



Elliptic flow of light nuclei in heavy ion collisions at STAR

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Outline

➤ Introduction

➤ Analysis Method

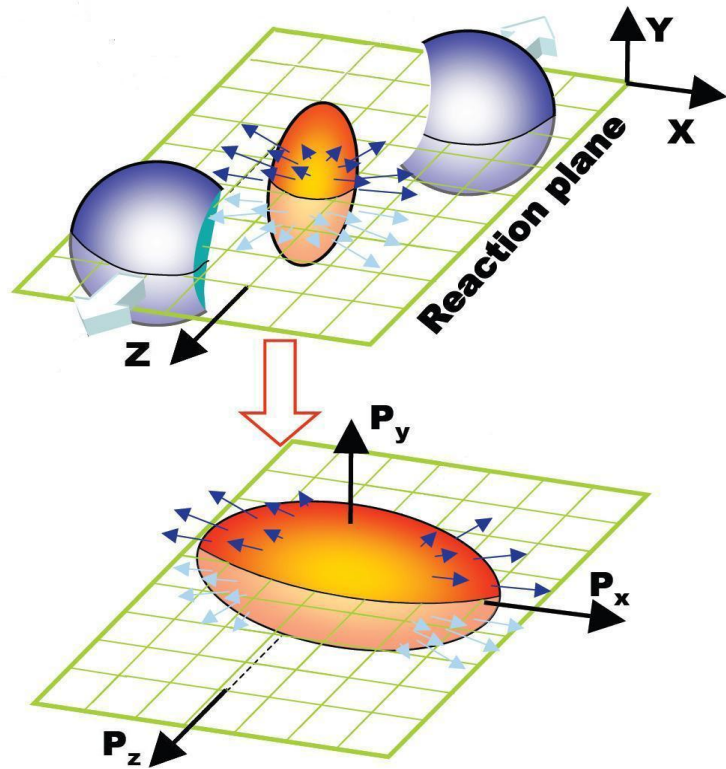
➤ Results and Discussions

- ✓ NCQ Scaling
- ✓ Centrality Dependence
- ✓ Mass Dependence
- ✓ Energy Dependence

➤ Summary



Introduction



$$\varepsilon = \frac{\langle y^2 - x^2 \rangle}{\langle y^2 + x^2 \rangle}$$

Coordinate-Space
Anisotropy



Momentum-Space
Anisotropy

$$E \frac{d^3 N}{d^3 p} = \frac{1}{2\pi} \frac{d^2 N}{p_T dp_T dy} \left(1 + 2 \sum_{n=1}^{\infty} v_n(p_T, y) \cos(n(\phi - \Psi_r)) \right)$$

$$v_n = \langle \cos(n(\phi - \Psi_r)) \rangle \quad \phi = \tan^{-1} \left(\frac{p_y}{p_x} \right)$$

➤ Elliptic flow (v_2) is a good probe of the early stage of the collision.



