

The STAR W Program at RHIC

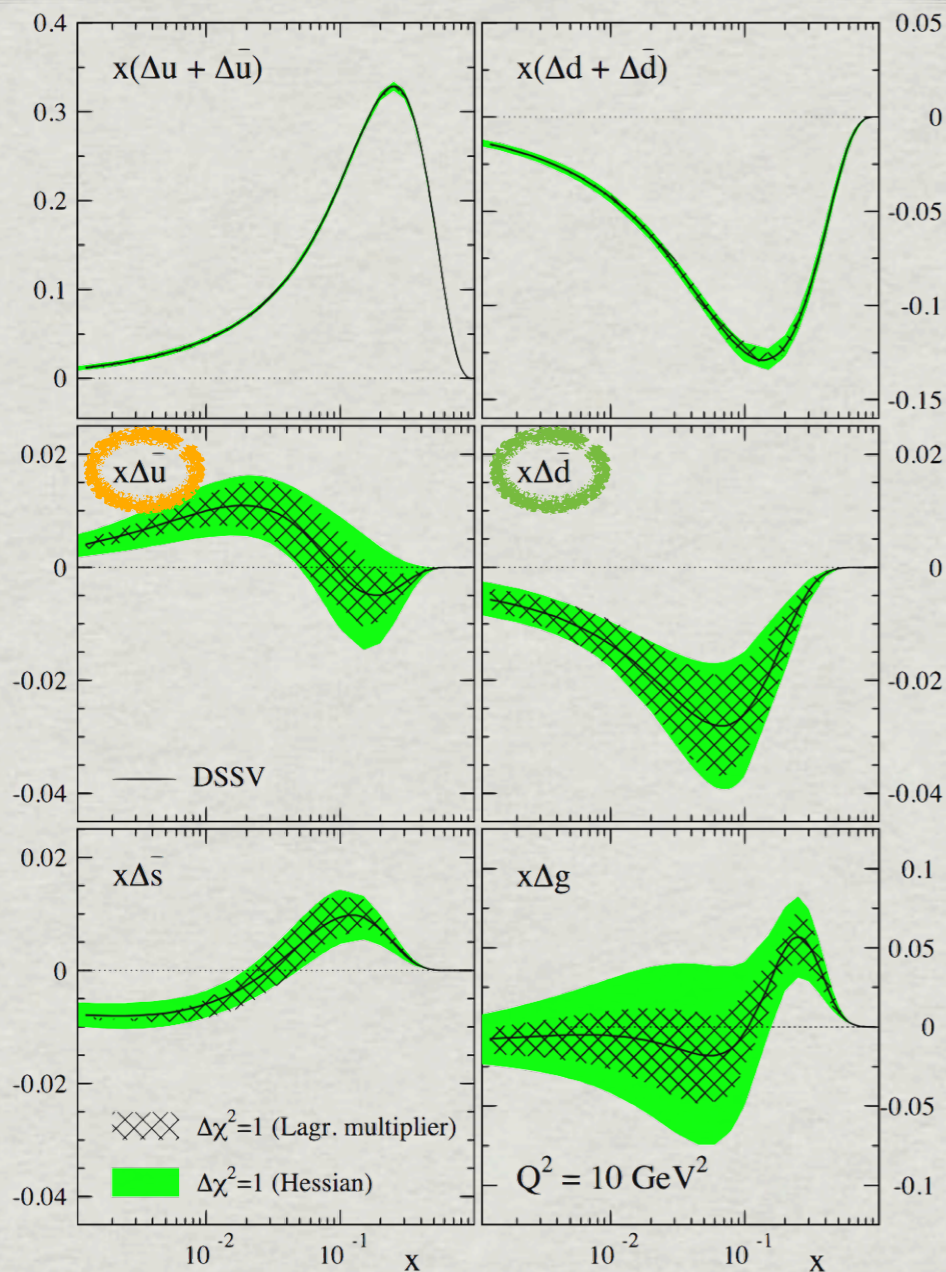
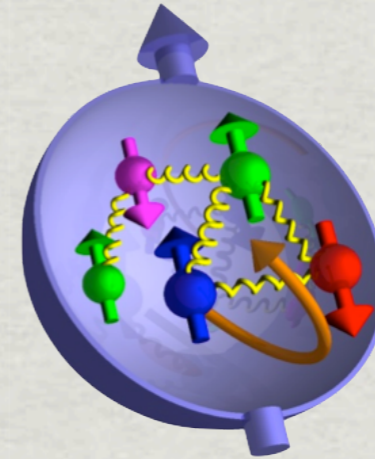
Justin Stevens for the STAR Collaboration

QCD-N' 2012



Proton Spin Puzzle

DSSV Global Analysis



PRD **80**, 034030 (2009)

$$\langle S_p \rangle = \frac{1}{2} = \frac{1}{2} \Delta\Sigma + \Delta G + L$$

$$\Delta\Sigma = \int (\Delta u + \Delta d + \Delta s + \Delta\bar{u} + \Delta\bar{d} - \Delta\bar{s}) dx$$

Integral of quark polarization is well measured in DIS to be ~30%, some info on decomposition from SIDIS but sea not well constrained

$$\Delta G = \int \Delta g(x) dx$$

First experimental evidence of non-zero Δg from 2009 RHIC data (previous two talks)

Polarized PDF

$$\Delta f(x) =$$

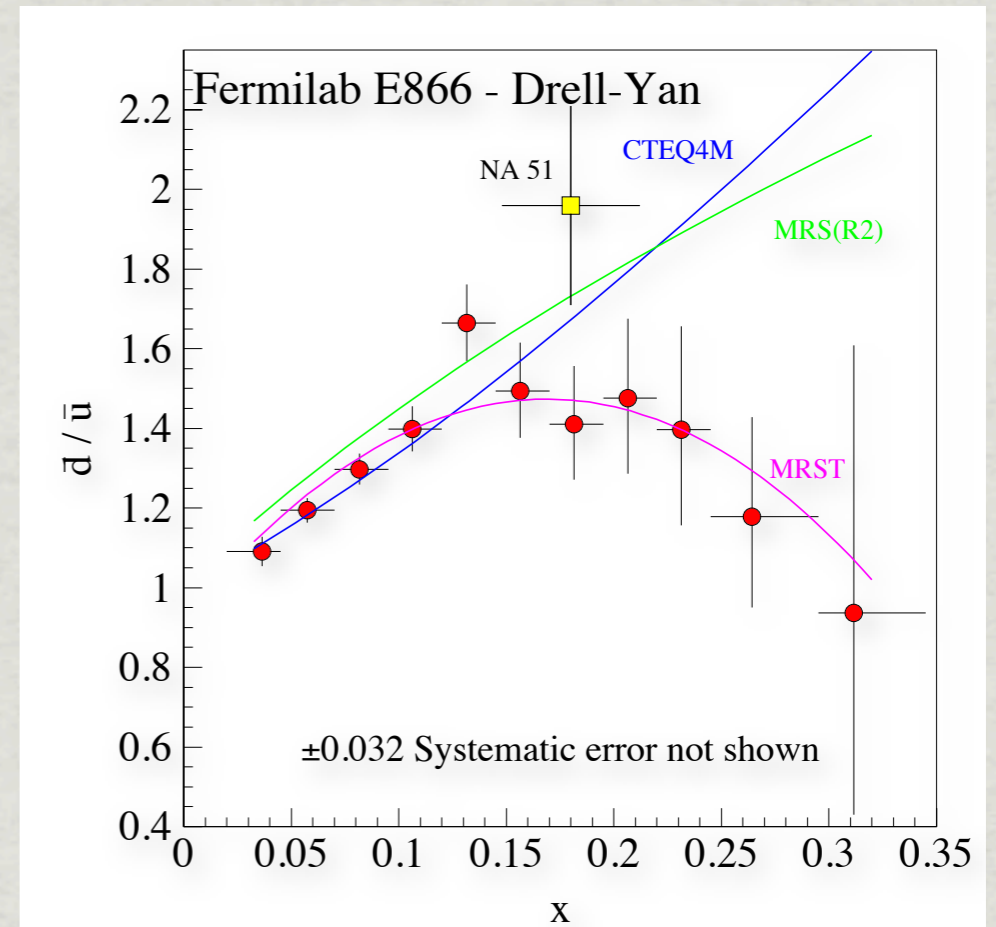


$$f^+(x) - f^-(x)$$

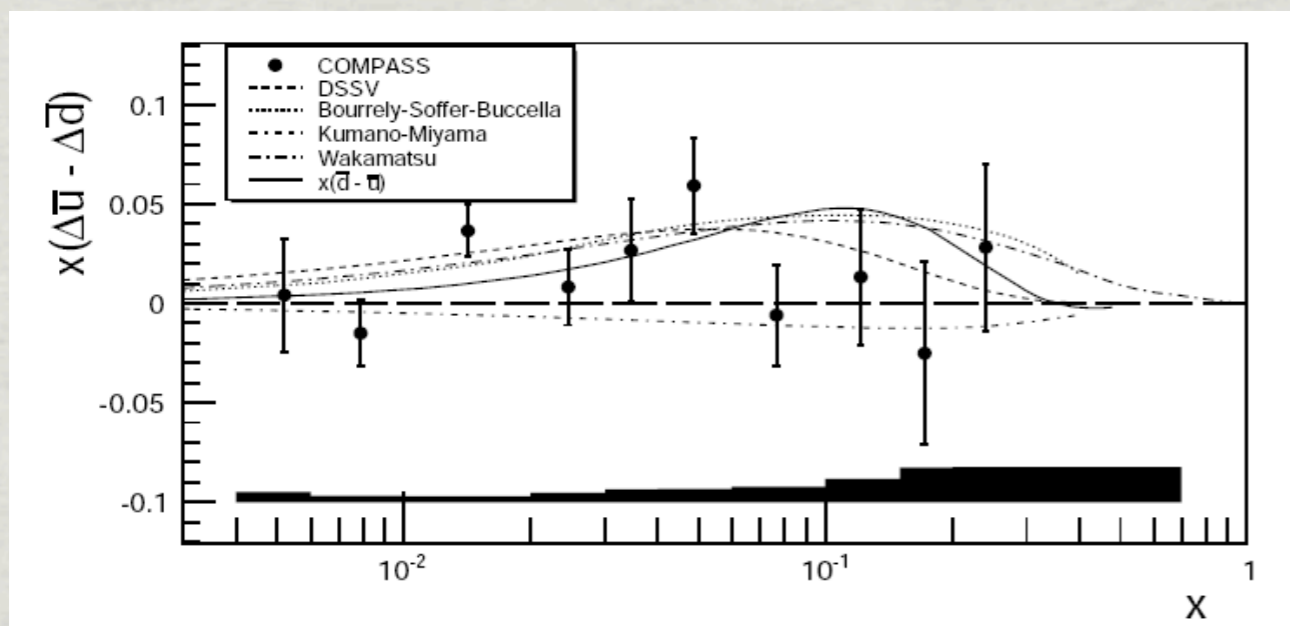
Flavor Asymmetry of the Sea

Unpolarized Flavor Asymmetry:

- * Quantitative calculation of Pauli blocking does not explain \bar{d}/\bar{u} ratio
- * Non-perturbative processes may be needed in generating the sea
- * E866 results are qualitatively consistent with pion cloud models, chiral quark soliton models, instanton models, etc.



PRD **64**, 052002 (2001)

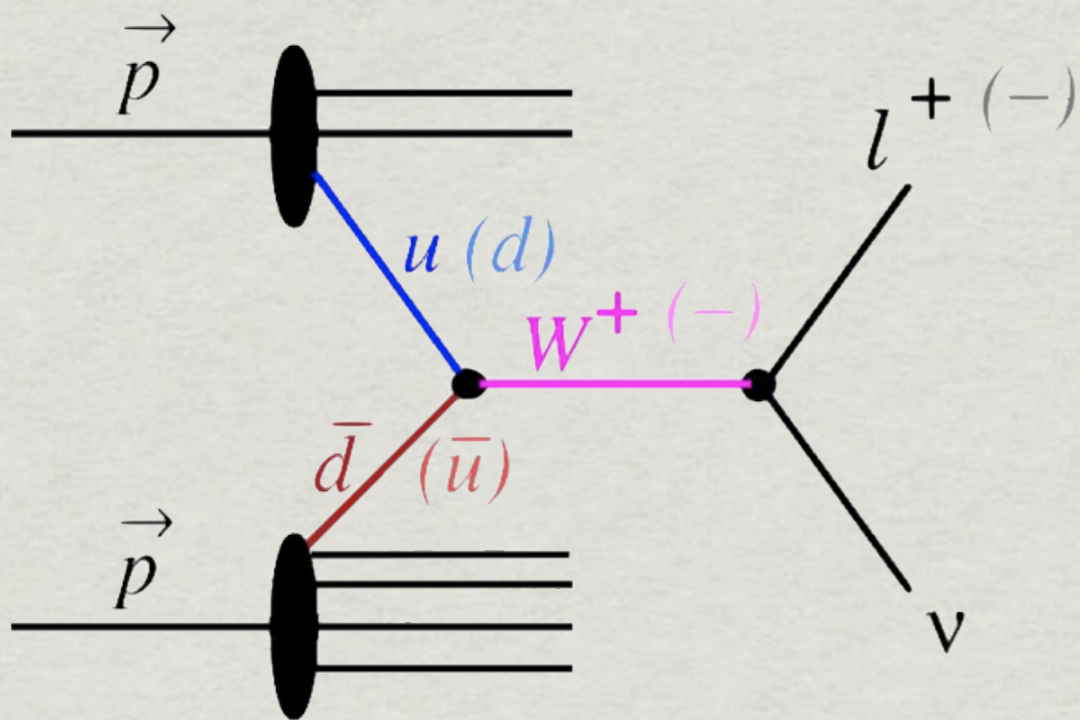


PLB **693**, 227 (2010)

Polarized Flavor Asymmetry:

- * Valence u and d distributions are well determined from DIS
- * Polarized flavor asymmetry $x(\Delta\bar{u} - \Delta\bar{d})$ could help differentiate models
- * SIDIS results depend on FFs

Why Ws?



$$u + \bar{d} \rightarrow W^+ \rightarrow e^+ + \nu$$

$$d + \bar{u} \rightarrow W^- \rightarrow e^- + \bar{\nu}$$

- * Ws couple directly to the quarks and antiquarks of interest
- * Detect Ws through e+/e- decay channels
- * V-A coupling of the weak interaction leads to perfect spin separation

Measure parity-violating single-spin asymmetry:

(Helicity flip in one beam while averaging over the other)

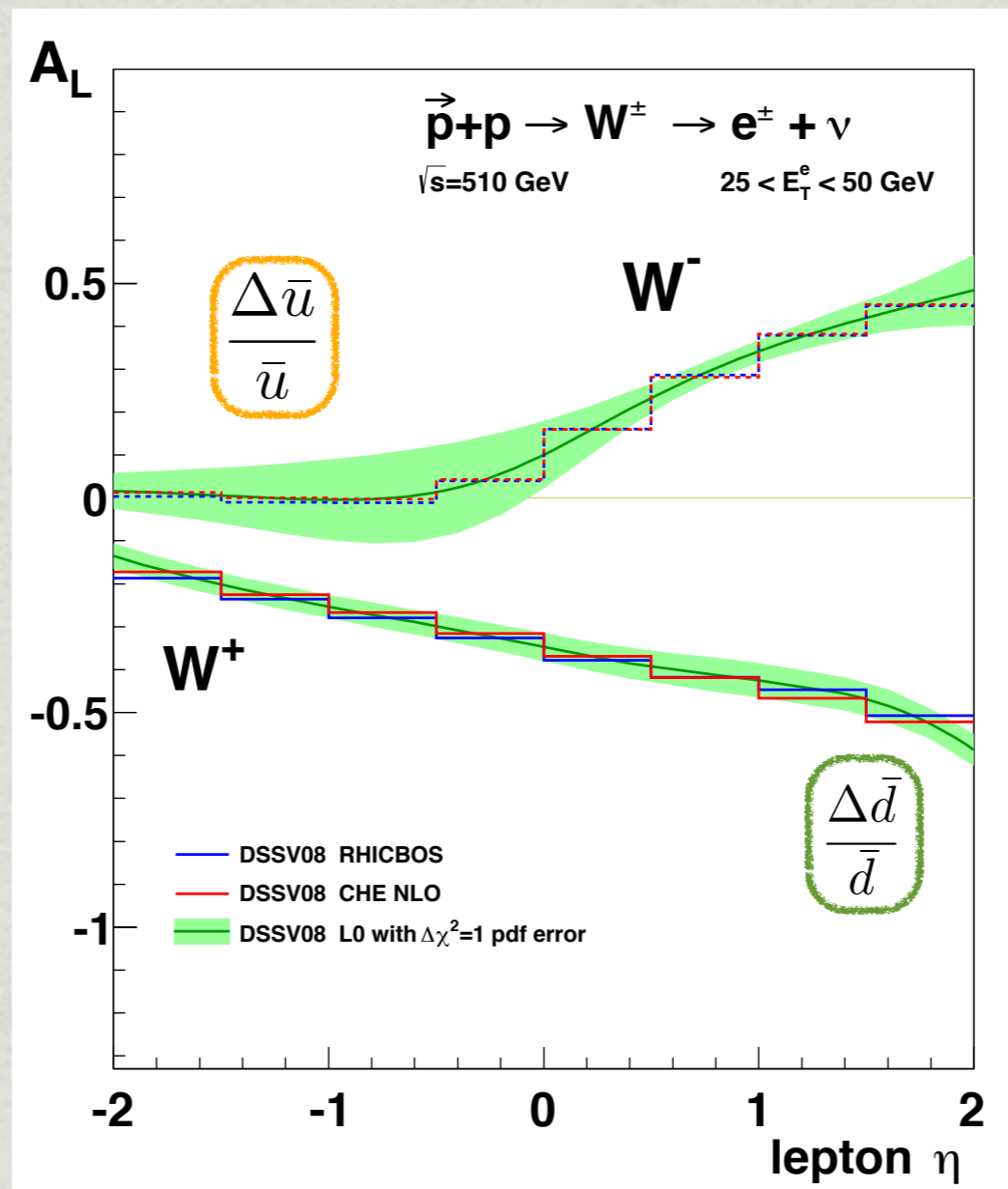
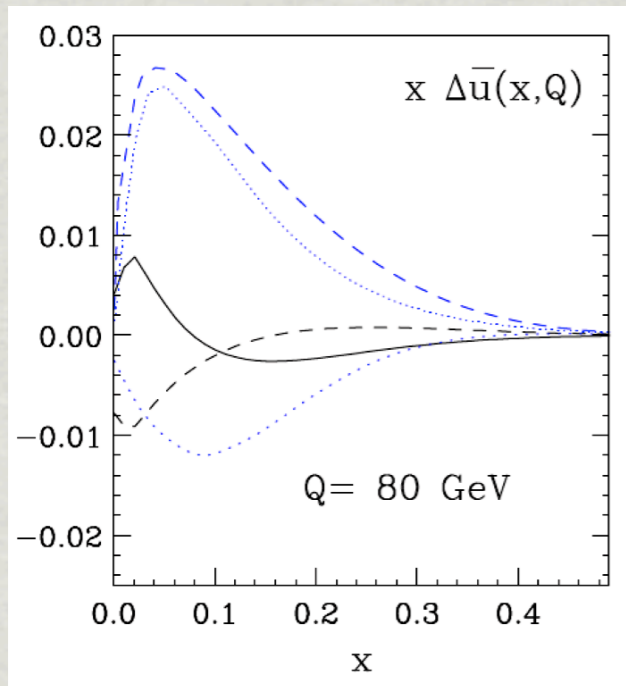
$$A_L = \frac{\sigma_+ - \sigma_-}{\sigma_+ + \sigma_-}$$

