dE/dx \rangle_{theory}} \right)$$

- \star <dE/dx>_{theory} is calculated using Bichsel function
- ★ Double Gaussian fit is done to calculate yield in each p_T and $\phi \Psi_2$ bin



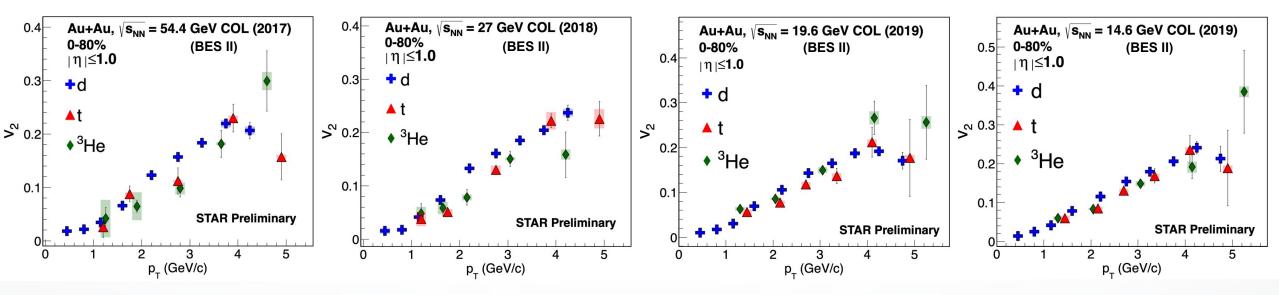








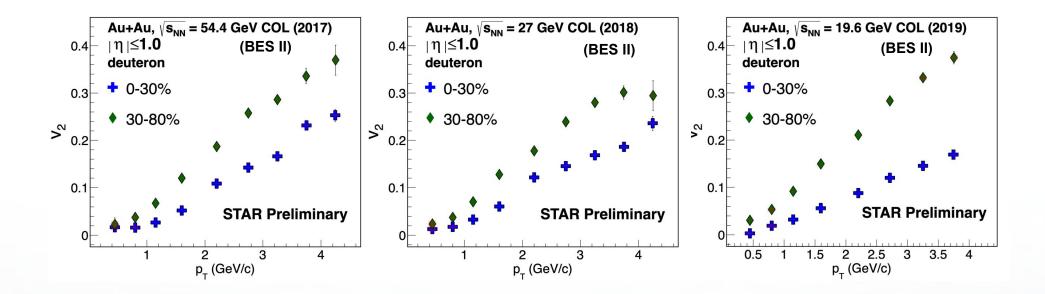
Elliptic flow of light nuclei



 \star The $v_2(p_T)$ for all nuclei species increases with increasing p_T for all collision energies



Centrality dependence

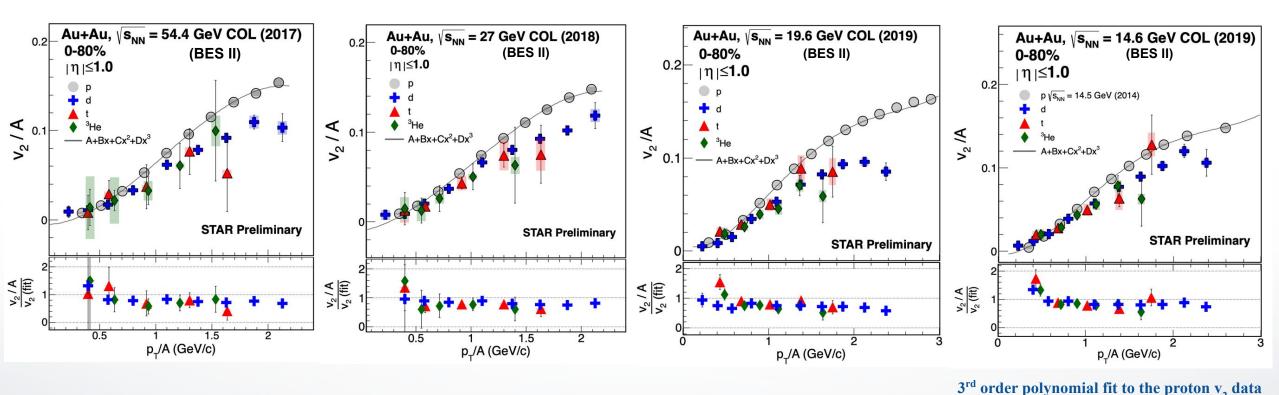


 \star v₂ of deuterons shows a strong centrality dependence

• Peripheral collisions have relatively larger v₂ due to their larger initial spatial anisotropy



Mass number scaling



 \star v₂/A of light nuclei is compared to the v₂ of protons to test mass number scaling

 \star v₂ of light nuclei obeys the mass number scaling within 20-30%



Summary

- \star v₂ of d, t, and ³He is presented in Au+Au collisions at $\sqrt{s_{NN}}$ = 14.6, 19.6, 27 and, 54.4 GeV (Collider)
 - v₂ shows a clear centrality dependence
 - Light nuclei v₂ seems to be obeying mass number scaling within 20-30%

Outlook

★ More exciting results to follow on light nuclei flow from BES-II energies



Thank you