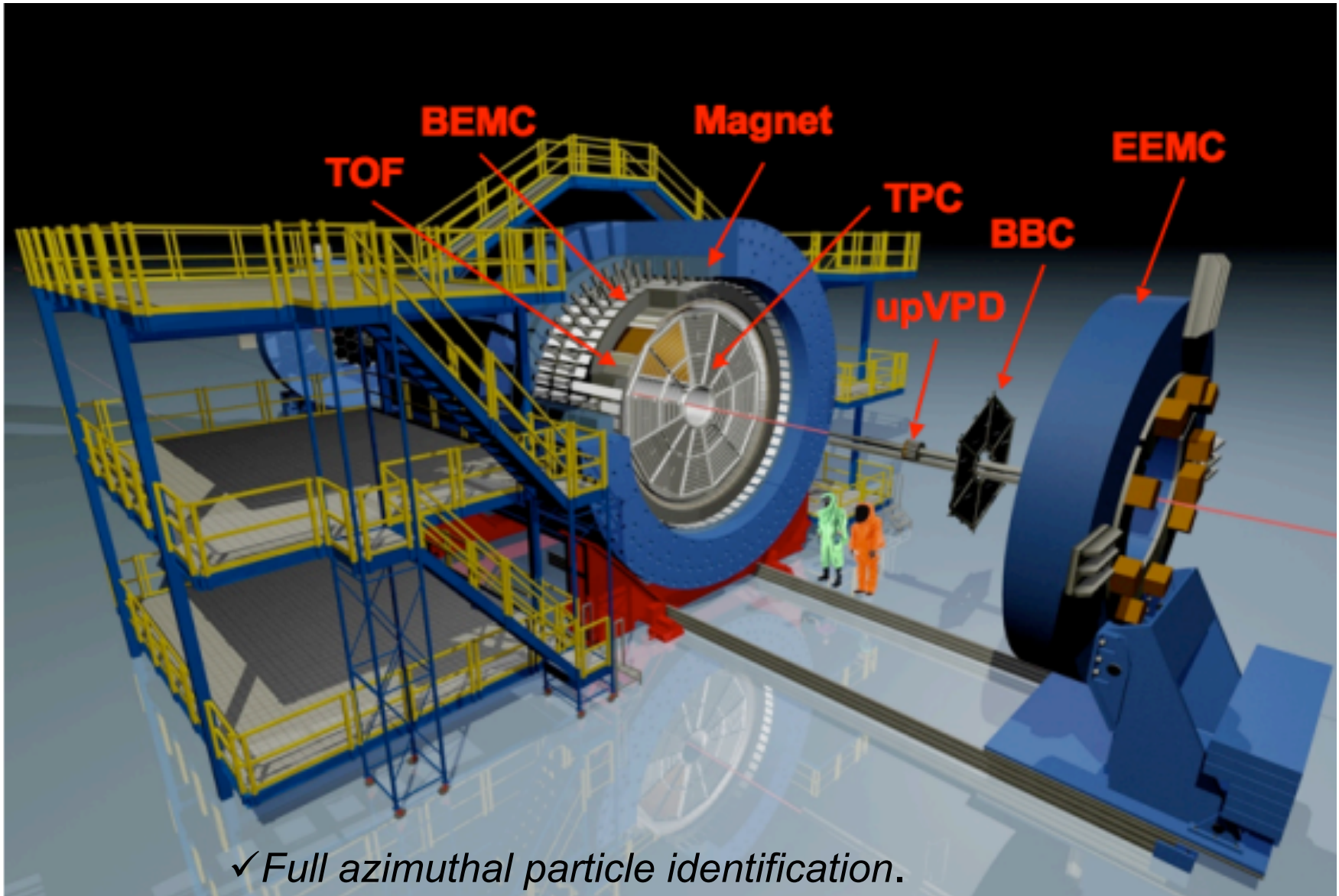
Motivation

- Light nuclei and anti-nuclei are formed through coalescence of nucleons (produced/participant) and anti-nucleons^{*}. This formation process is generally believed to happen at a late stage of the evolution due to their small binding energy.
- Production of light nuclei provides information of space-momentum correlation among these nucleons.
- By studying the v_2 of light nuclei and comparing to that of their constituent nucleons, we will have a better understanding of coalescence process for hadronization.

^{*}H.H. Gutbrod et al., Phys. Rev. Lett. 37, 667 (1976).



STAR Detector



✓ *Full azimuthal particle identification.*



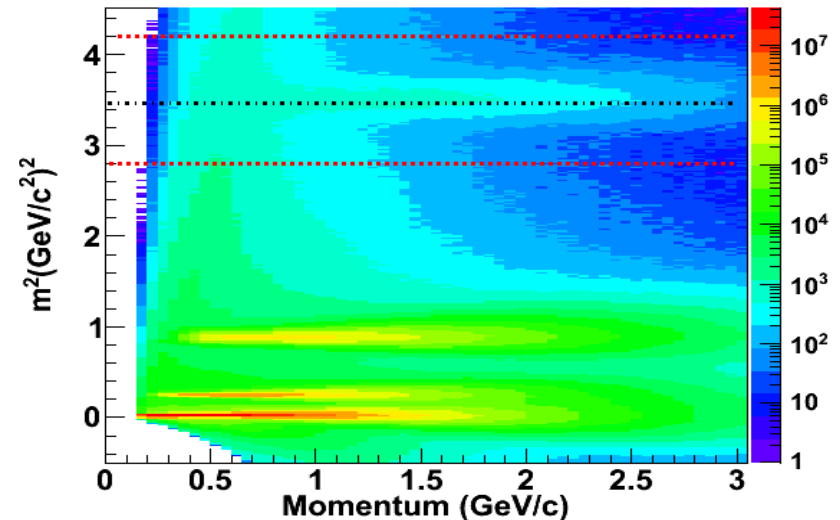
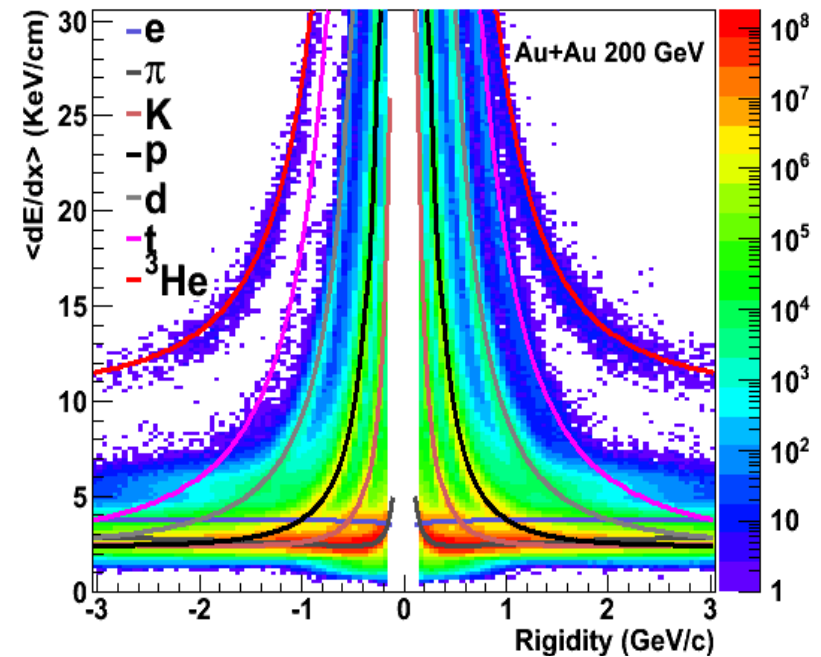
Data Set

Data : Run 7 Au+Au 200 GeV
Number of events: ~62 M

Data : Run 10 Au+Au 39 GeV
Number of events: ~14 M

➤ Measurement of the ionization energy loss (dE/dx) of charged tracks in the TPC gas are used to identify the light nuclei.

➤ In run 10 Au+Au 39 GeV TOF information is used to enhance the light nuclei identification at high momentum region.