$$A_L^{W^-} \propto \frac{-\Delta d(x_1)\bar{u}(x_2) + \Delta\bar{u}(x_1)d(x_2)}{d(x_1)\bar{u}(x_2) + \bar{u}(x_1)d(x_2)}$$

$$A_L^{W^+} \propto \frac{-\Delta u(x_1)\bar{d}(x_2) + \Delta\bar{d}(x_1)u(x_2)}{u(x_1)\bar{d}(x_2) + \bar{d}(x_1)u(x_2)}$$

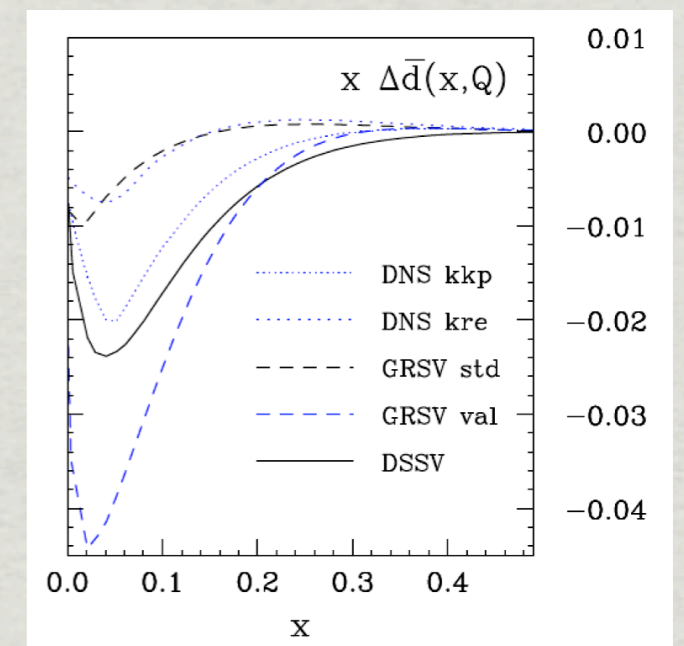
Expectations for $W A_L$

$$A_L^{W^-} \propto \frac{-\Delta d(x_1)\bar{u}(x_2) + \Delta\bar{u}(x_1)d(x_2)}{d(x_1)\bar{u}(x_2) + \bar{u}(x_1)d(x_2)} \quad A_L^{W^+} \propto \frac{-\Delta u(x_1)\bar{d}(x_2) + \Delta\bar{d}(x_1)u(x_2)}{u(x_1)\bar{d}(x_2) + \bar{d}(x_1)u(x_2)}$$



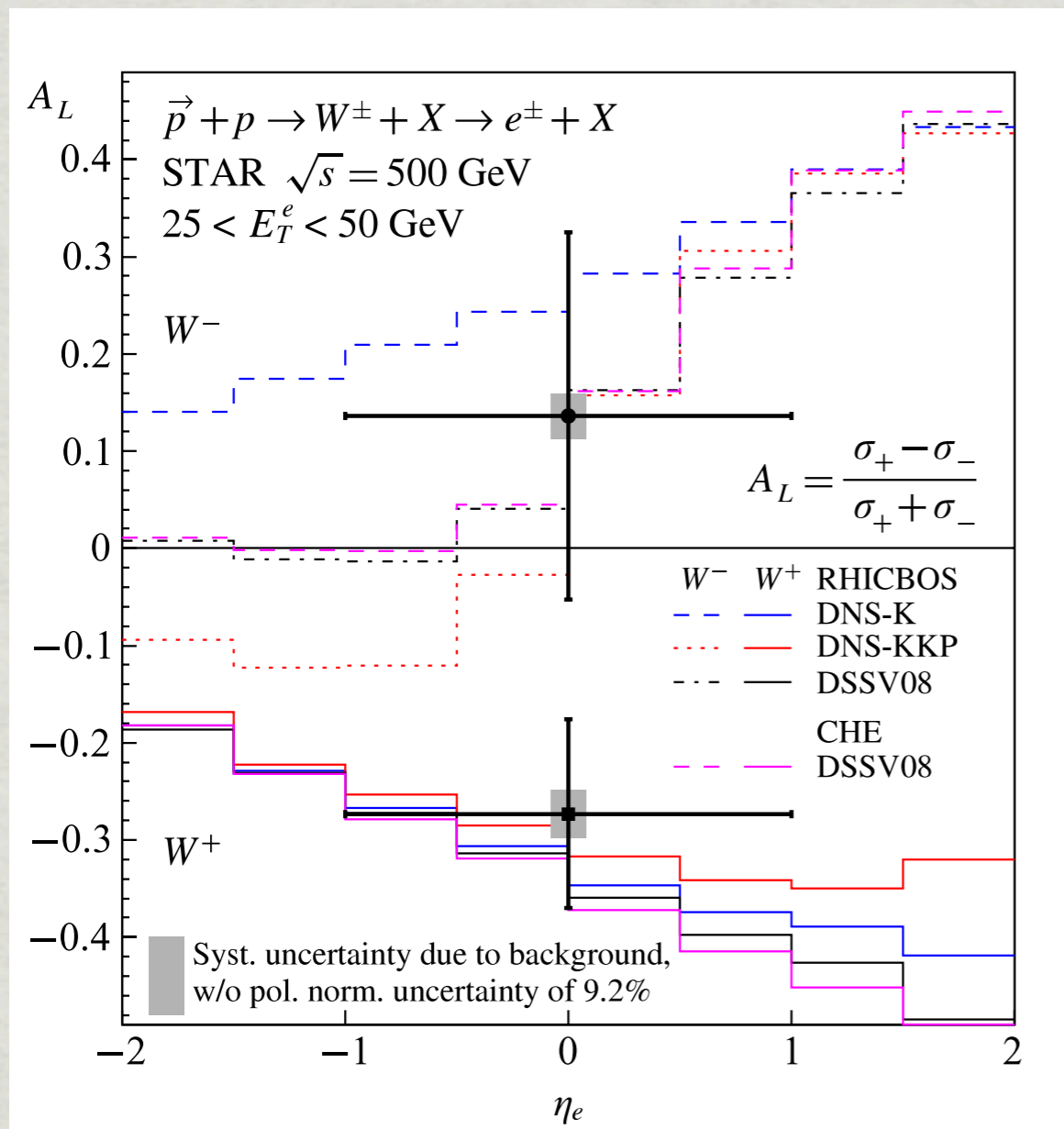
- * Large parity-violating asymmetries expected
- * Simplified interpretation at forward and backward rapidity

- * DSSV $\Delta\chi^2=1$ band underestimates the theoretical uncertainty (and Lagrange multiplier estimates for a $\Delta\chi^2/\chi^2 = 2\%$ error are in progress)

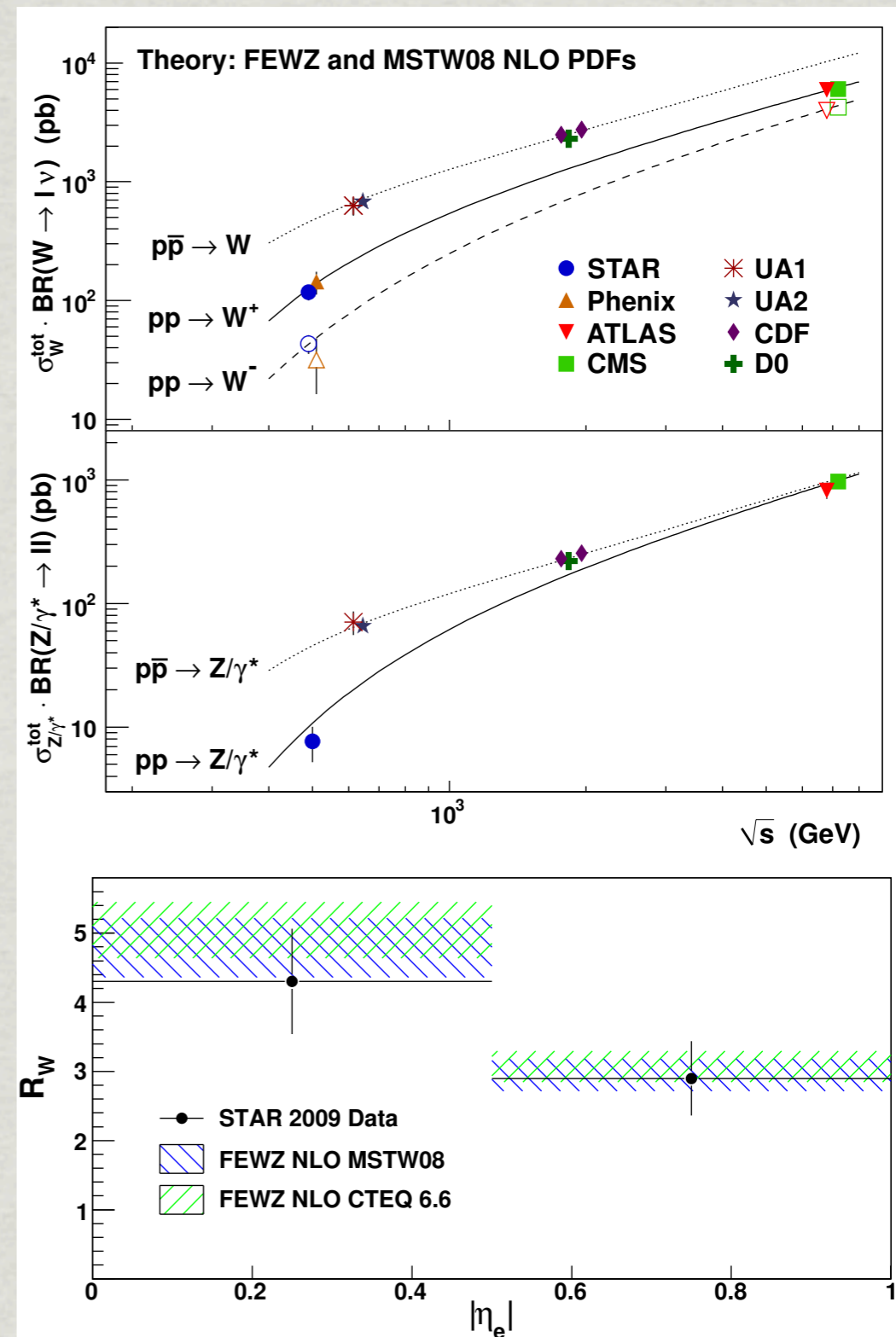


STAR 2009 W Measurements

Very successful first
500 GeV physics run!



PRL 106, 062002 (2011)

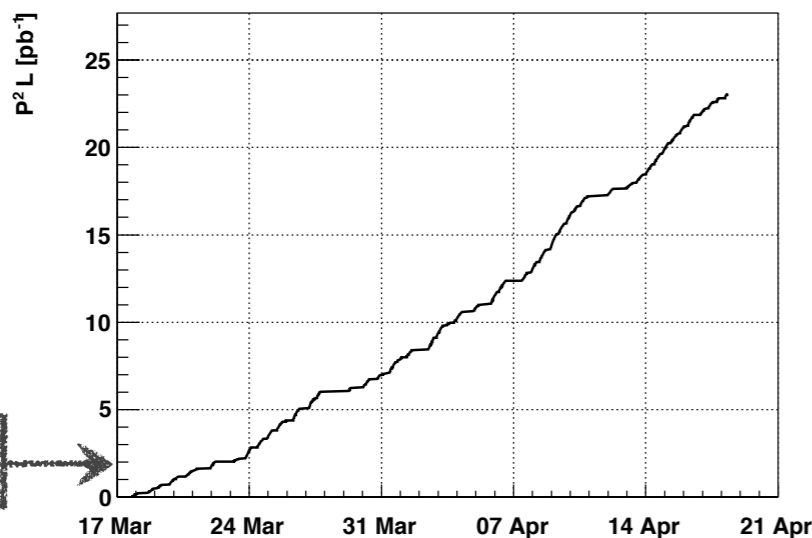


PRD 85, 92010 (2012)

NEW!

2012 STAR Dataset

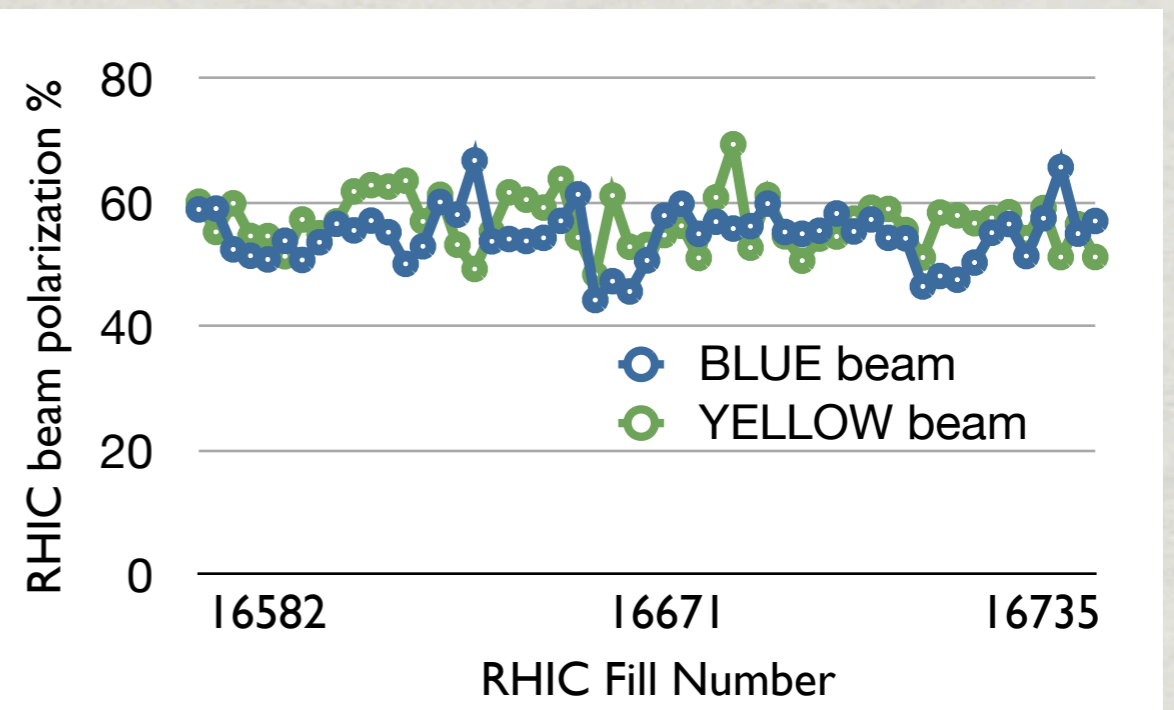
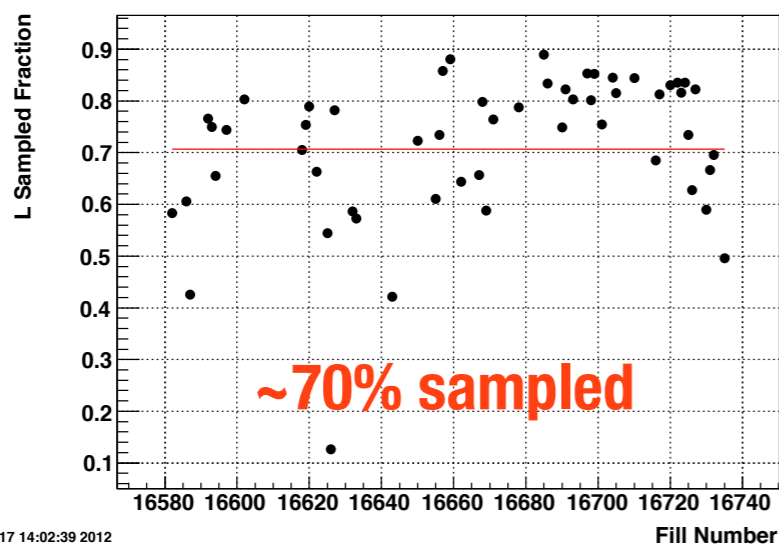
W Trigger FOM = P^2L