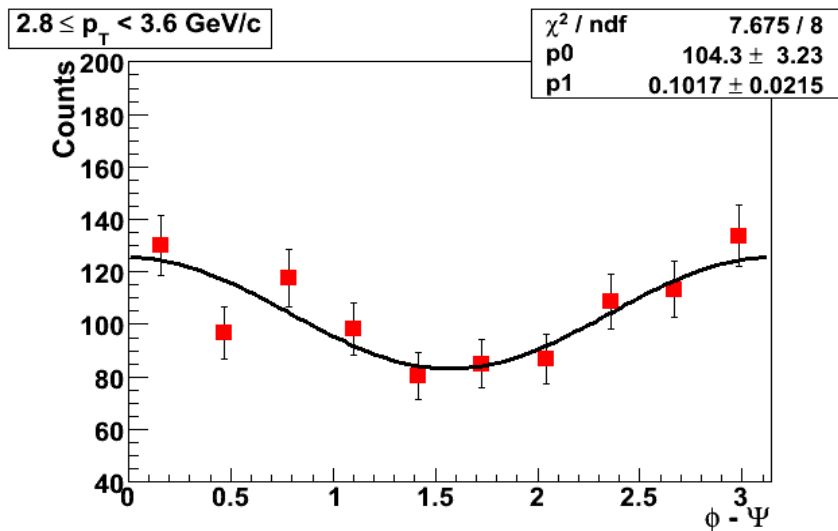
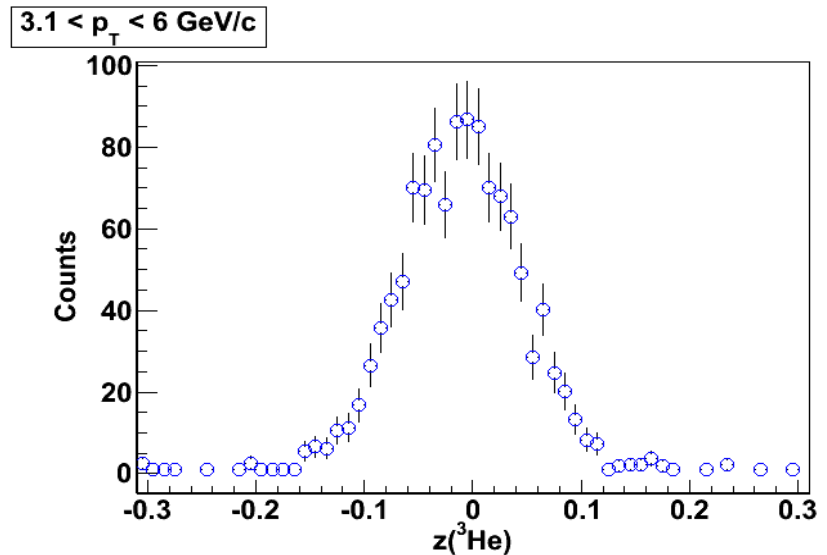


Analysis Method

➤ ^3He signal is almost background free for $p_T > 1.4 \text{ GeV}/c$.

$$Z = \ln\left(\frac{(dE/dx)_{\text{measure}}}{(dE/dx)_{\text{predict}}}\right)$$

➤ TPC is used to determine the event plane.



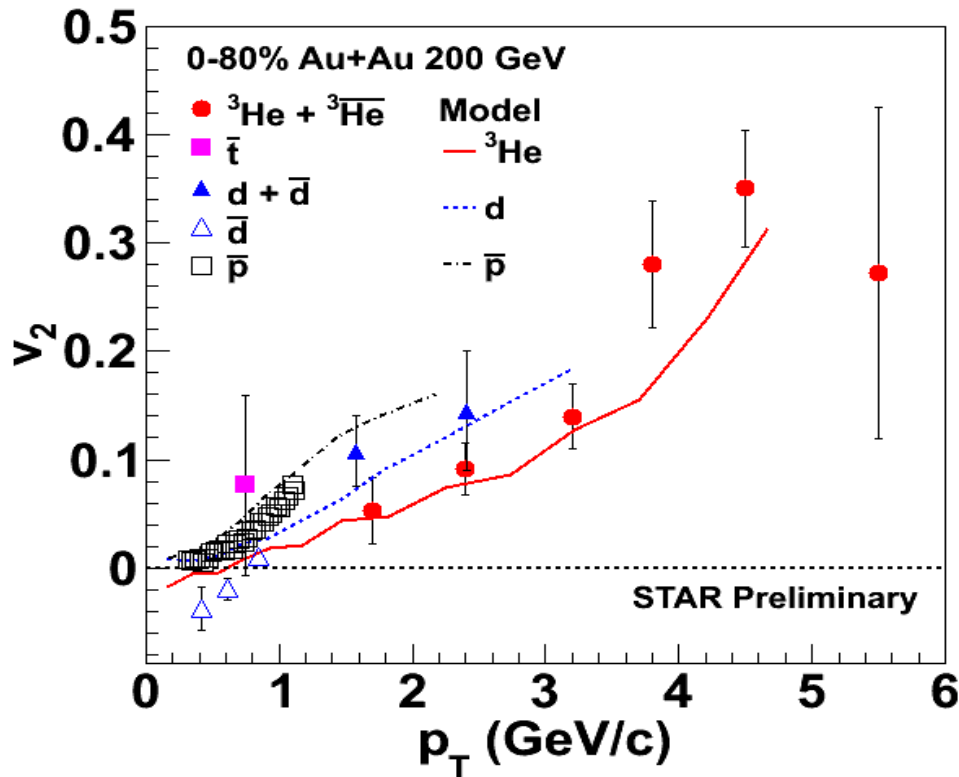
$$\frac{dN}{d\phi} \propto 1 + 2v_2^{obs} \cos[2(\phi - \Psi_R)]$$

$$v_2^{final} = \frac{v_2^{obs}}{R_2}$$

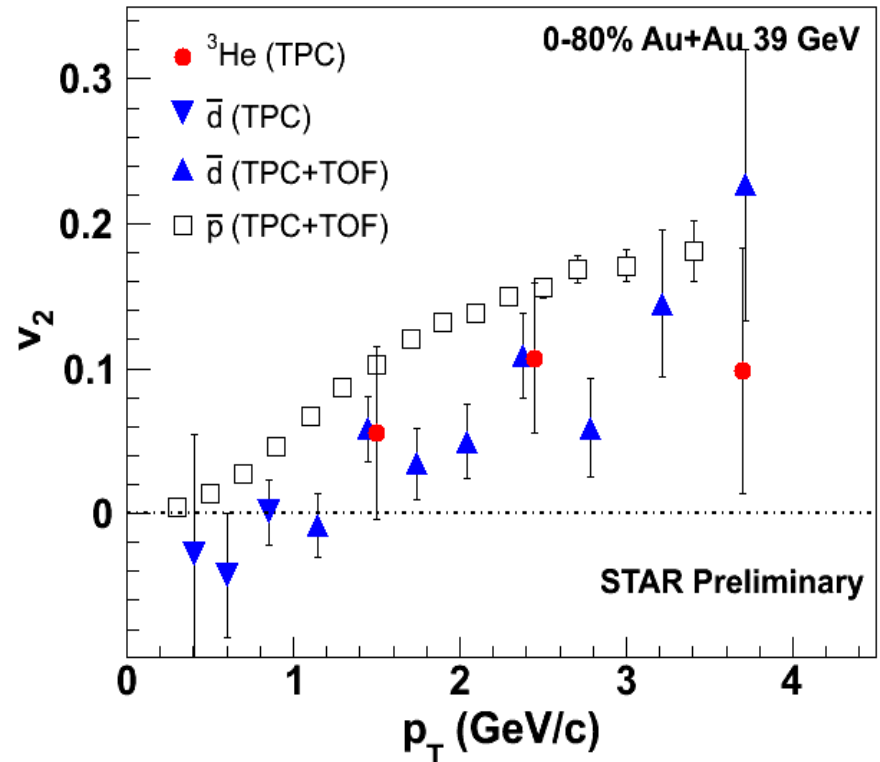
R_2 is the second order event plane resolution



V₂ VS. p_T



\bar{p} d(d) data: STAR, arXiv:0909.0566 [nucl-ex]



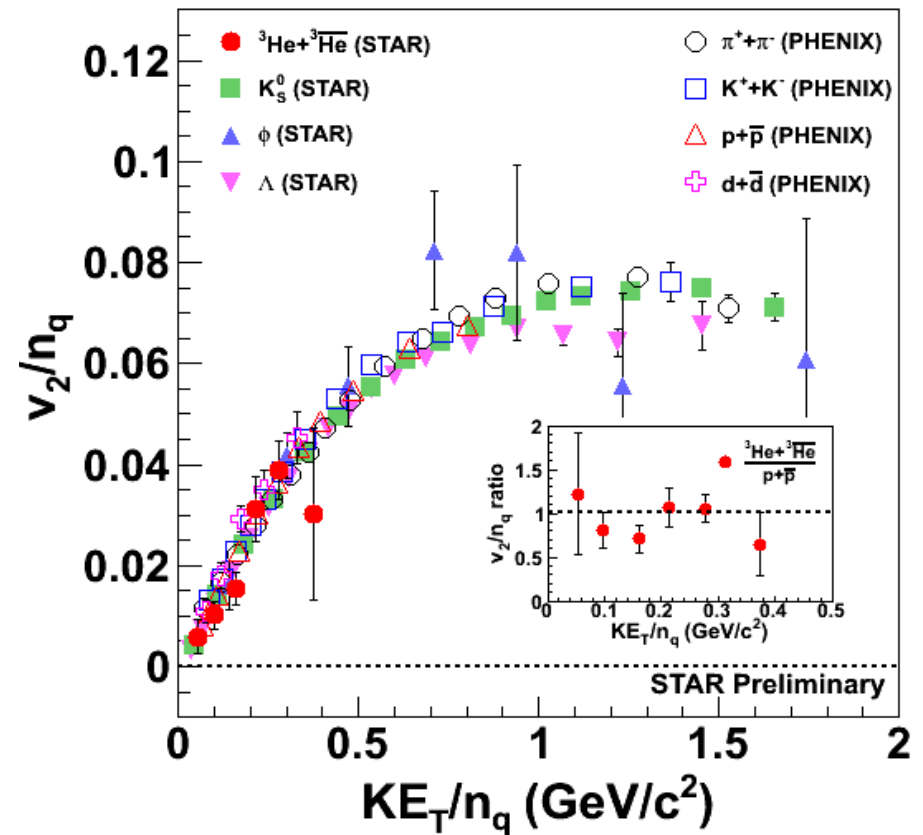
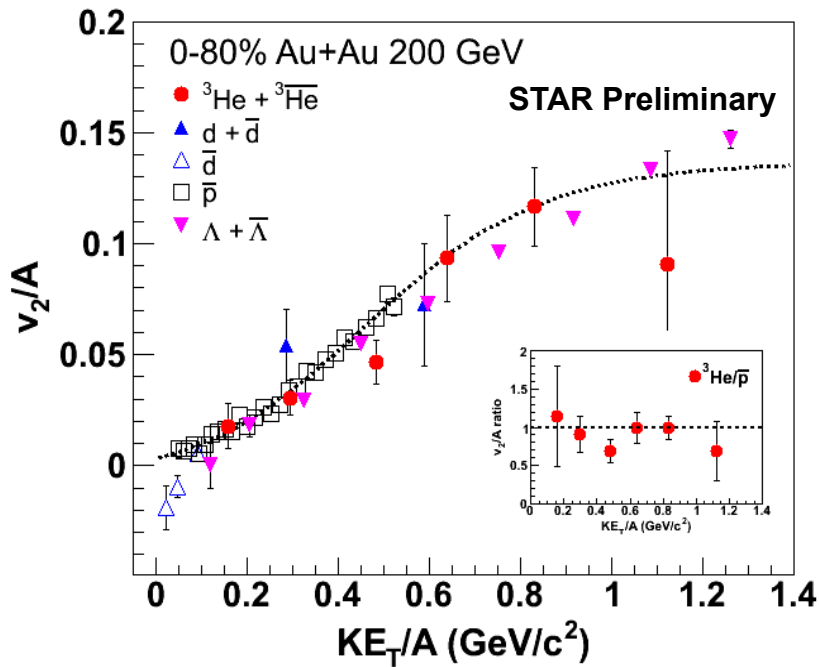
\bar{p} : A. Schmah (STAR), QM 2011

- v_2 of light nuclei increases with p_T in both beam energies.
- v_2 of light nuclei is well described by the dynamical coalescence model in Au+Au 200 GeV.