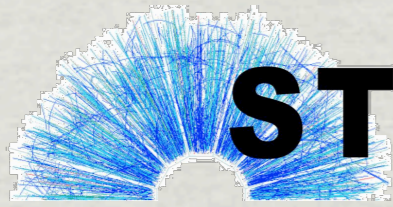
	L (pb^{-1})	P	P^2L (pb^{-1})
Run 9	12	0.40	1.9
Run 12	72	0.56	22.6

Note: For Run 12 expect ~10% more statistics with final calibrations

Lumi Sampled Fraction



**Excellent RHIC and STAR
performance. Thanks!**



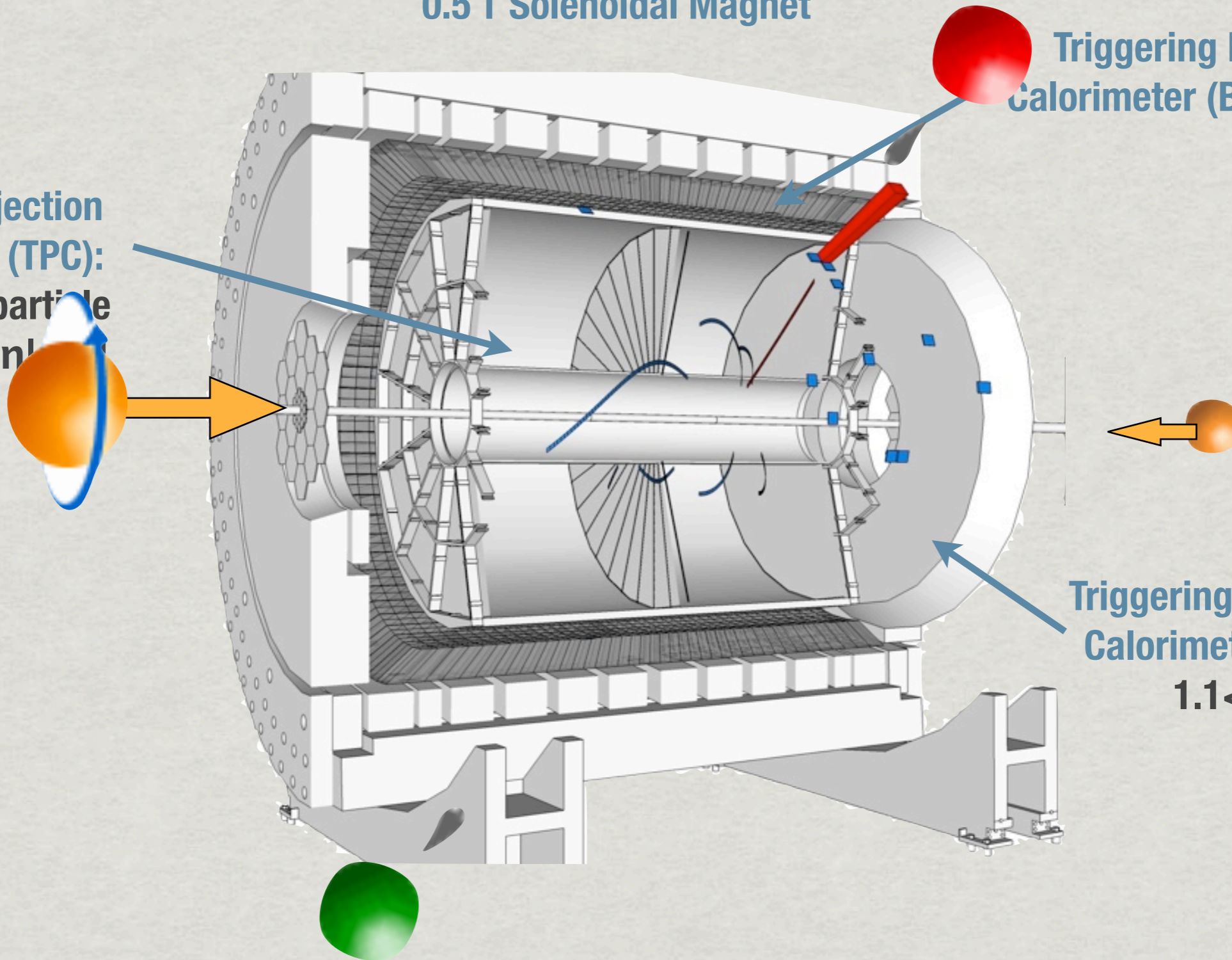
STAR Detector Overview

0.5 T Solenoidal Magnet

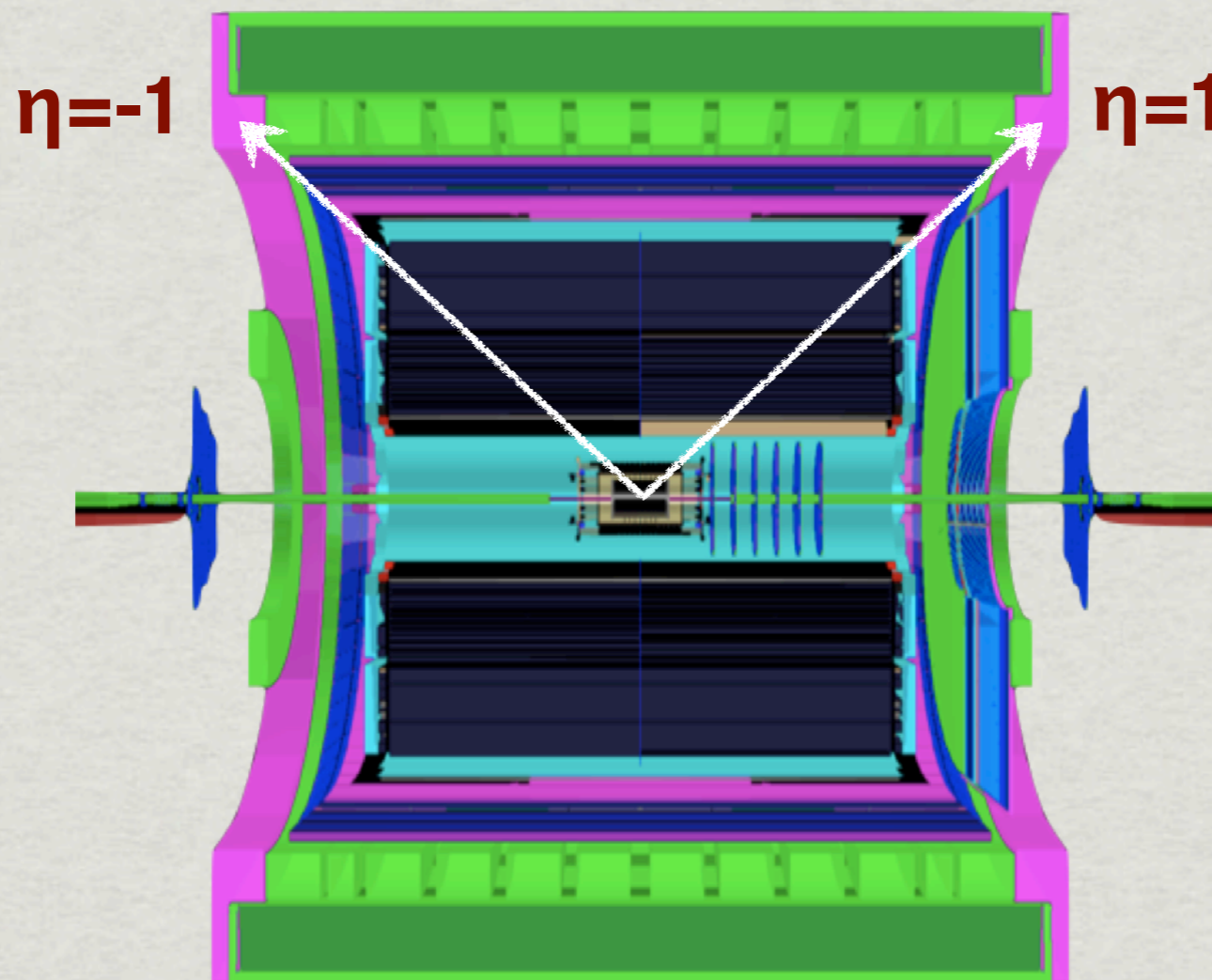
Triggering Barrel EM Calorimeter (BEMC): $|\eta| < 1$

Time Projection Chamber (TPC):
Charged particle tracking $|\eta| < 1$

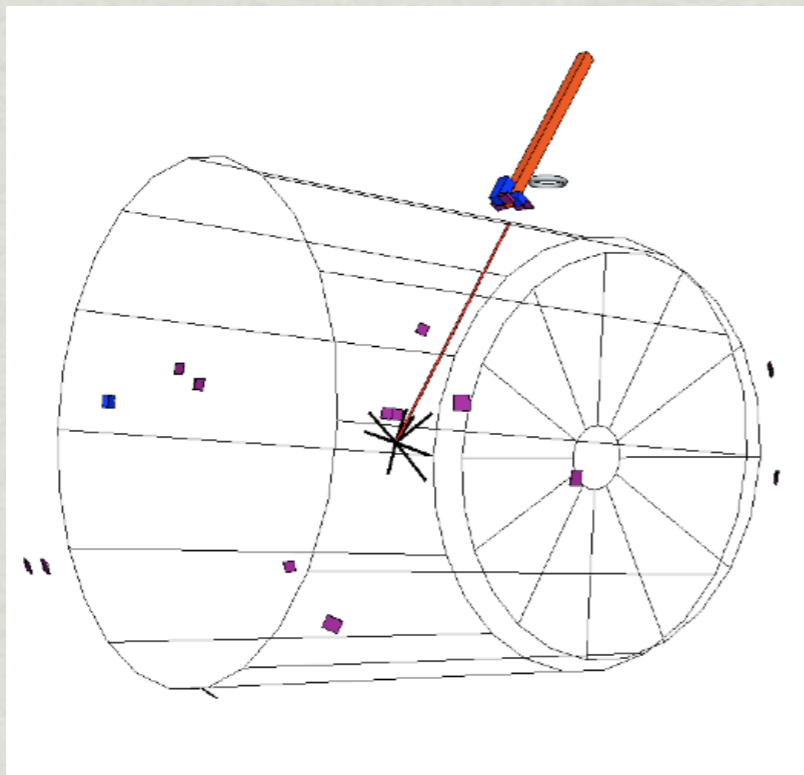
Triggering Endcap EM Calorimeter (EEMC):
 $1.1 < \eta < 2$



Mid-Rapidity W s at STAR



What do W decays look like?

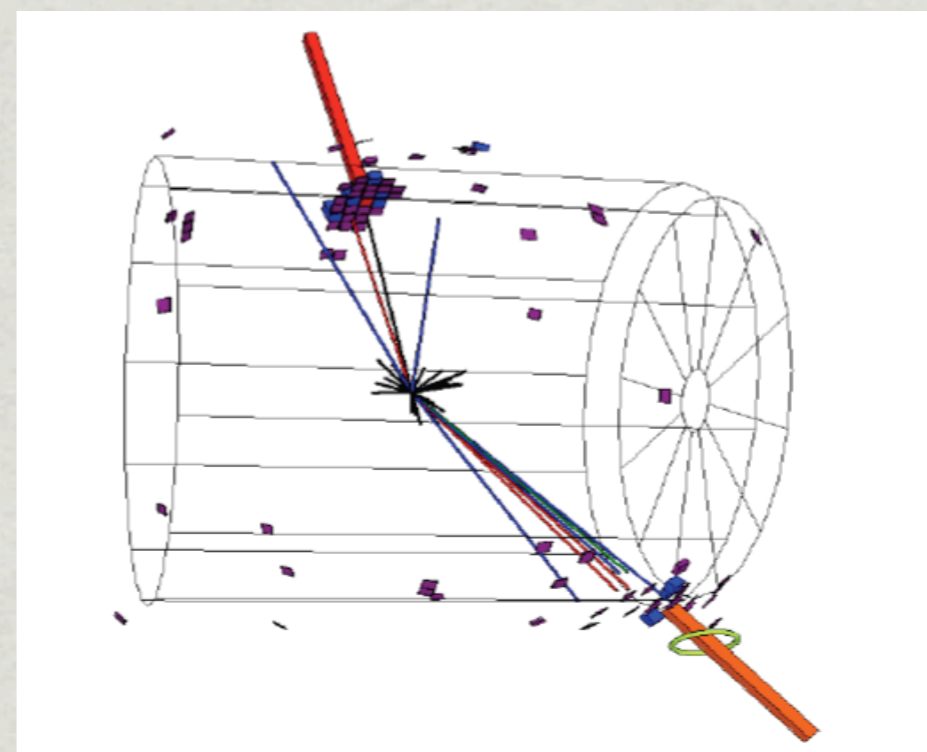


$W \rightarrow e + \nu$ Candidate Event

- Isolated track pointed to isolated EM cluster in calorimeter
- Large “missing energy” opposite the electron candidate

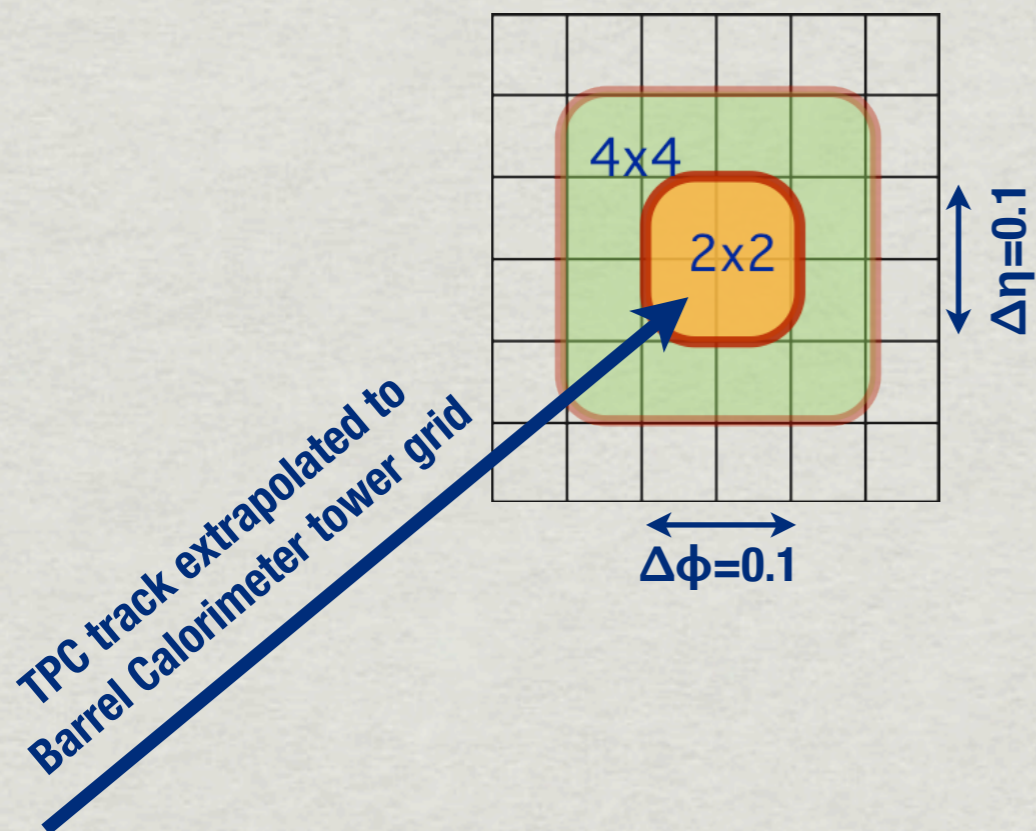
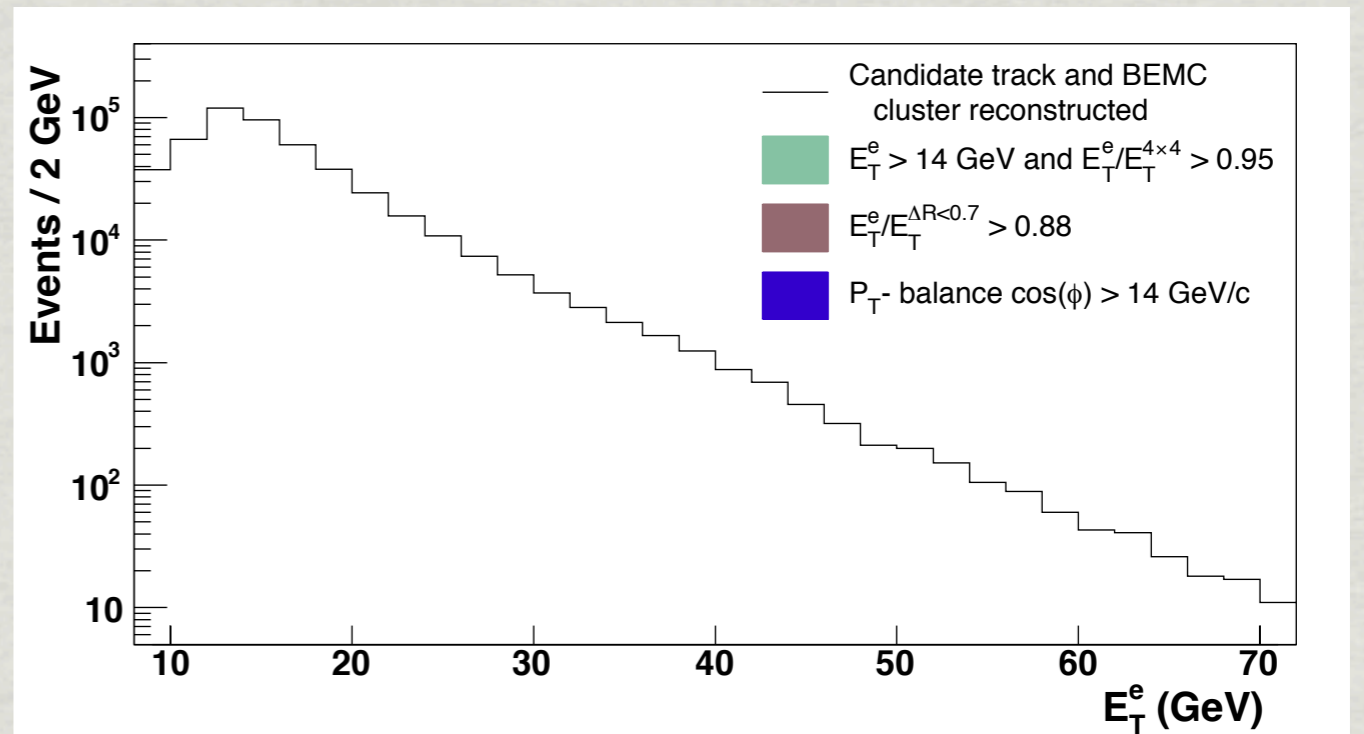
Di-jet Background Event