Model Calculations: S. Zhang et al., Phys. Lett. B 684 (2010) 224



NCQ Scaling: Au+Au 200 GeV



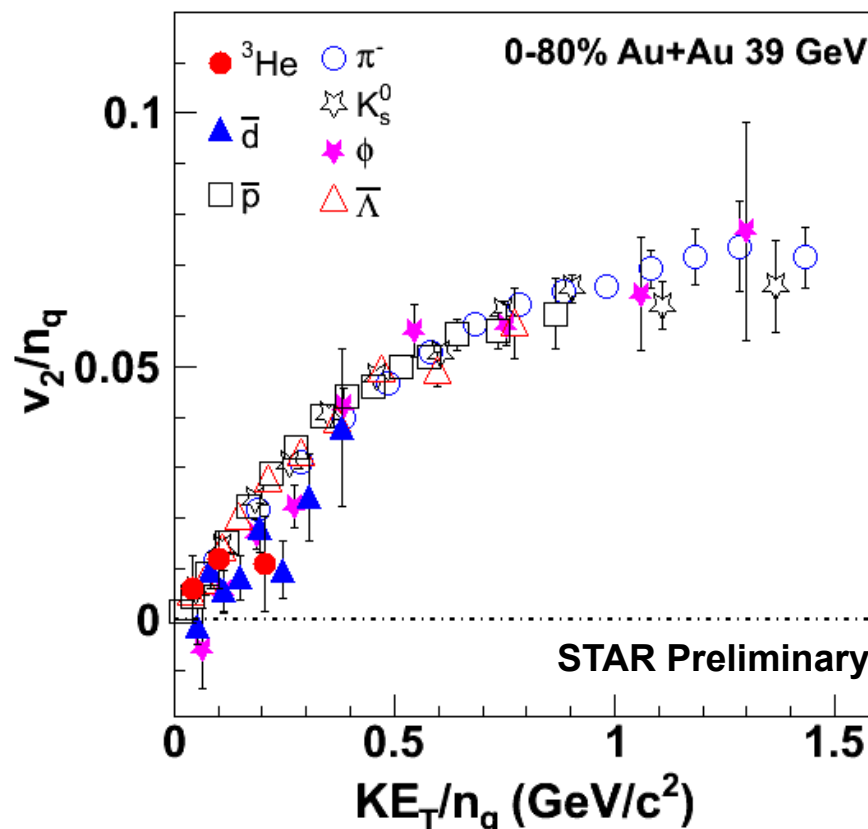
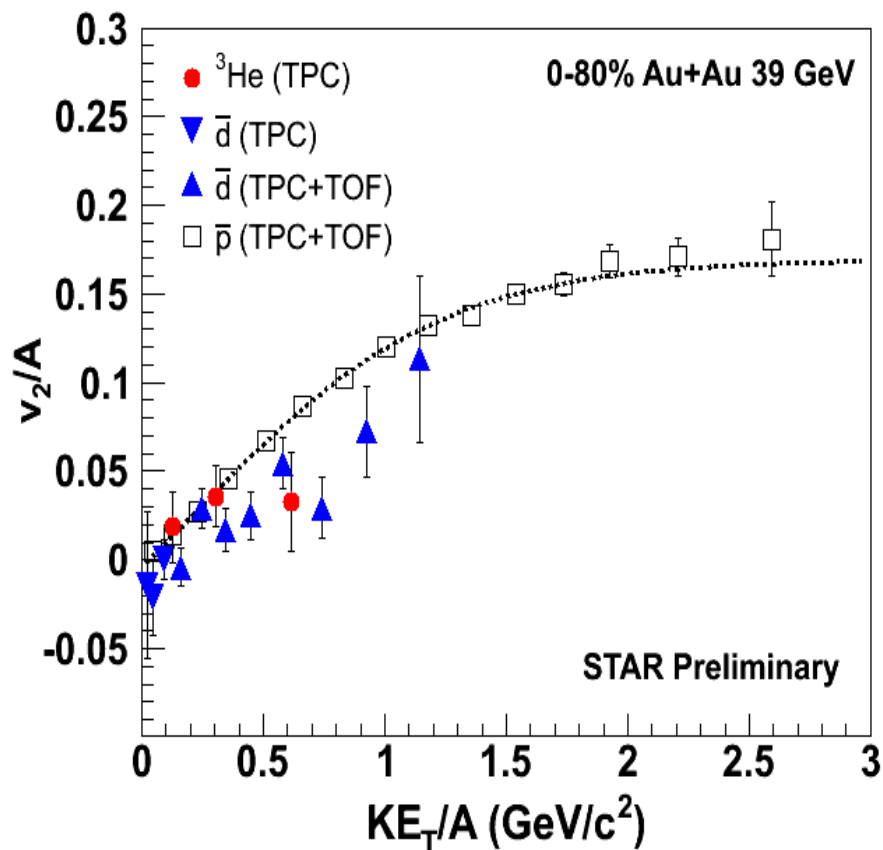
STAR, PRC 72, 014904 (2005)
 STAR, arXiv:0909.0566 [nucl-ex]
 PHENIX, PRL 99, 052301 (2007)

$$d(p+n) : n_q = 2 \times 3 \quad {}^3\text{He}(2p+n) : n_q = 3 \times 3$$

- v_2 of light nuclei follow the atomic mass number (A) scaling within errors.
- v_2 of light nuclei scaled to the number of constituent quarks (NCQ) of their constituent nucleons, are consistent with NCQ scaled v_2 of baryons and mesons.



NCQ Scaling: Au+Au 39 GeV

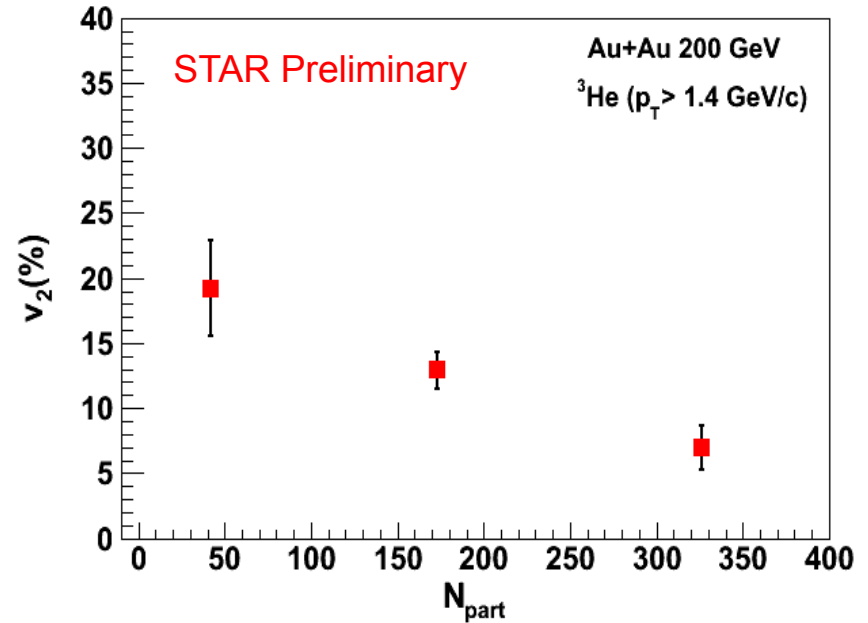
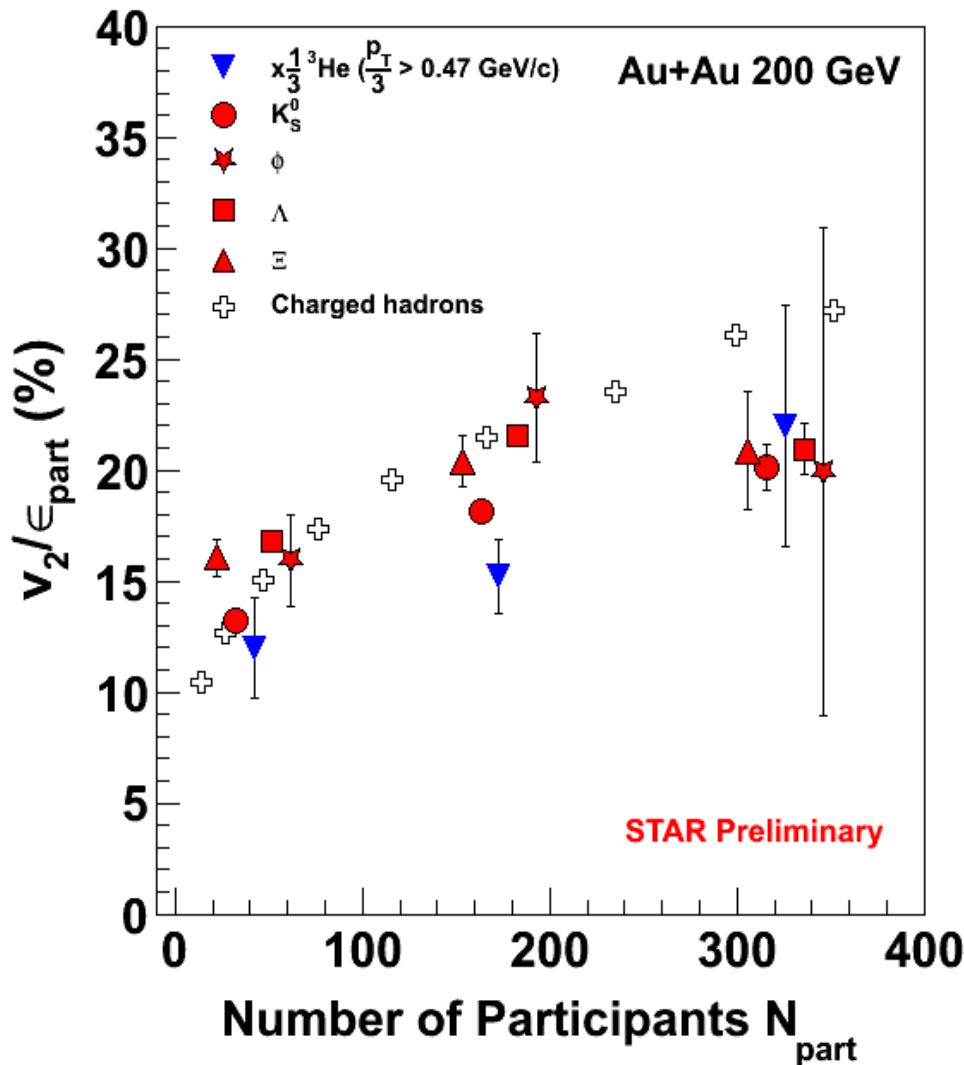


A. Schmah (STAR), QM 2011

- v_2 of light nuclei seem to follow the atomic mass number (A) scaling in Au+Au 39 GeV.
- NCQ scaling holds good for v_2 of light nuclei in Au+Au 39 GeV.