- Several tracks pointing to EM energy deposit in several towers
- Vector p_T sum is balanced by opposite jet, “missing energy” is small



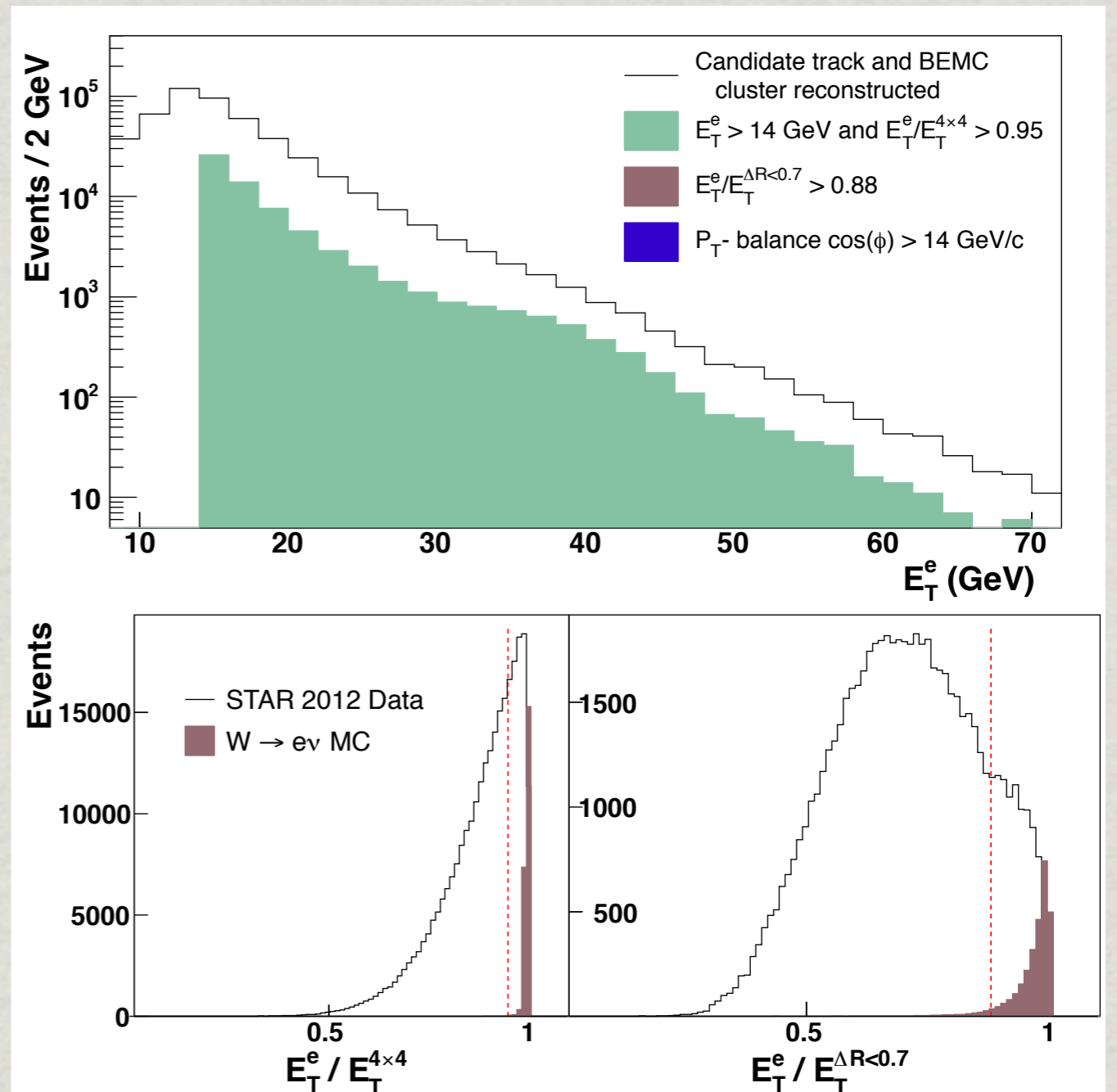
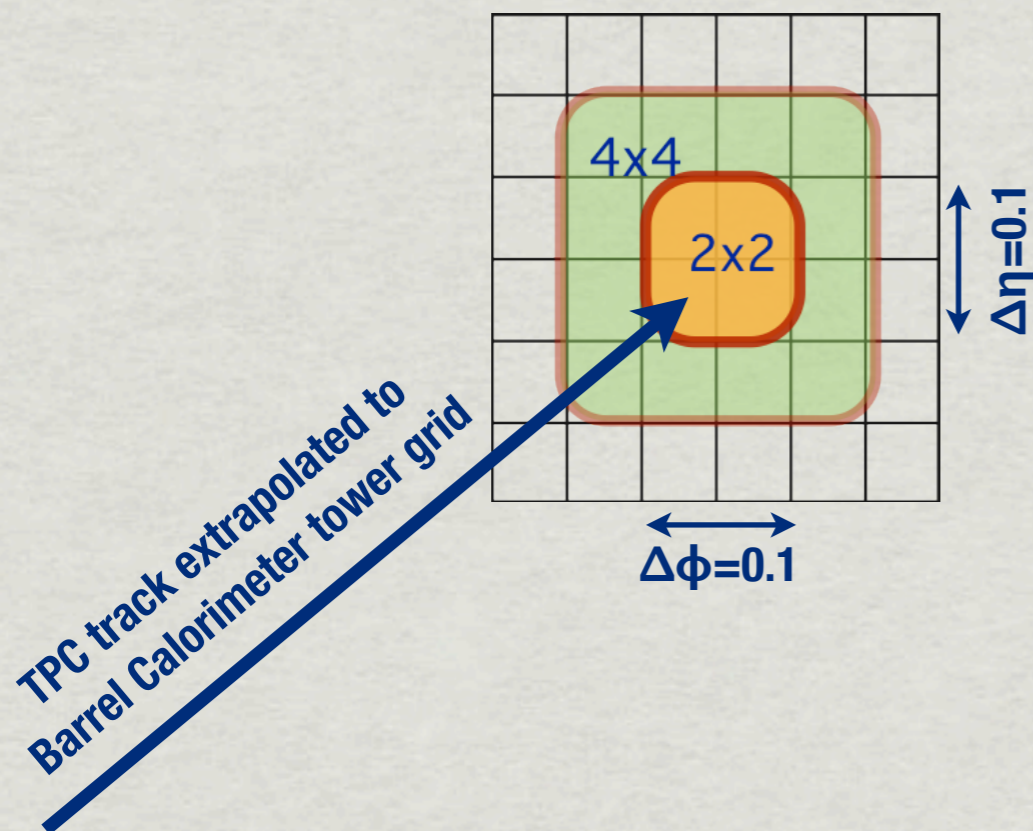
How do we find Ws?

- Match $p_T > 10$ GeV track to BEMC cluster



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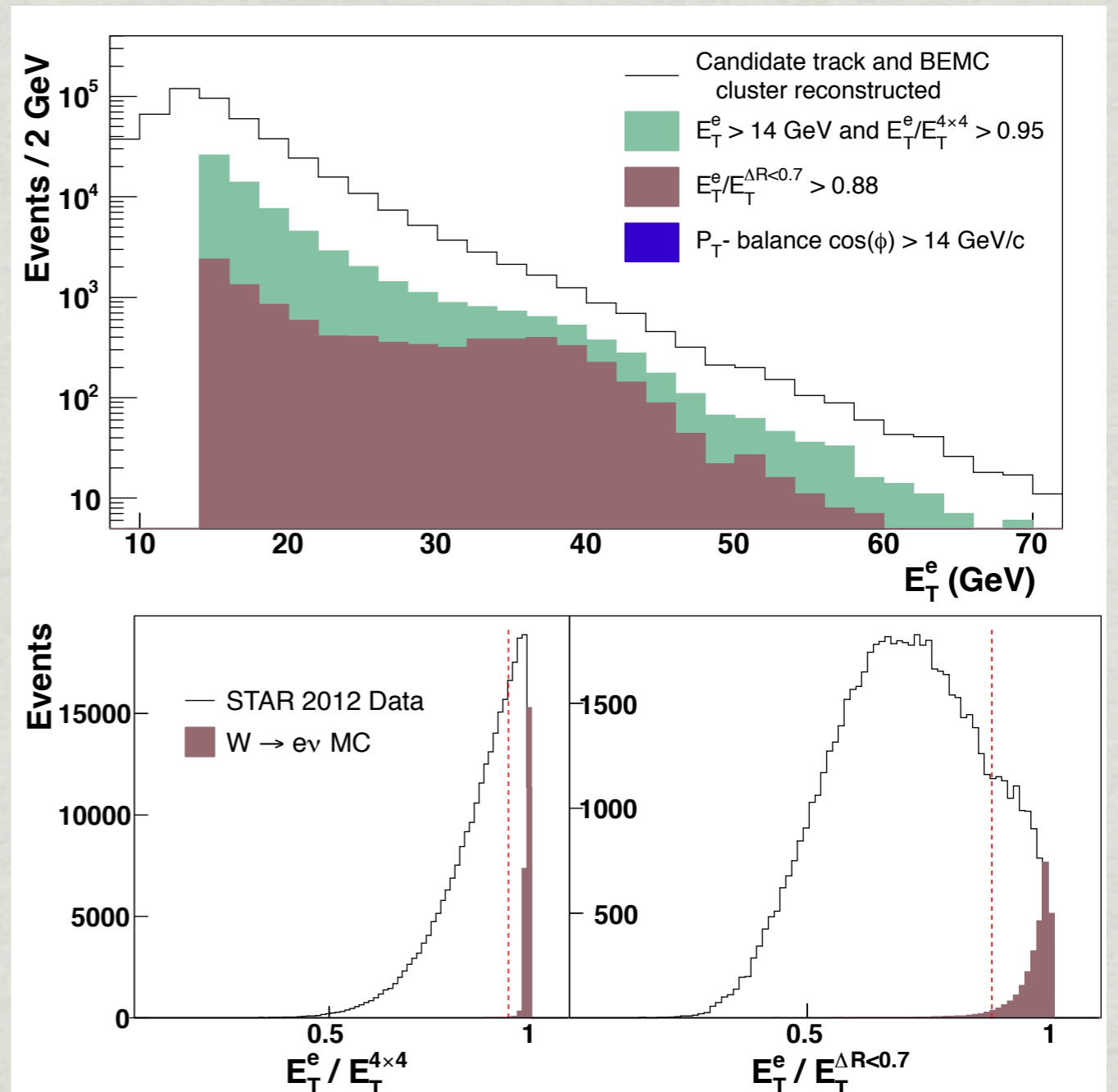
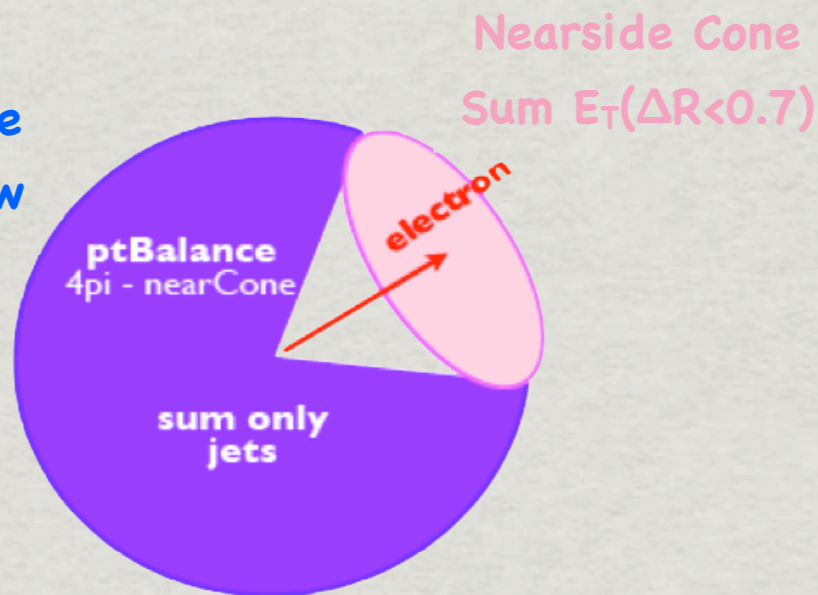
- Match $p_T > 10$ GeV track to BEMC cluster
- Isolation Ratios



How do we find Ws?

- Match $p_T > 10$ GeV track to BEMC cluster
- Isolation Ratios

Transverse Plane View



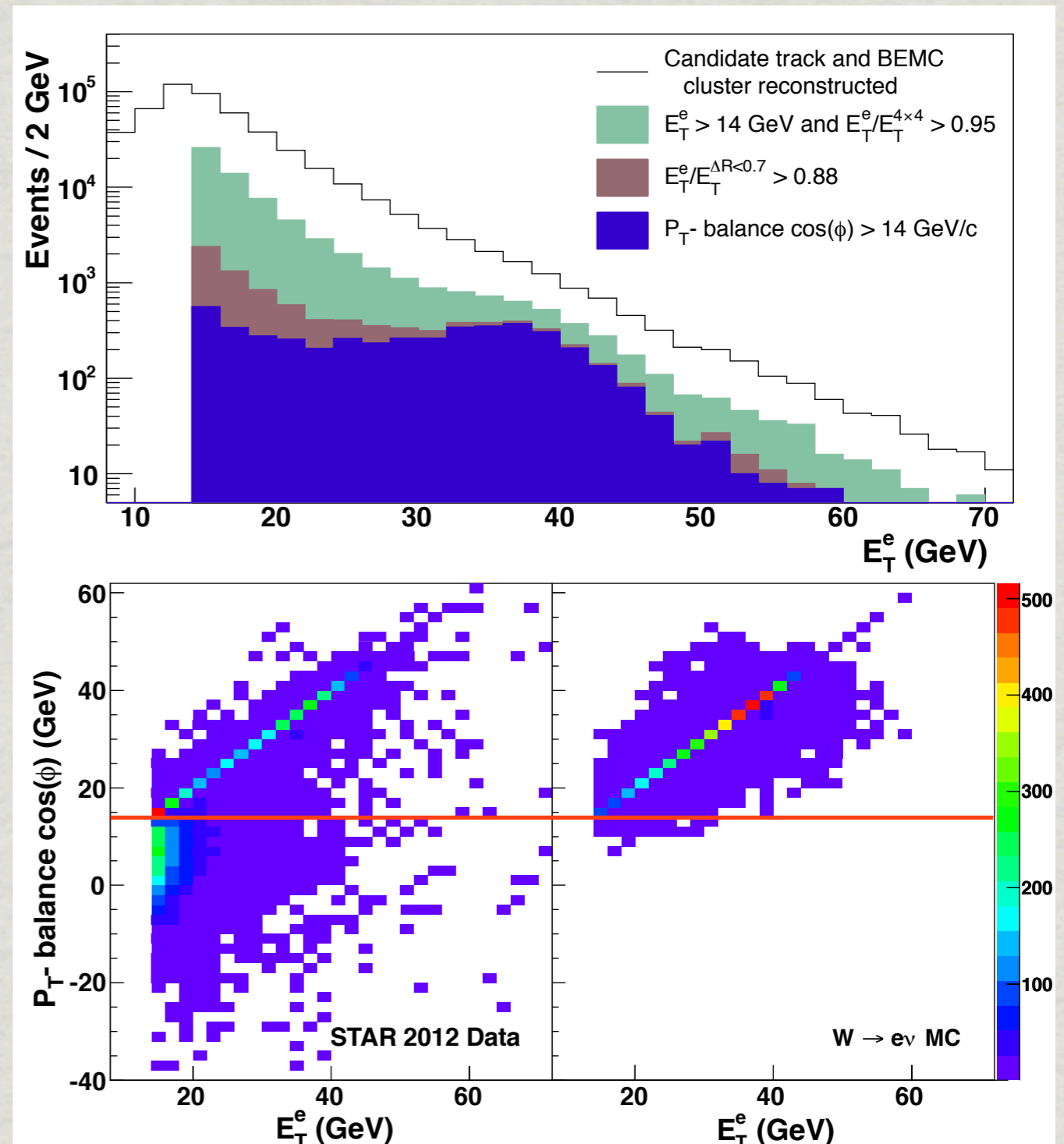
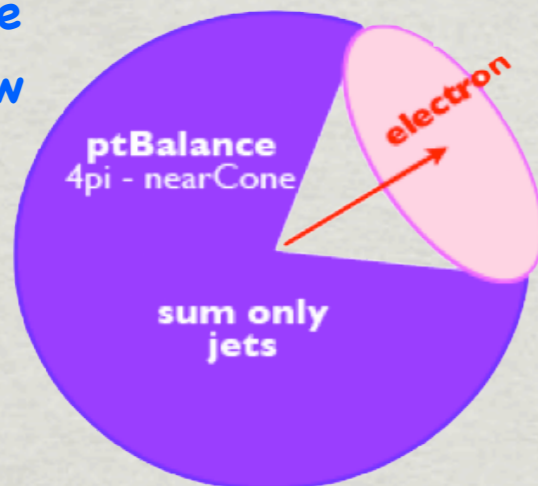
How do we find Ws?