Centrality Dependence



➤ Similar to other hadrons, at more central collision the larger value of v_2/ϵ indicates stronger collective expansion.

STAR, PRC 77, 054901 (2008)

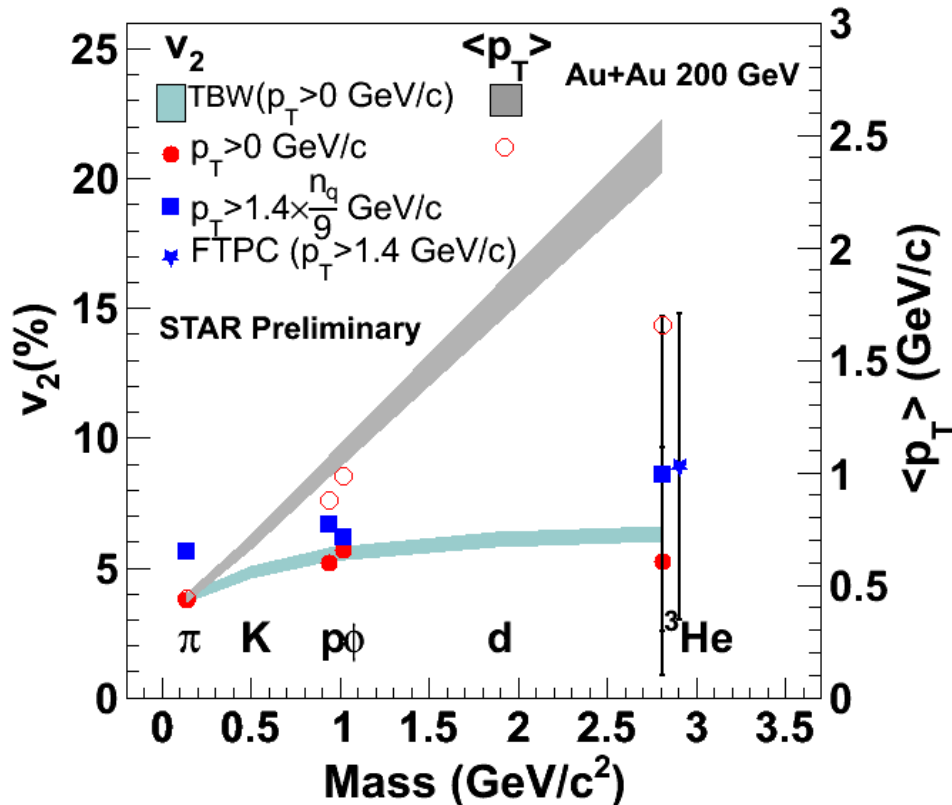


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Mass Dependence



➤ Tsallis Blast-wave (TBW) parameters obtained from the simultaneous fits to p_T spectra and v_2 of π , K and p are used to predict the p_T spectra and v_2 of d and ^3He .

➤ Both v_2 and $\langle p_T \rangle$ trends are consistent with expectations from TBW model fit.

➤ v_2 values up to ^3He mass has reasonable agreement within the Blast-wave formalism but differences seen in $\langle p_T \rangle$ beyond proton mass.

➤ v_2 of ^3He measured from different event plane detectors (TPC and FTPC) are consistent with each other.

✓ Systematic uncertainties are under study.

π , p spectra: STAR, PRL 97, 152301 (2006)

π , p v_2 : STAR, PRC 72, 014904 (2005)

STAR, PRC 77, 054901 (2008)

ϕ spectrum and v_2 : STAR, PRL 99, 112301 (2007)

^3He spectrum: STAR, arXiv:0909.0566 [nucl-ex]

TBW: Z. Tang et al., PRC 79, 051901 (2009)

Z. Tang et al., arXiv:nucl-ex/1101.1912

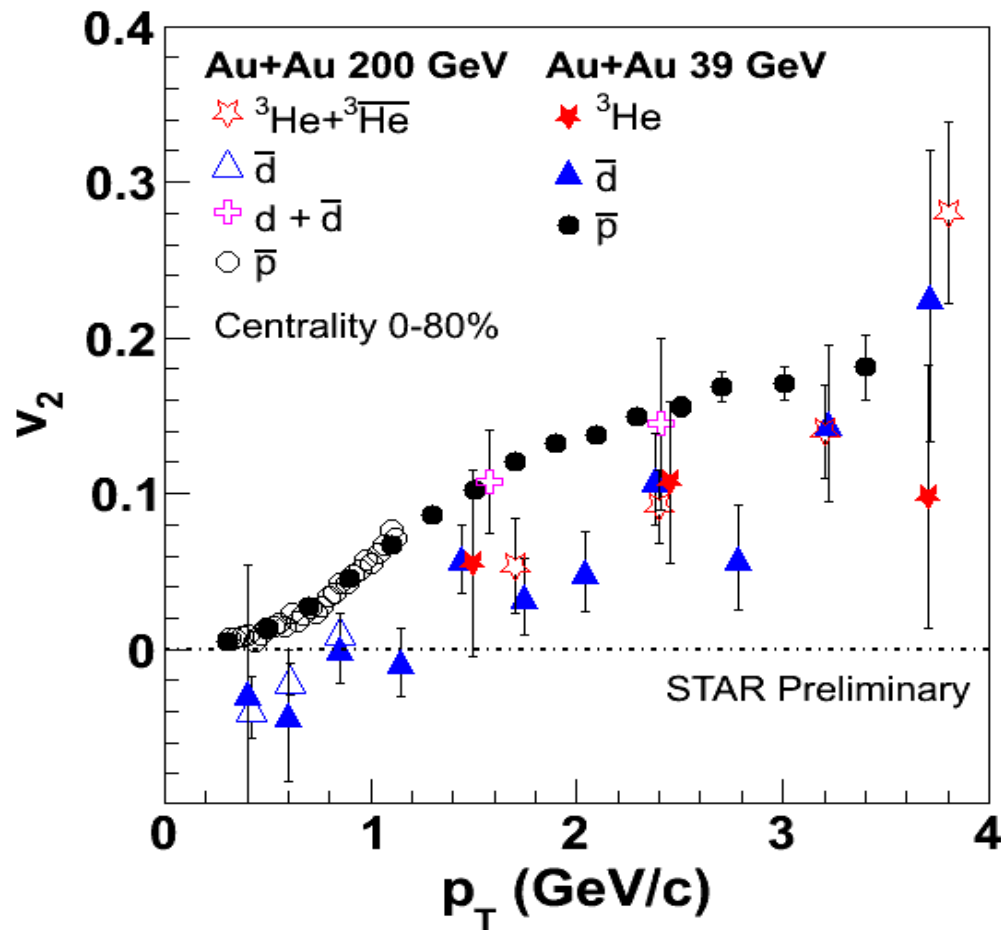


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Energy Dependence

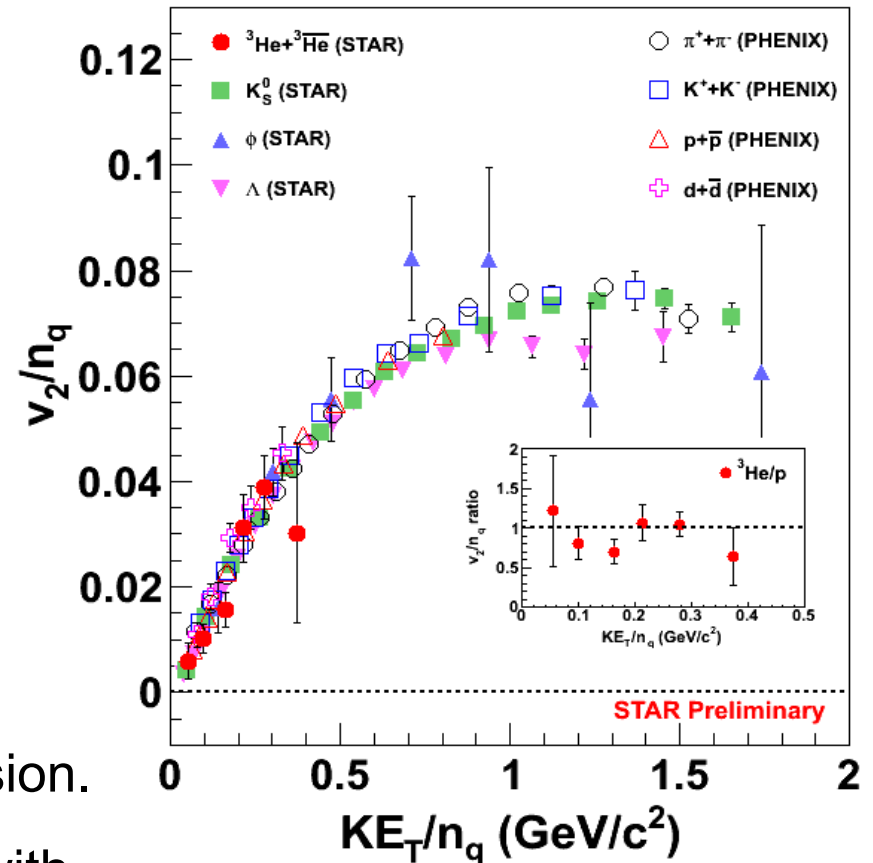


➤ Light nuclei v_2 is comparable in both beam energies.



Summary

- v_2 of light nuclei are measured in Au+Au 200 GeV and 39 GeV using event plane method.
- p_T dependence of v_2 is well described by Dynamical Coalescence Model.
- v_2 of light nuclei seem to follow the atomic mass number scaling.
- Number of constituent quark scaling holds good for v_2 of light nuclei.
- At more central collision the large value of v_2/ε indicates stronger collective expansion.
- Both v_2 and $\langle p_T \rangle$ trends are consistent with expectations from Blast-wave model fit.
- Light nuclei v_2 is comparable in both beam energies.





Thank You!

Back-up



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Tsallis Blast-Wave Model

$$\frac{dN}{m_T dm_T d\phi} \propto m_T \int_0^{2\pi} d\phi_s \int_{-Y}^{+Y} dy \cosh(y) \times \int_0^R r dr \left(1 + \frac{q-1}{T} E_T\right)^{-1/(q-1)}$$

where

Z. Tang et al., PRC 79, 051901 (2009)

Z. Tang et al., arXiv:nucl-ex/1101.1912

$$E_T = m_T \cosh(y) \cosh(\rho) - p_T \sinh(\rho) \cos(\phi_b - \phi)$$

$$\rho = \sqrt{(r \cos(\phi_s)/R_X)^2 + (r \sin(\phi_s)/R_Y)^2} (\rho_0 + \rho_2 \cos(2\phi_b))$$

$$\tan(\phi_b) = (R_X/R_Y)^2 \tan(\phi_s)$$

ρ = Flow profile in transverse rapidity

ϕ_s = Azimuthal angle of the co-ordinate space

ϕ_b = Angle of the flow direction

q = Degree of non-equilibrium

T (MeV)	q	ρ_0	ρ_2	$\frac{R_X}{R_Y}$	χ^2/ndf
88.3	1.049	0.86	0.054	0.89	512.6
± 1.2	± 0.003	± 0.013	± 0.001	± 0.002	/137