- Match $p_T > 10$ GeV track to BEMC cluster
- Isolation Ratios
- P_T -balance

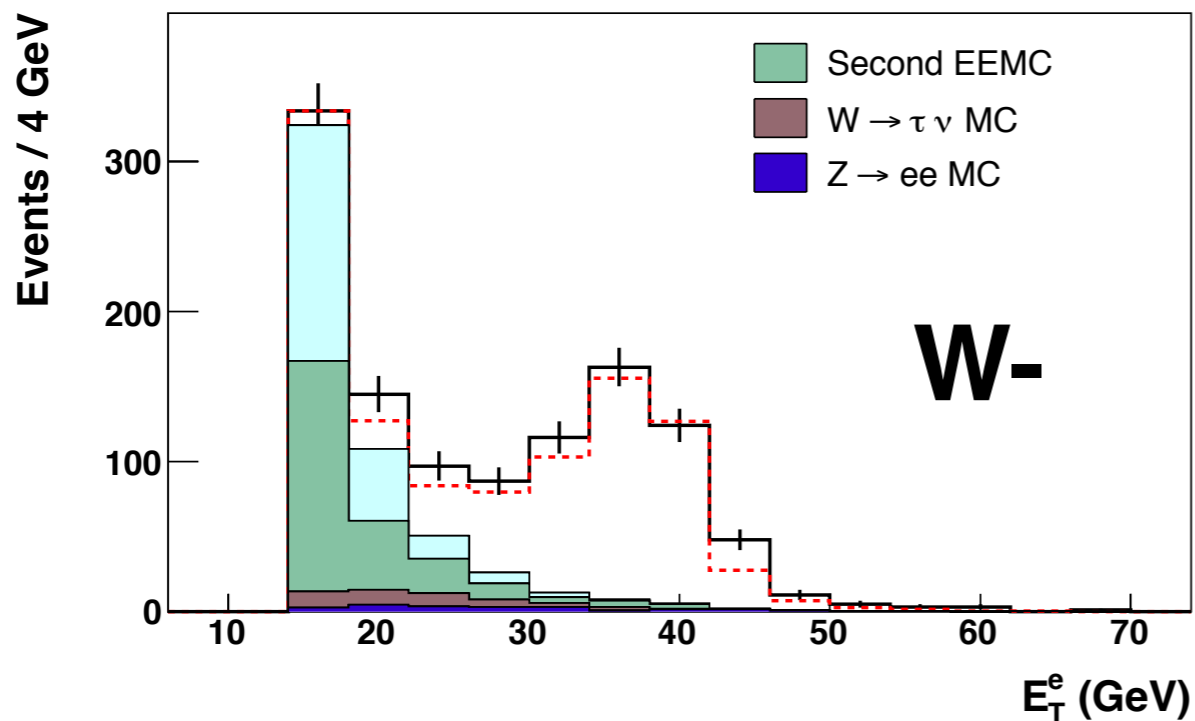
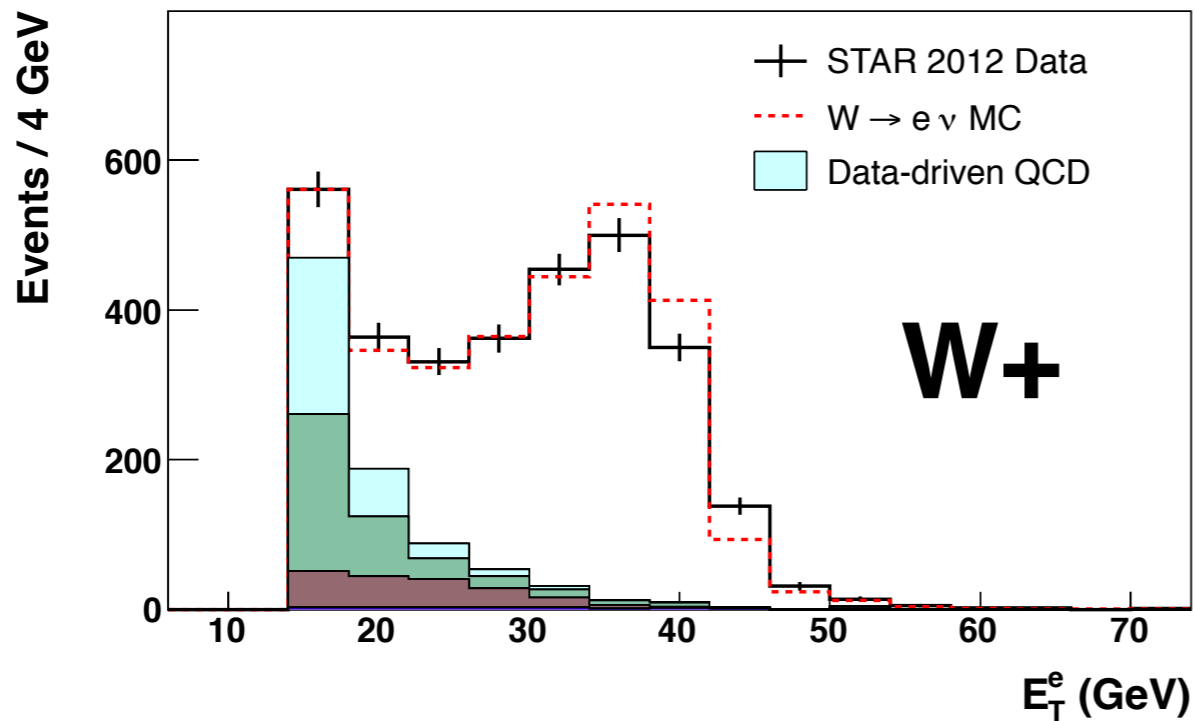
$$\vec{p}_T^{bal} = \vec{p}_T^e + \sum_{\Delta R > 0.7} \vec{p}_T^{jets}$$

$$P_T\text{-balance } \cos(\phi) = \frac{\vec{p}_T^e \cdot \vec{p}_T^{bal}}{|\vec{p}_T^e|}$$

Transverse
Plane View



Mid-rapidity Background Estimation



W Signal

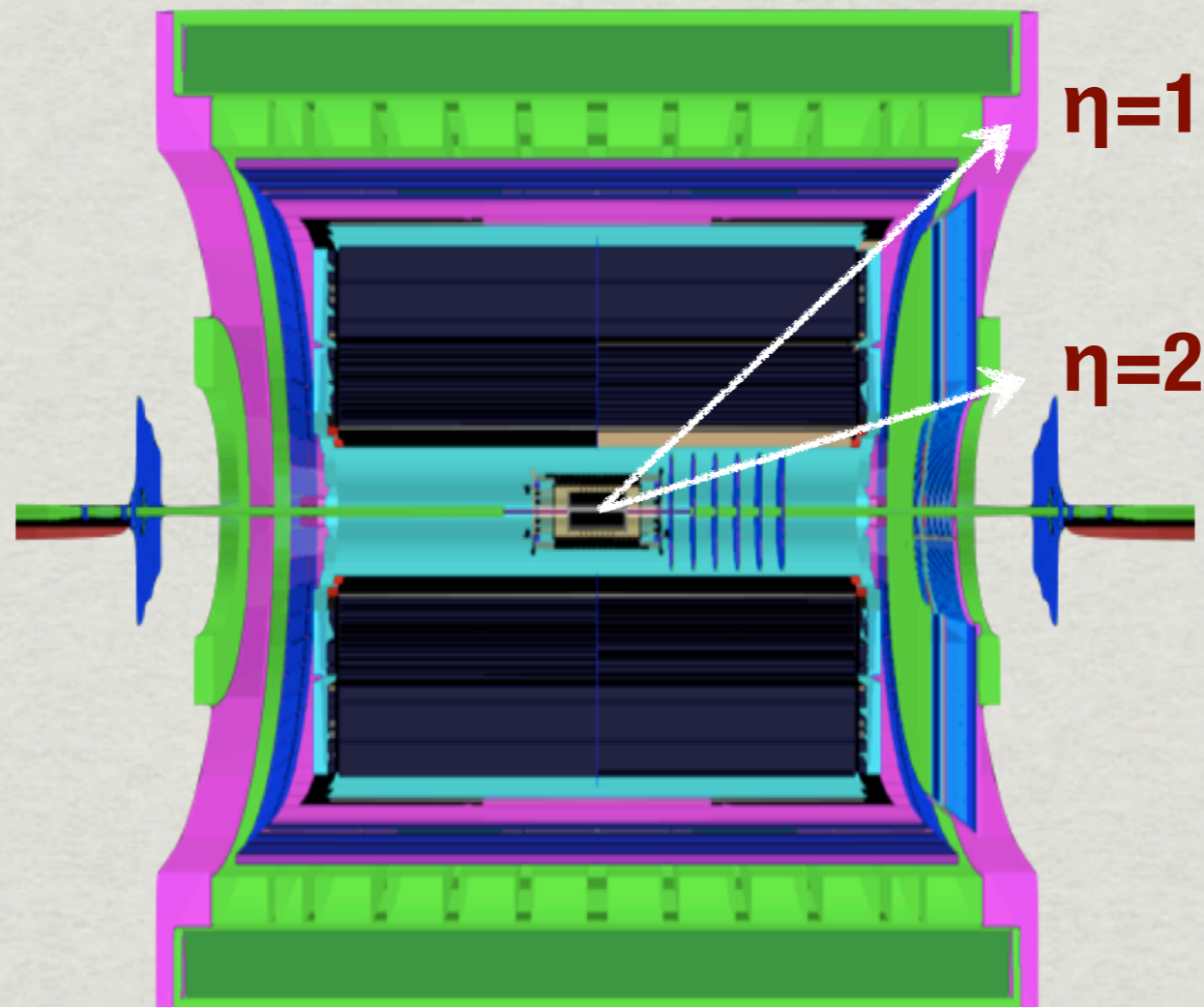
- * “Jacobian Peak”

Background Estimation

- * Electroweak
 - * $Z \rightarrow ee$ MC
 - * $W \rightarrow \tau \nu$ MC
- * Second EEMC
- * Data-driven QCD

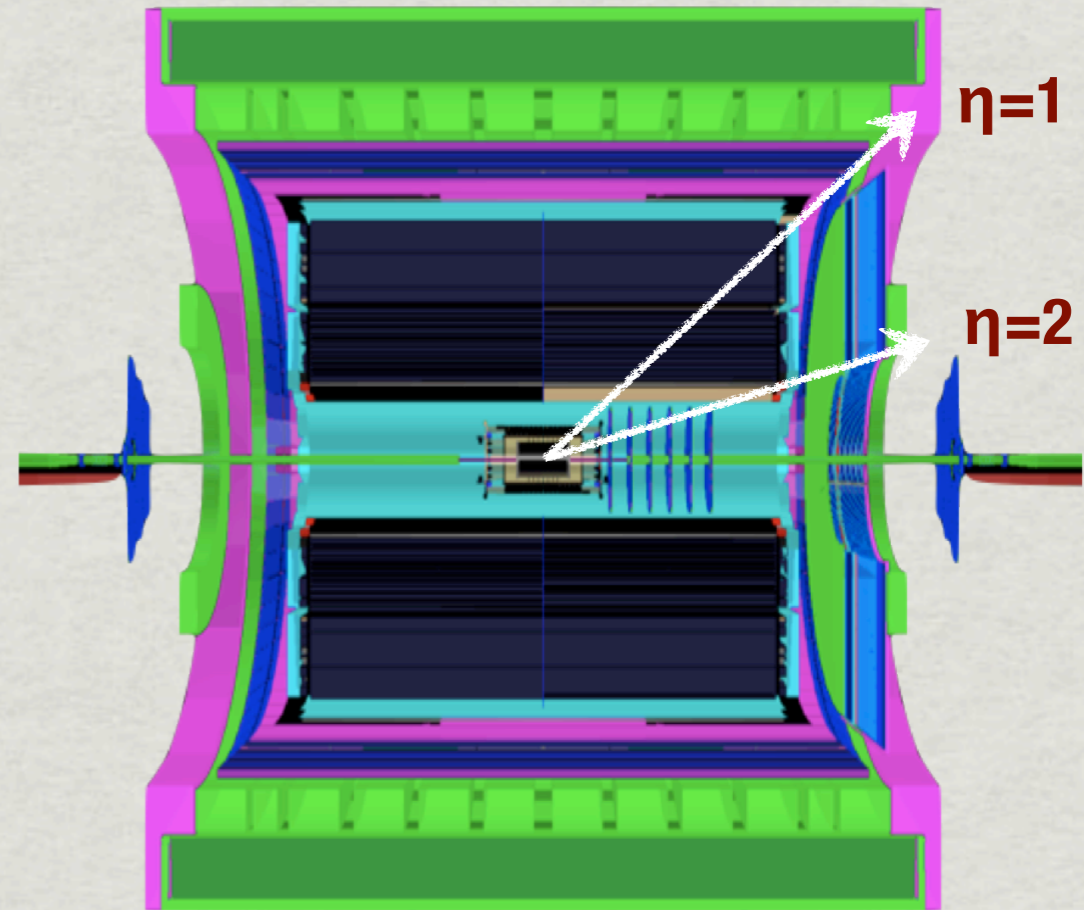
NEW!

Forward Rapidity Ws at STAR

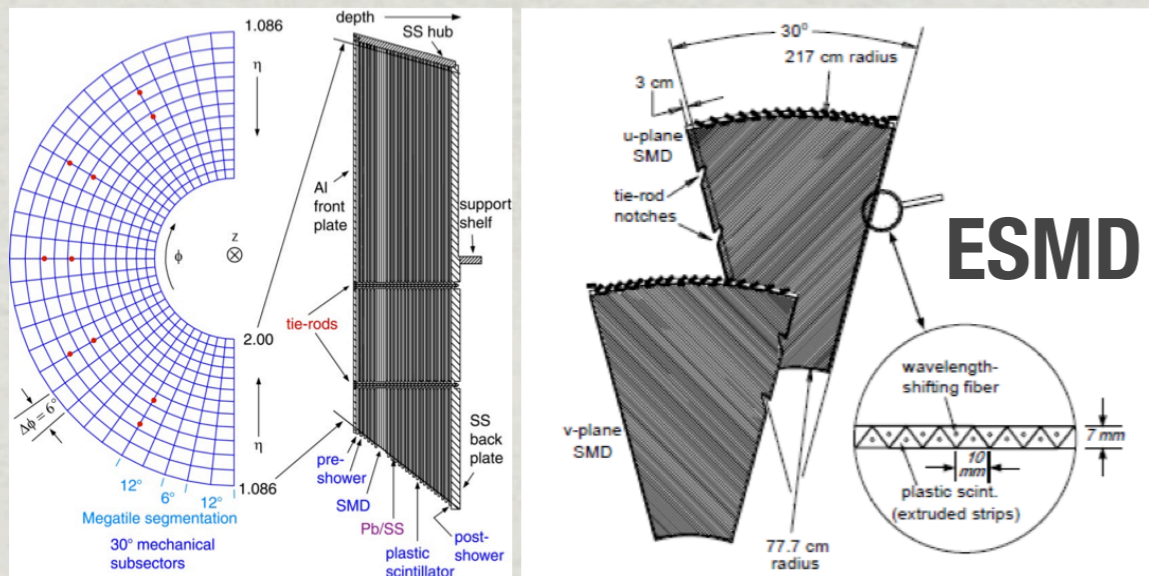


Forward Rapidity Analysis

- * Similar concept as mid-rapidity:
 - * Utilize TPC which extends to $\eta \sim 1.4$ to reconstruct high p_T TPC track
 - * Use isolation ratios and vector p_T imbalance to reduced QCD background
- * Improve background rejection by using the Endcap Shower Maximum Detector



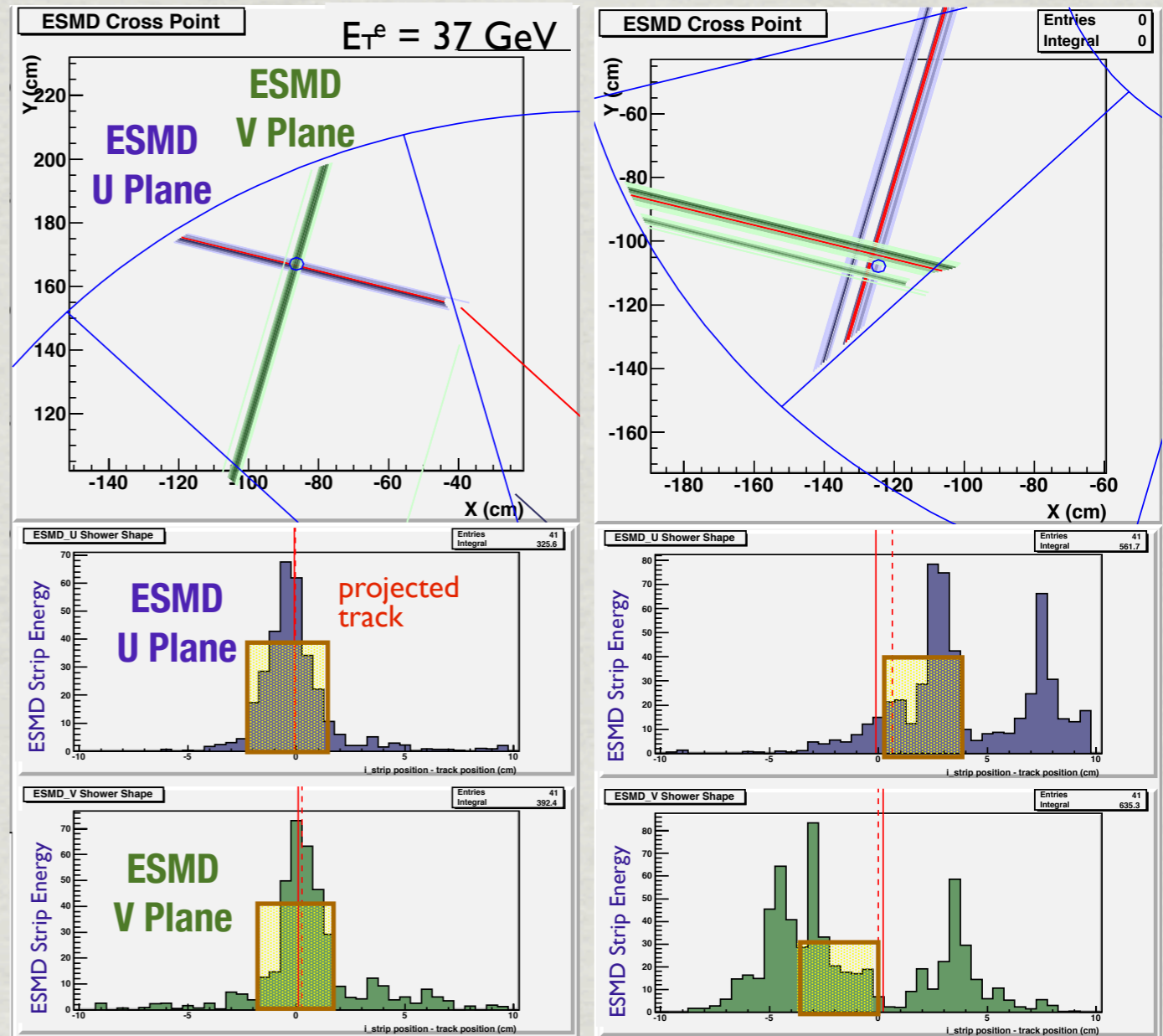
Endcap EM Calorimeter (EEMC)



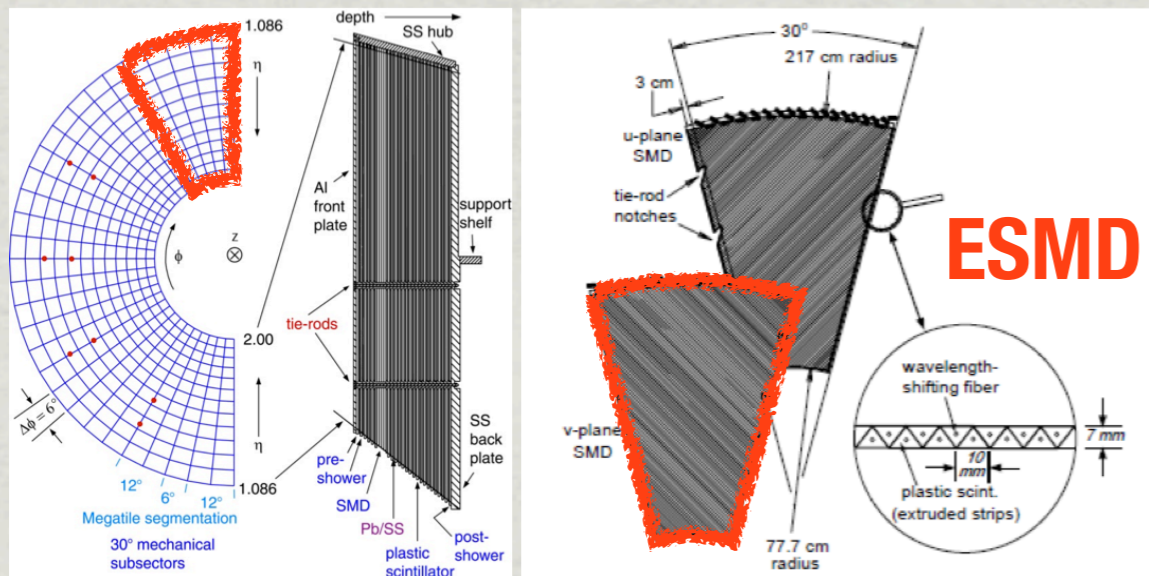
Forward Rapidity Analysis

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 - Utilize TPC which extends to $\eta \sim 1.4$ to reconstruct high p_T TPC track
 - Use isolation ratios and vector p_T imbalance to reduced QCD background
- Improve background rejection by using the Endcap Shower Maximum Detector

2012 data events which satisfy all previous cuts
 Signal Example Background Example



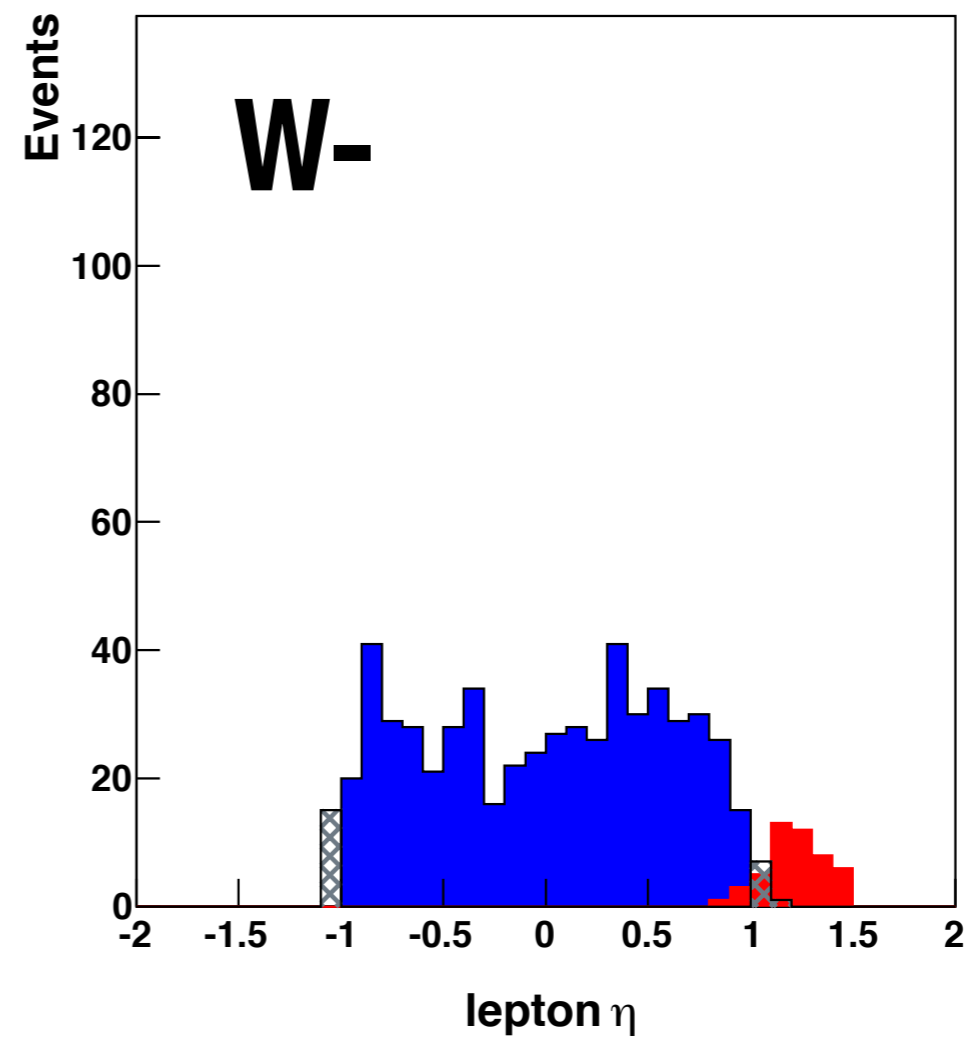
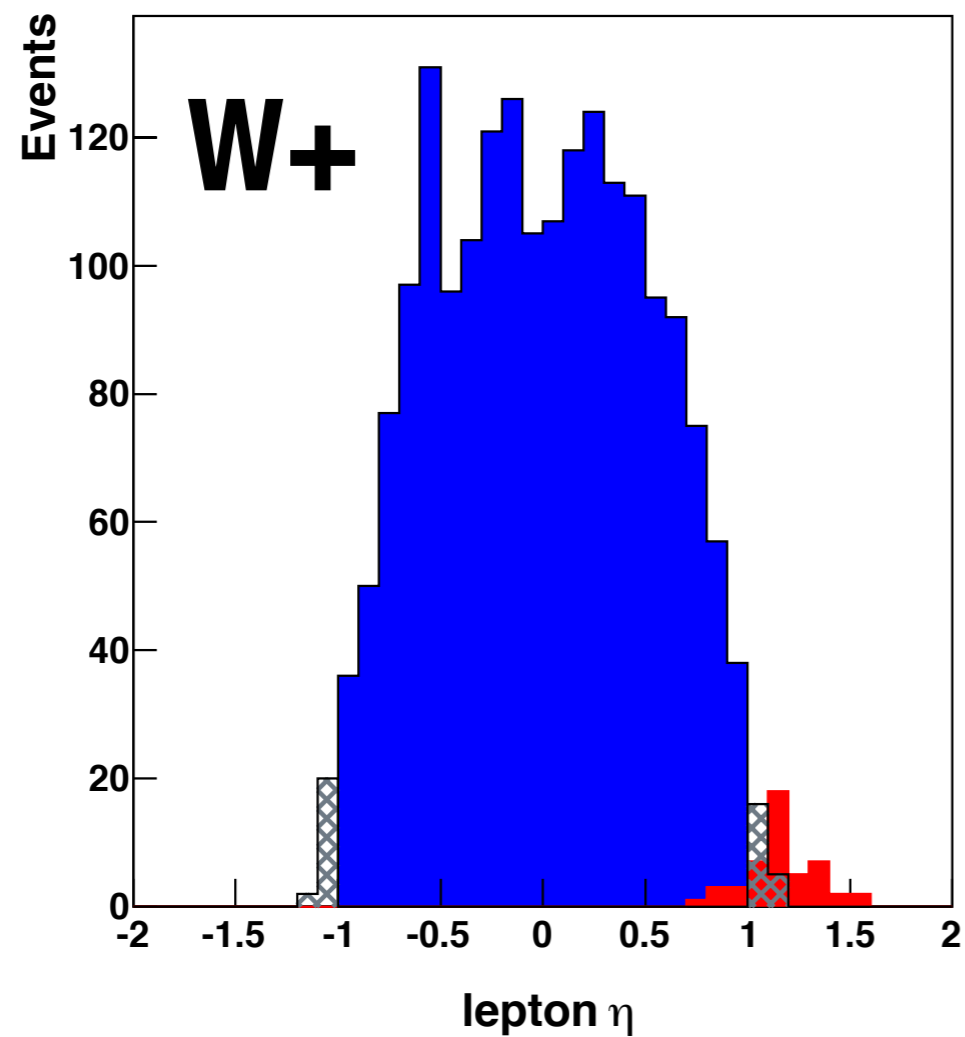
Endcap EM Calorimeter (EEMC)



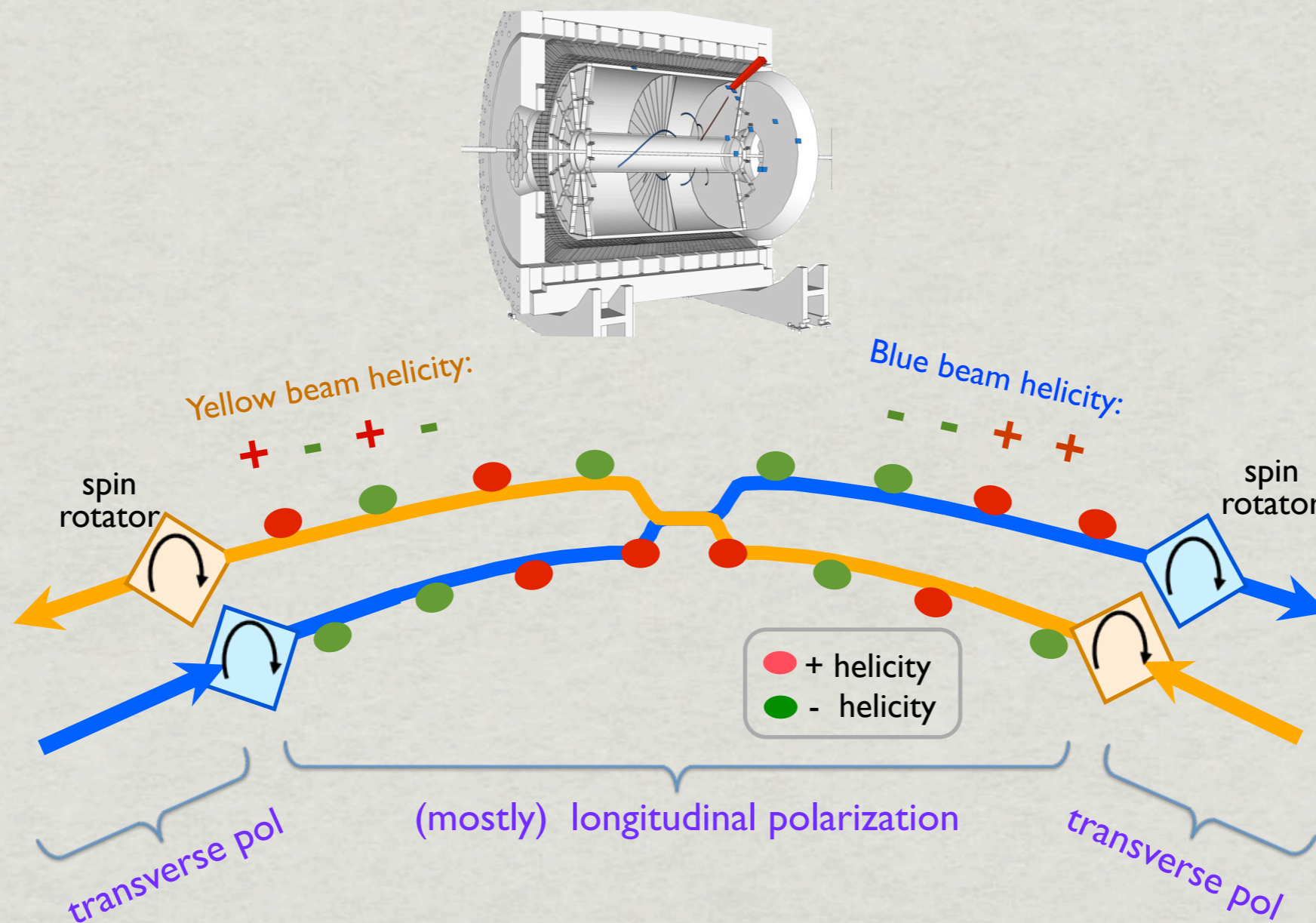
2012 STAR W Candidate Yields vs η

Mid-rapidity (Barrel) Ws

Forward rapidity (Endcap) Ws



Spin Asymmetry: $A_L(\eta_e)$



STAR sees four helicity configurations

Single Asymmetric η Slice

$$\begin{aligned}
 N_{\eta STAR}^{++} / \mathcal{L}^{++} &= C_{\eta STAR} [1 + A_L(\eta) P_1^L + A_L(-\eta) P_2^L + A_{LL}(|\eta|) P_1^L P_2^L] \quad \#1 \\
 N_{\eta STAR}^{+-} / \mathcal{L}^{+-} &= C_{\eta STAR} [1 + A_L(\eta) P_1^L - A_L(-\eta) P_2^L - A_{LL}(|\eta|) P_1^L P_2^L] \quad \#2 \\
 N_{\eta STAR}^{-+} / \mathcal{L}^{-+} &= C_{\eta STAR} [1 - A_L(\eta) P_1^L + A_L(-\eta) P_2^L - A_{LL}(|\eta|) P_1^L P_2^L] \quad \#3 \\
 N_{\eta STAR}^{--} / \mathcal{L}^{--} &= C_{\eta STAR} [1 - A_L(\eta) P_1^L - A_L(-\eta) P_2^L + A_{LL}(|\eta|) P_1^L P_2^L] \quad \#4
 \end{aligned}$$

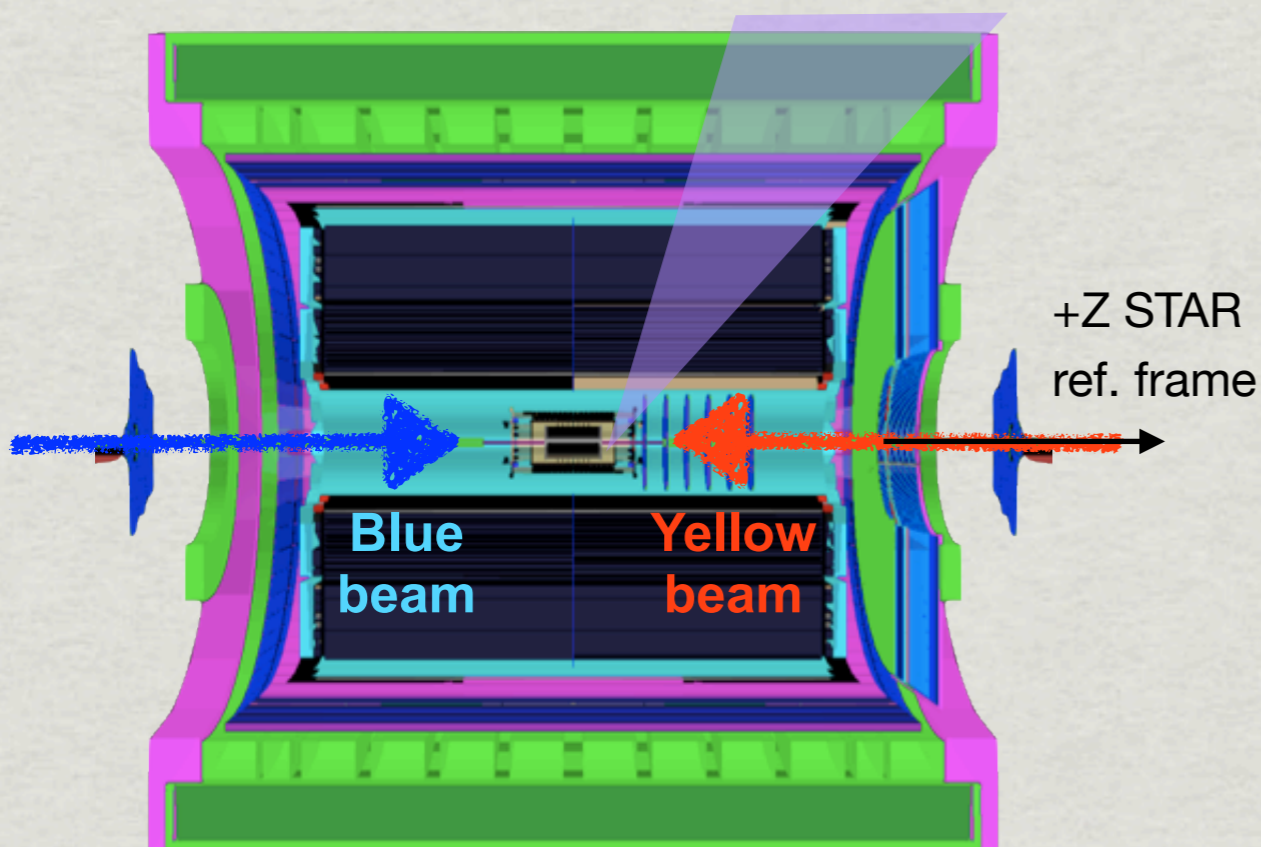
$$A_L = \frac{\sigma_+ - \sigma_-}{\sigma_+ + \sigma_-} = \frac{1}{P} \frac{N_+ / \mathcal{L}_+ - N_- / \mathcal{L}_-}{N_+ / \mathcal{L}_+ + N_- / \mathcal{L}_-}$$

Blue beam
pol: $A_L(+\eta)$

Yellow beam
pol: $A_L(-\eta)$

$A_{LL} |\eta|$

Single η slice



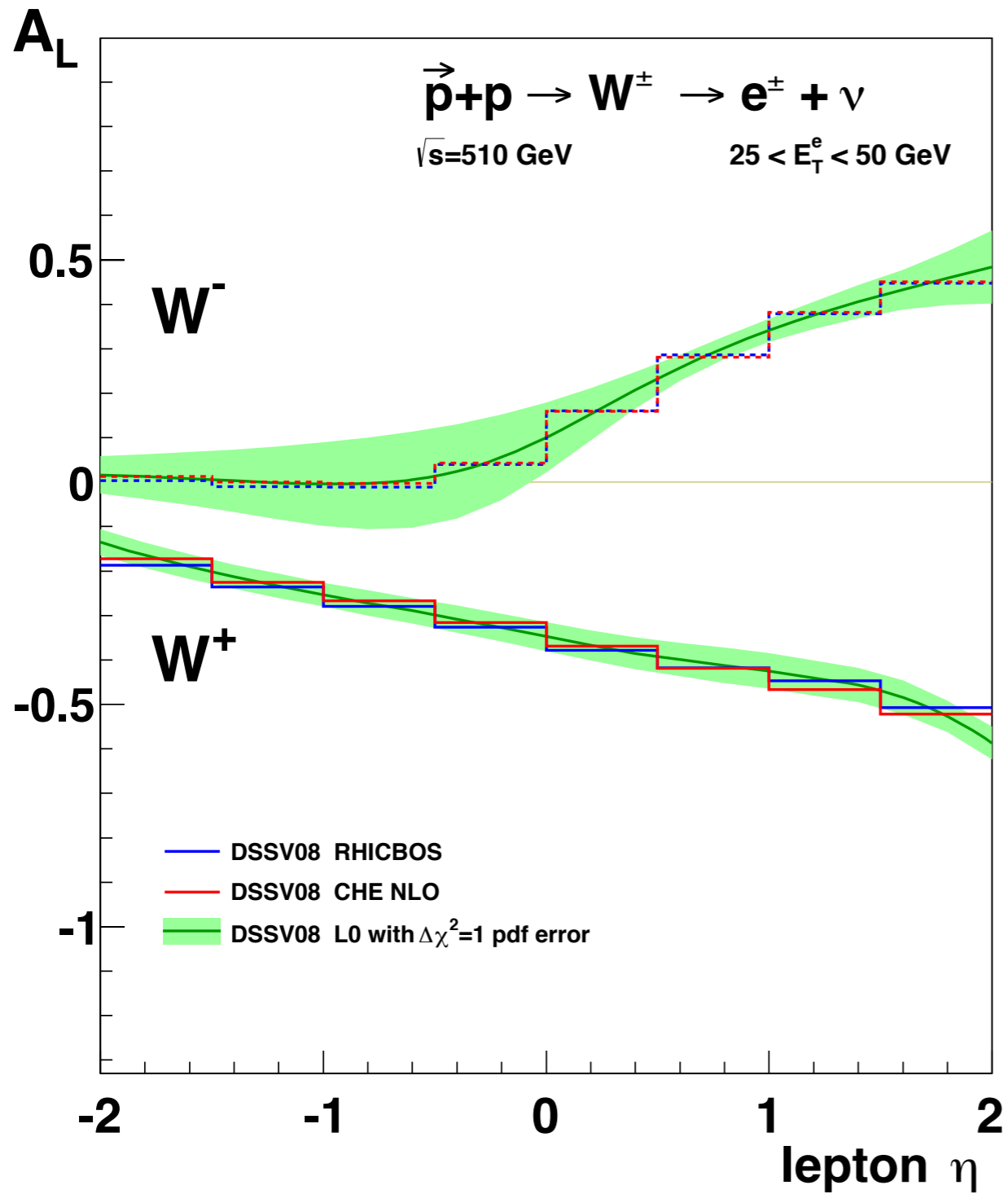
A_L for both beams measured from 4 spin dependent yields

$$A_L^{sig}(+\eta) = \frac{1 + 2 - 3 - 4}{P_1^L \cdot \Sigma 1 \dots 4}$$

$$A_L^{sig}(-\eta) = \frac{1 - 2 + 3 - 4}{P_2^L \cdot \Sigma 1 \dots 4}$$

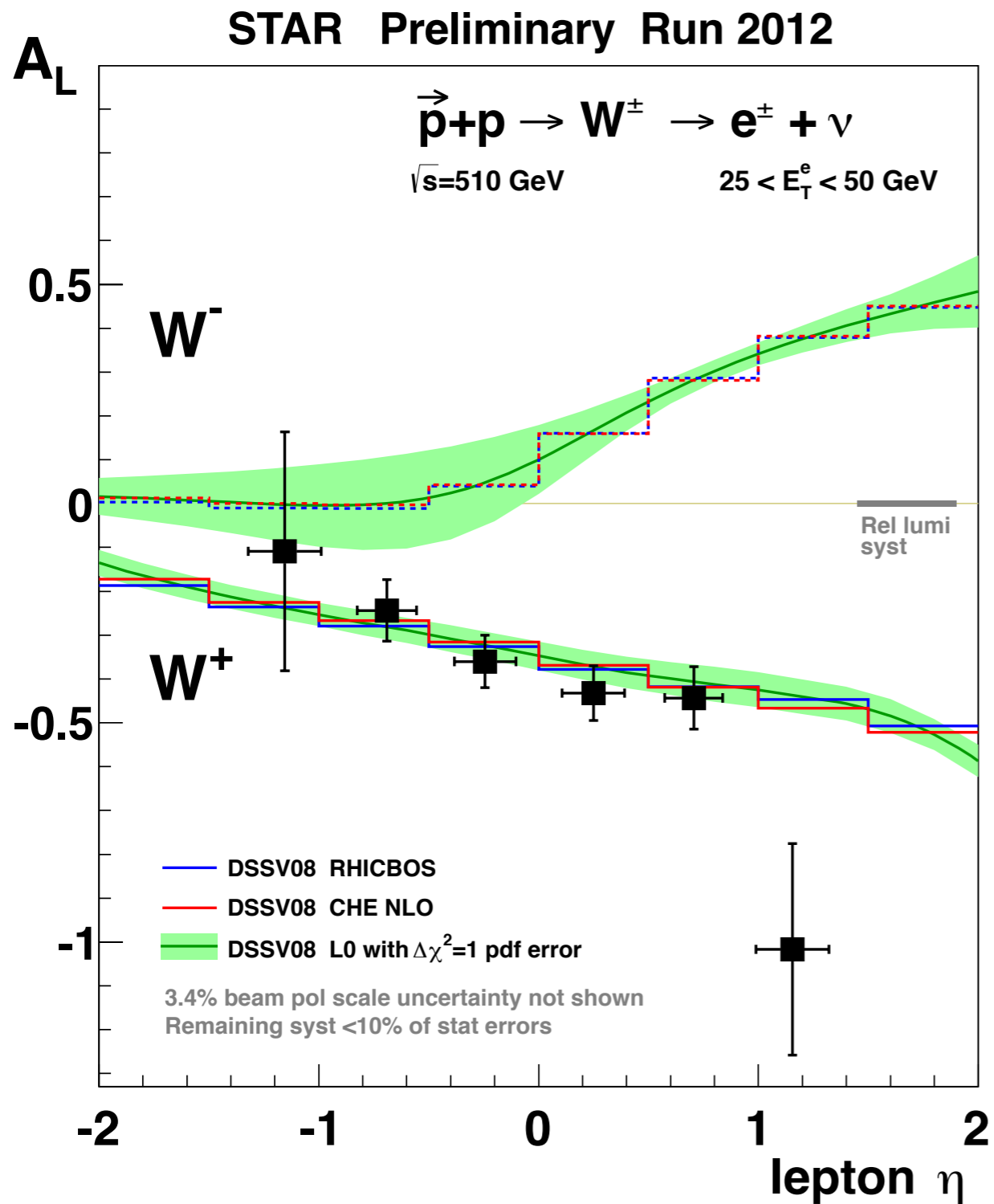
NEW!

STAR 2012 $W A_L(\eta)$



NEW!

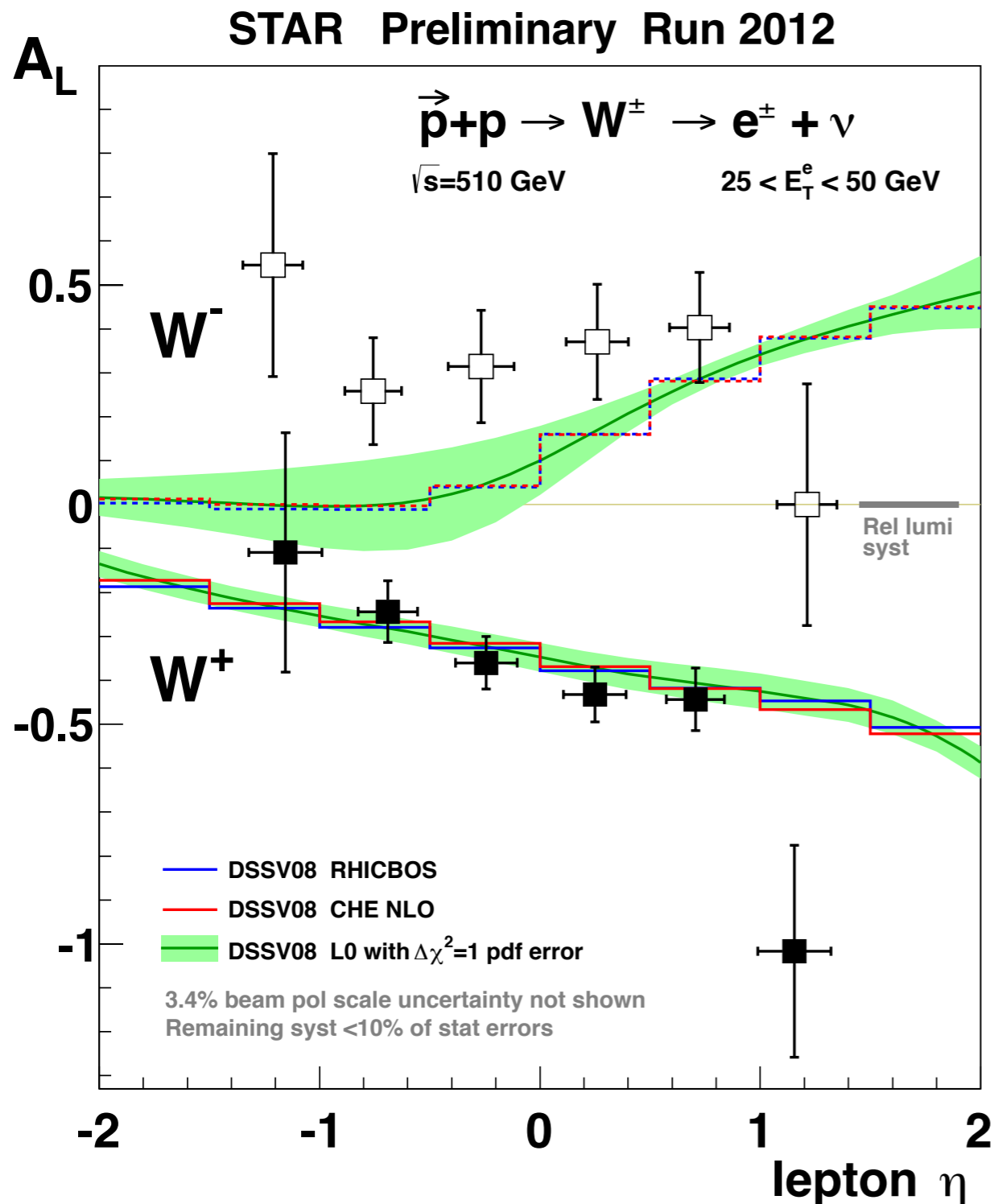
STAR 2012 $W A_L(\eta)$



- * $A_L(W^+)$ is consistent with theoretical predictions using the DSSV polarized PDFs

NEW!

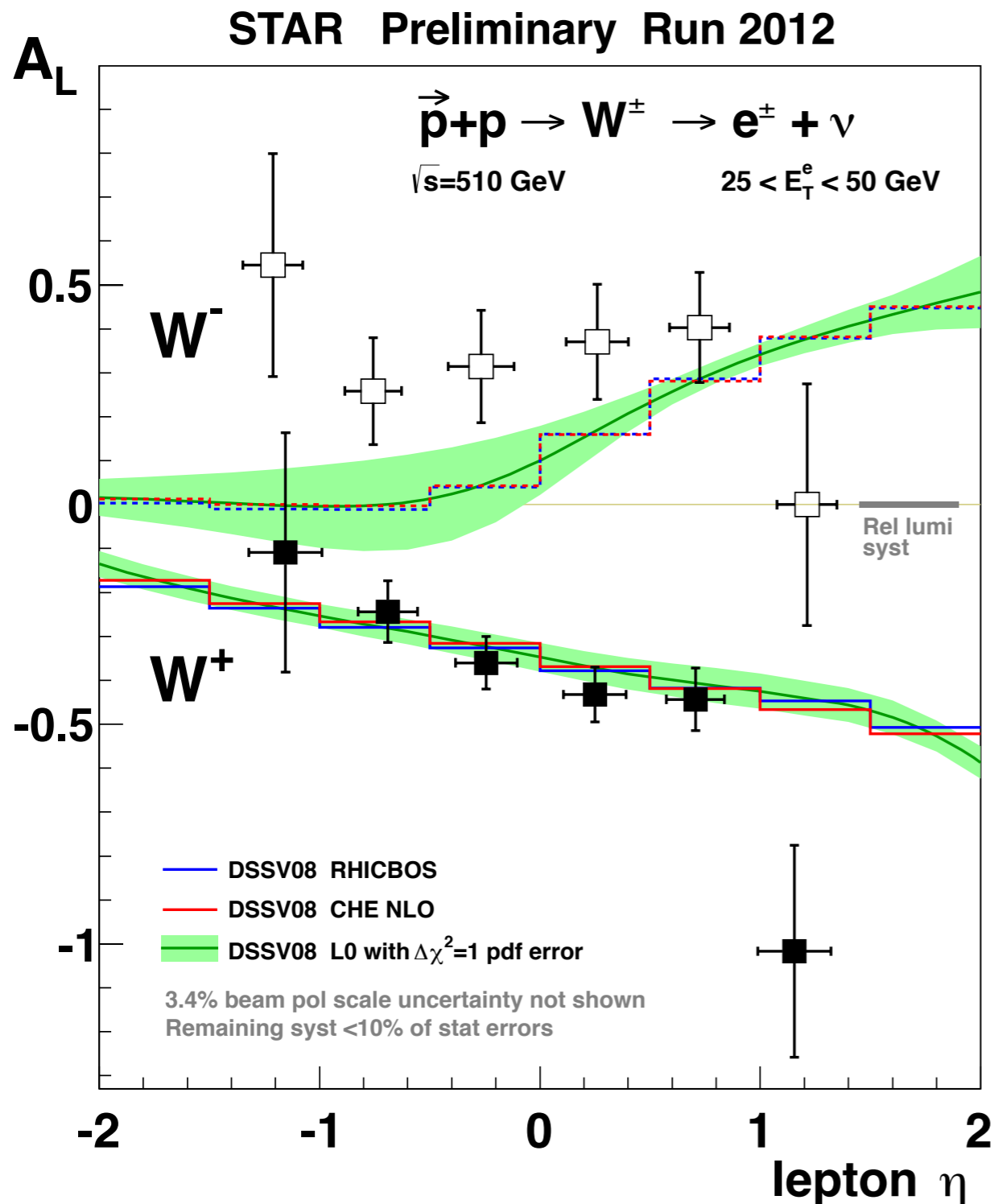
STAR 2012 $W A_L(\eta)$



- * $A_L(W^-)$ is systematically larger than the DSSV predictions
- * The enhancement at $\eta_e < 0$, in particular, is sensitive to the $\Delta\bar{u}$ polarized antiquark distribution
- * $A_L(W^+)$ is consistent with theoretical predictions using the DSSV polarized PDFs

NEW!

STAR 2012 $W A_L(\eta)$

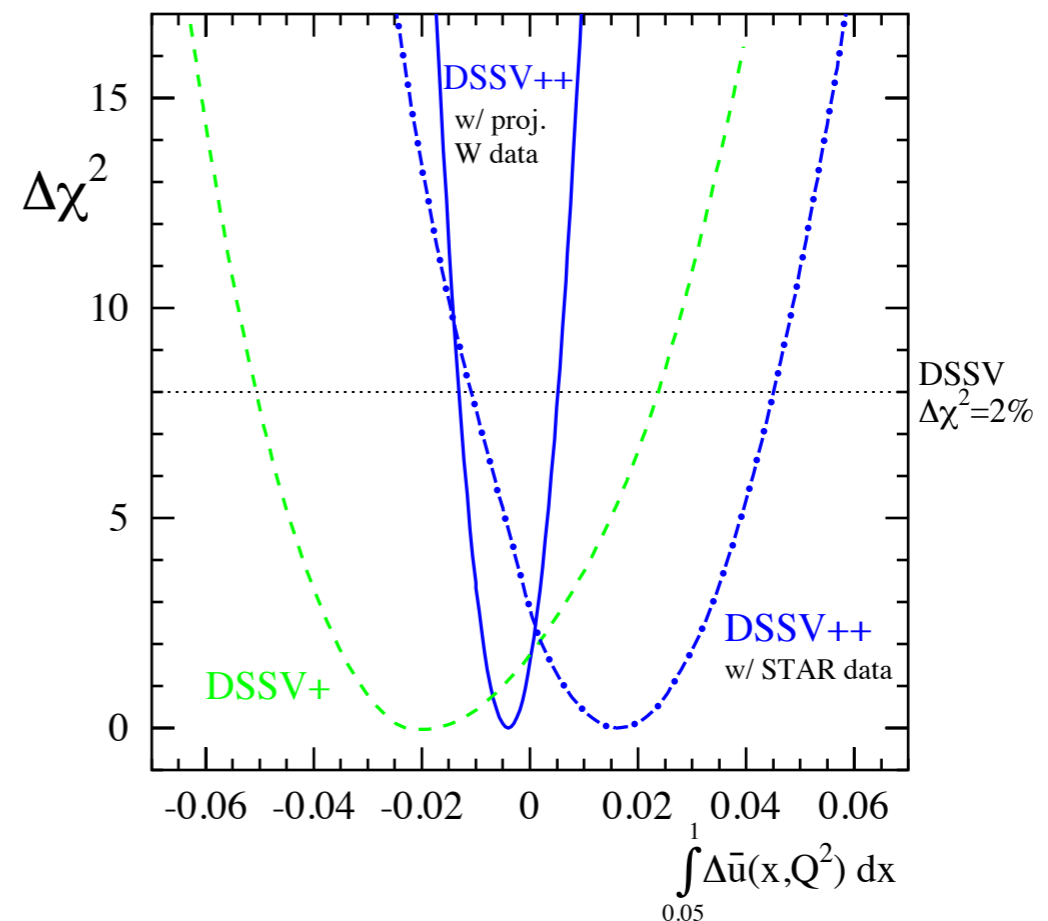
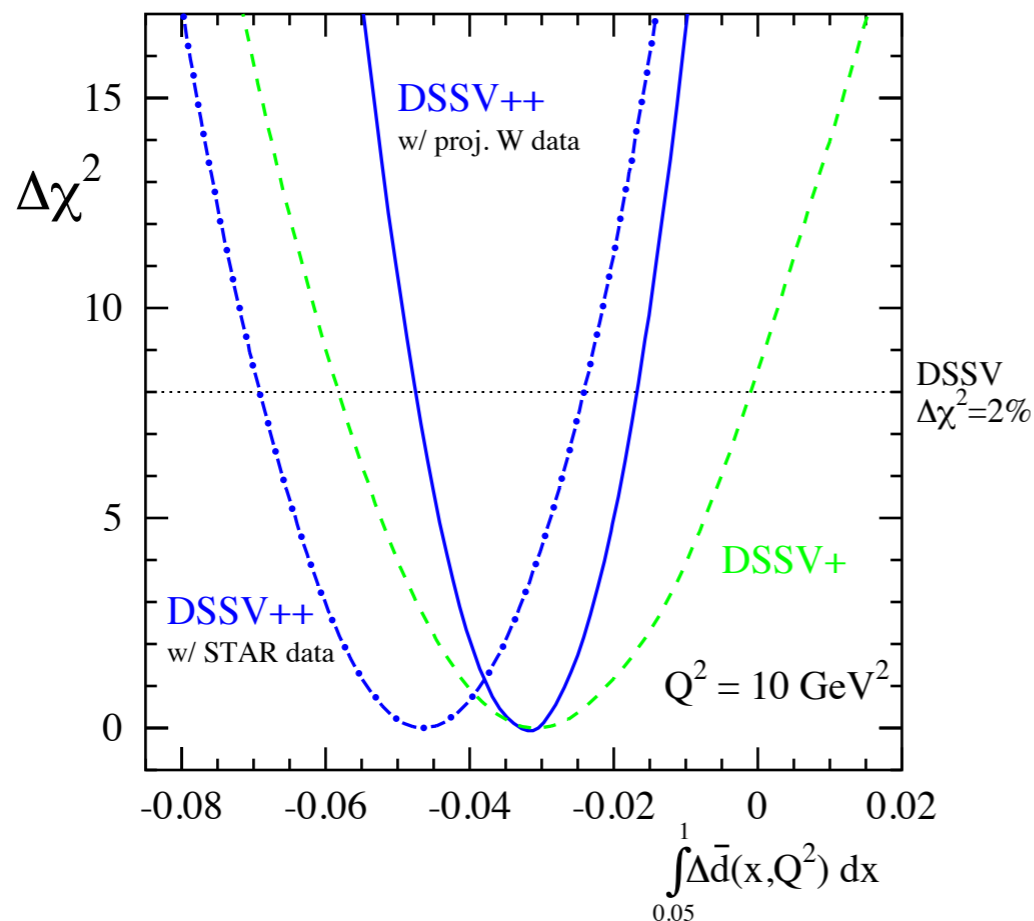
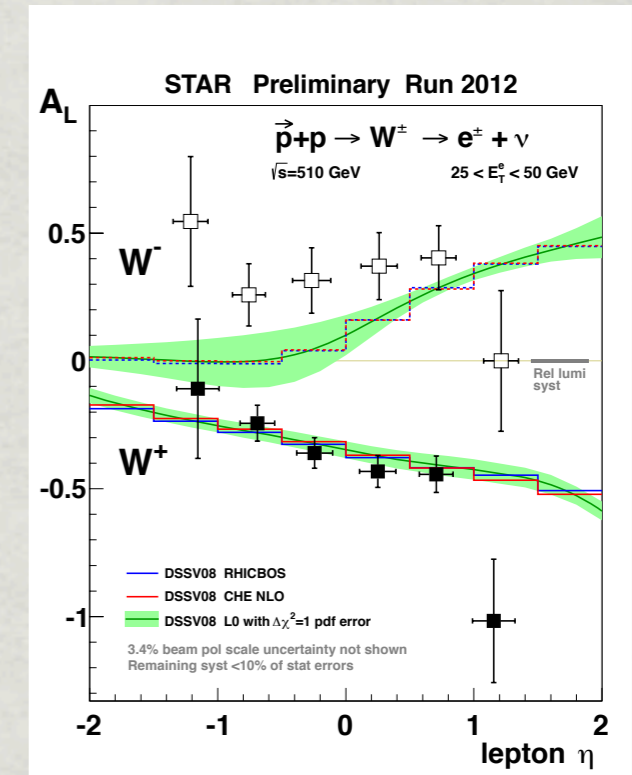


- * $A_L(W^-)$ is systematically larger than the DSSV predictions
- * The enhancement at $\eta_e < 0$, in particular, is sensitive to the $\Delta\bar{u}$ polarized antiquark distribution
- * $A_L(W^+)$ is consistent with theoretical predictions using the DSSV polarized PDFs
- * The systematic uncertainties for A_L are well under control for $|\eta_e| < 1.4$

NEW!

DSSV++ Global Analysis

- ✱ DSSV++ is a new preliminary global analysis from the DSSV group that includes RHIC 2009 A_{LL} data and STAR 2012 $W A_L$ data
- ✱ Significant shift in $\Delta\bar{u}$ due to $A_L W^-$

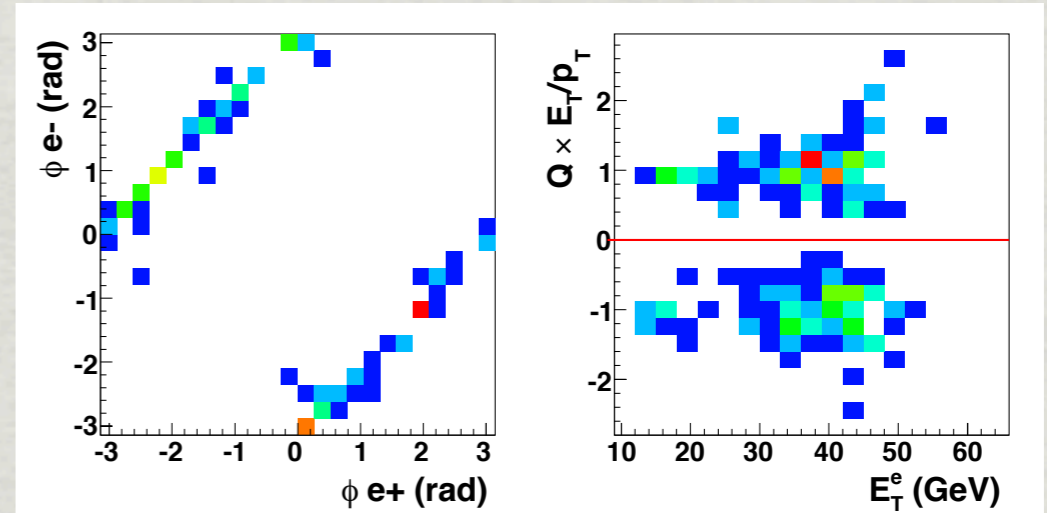
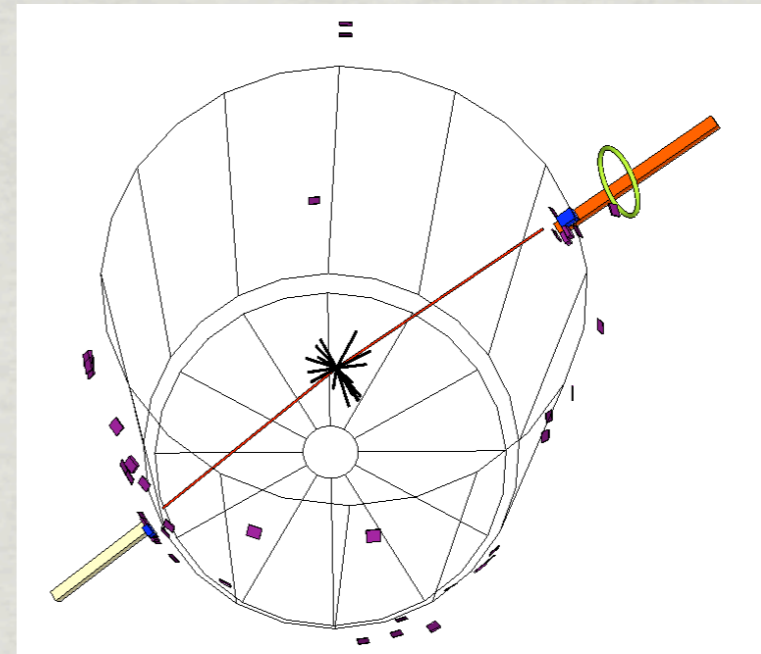


Special thanks to the DSSV group!

NEW!

STAR 2012 Z A_L

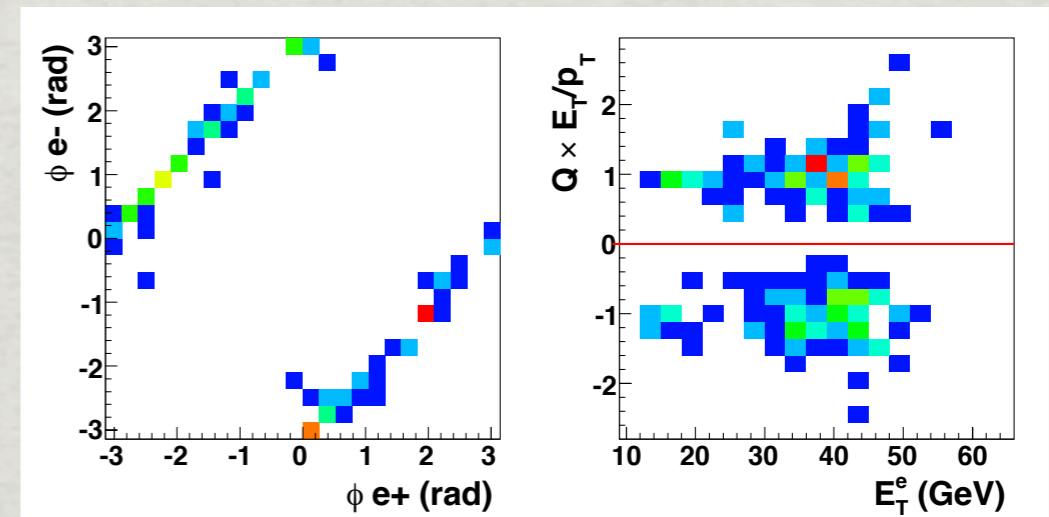
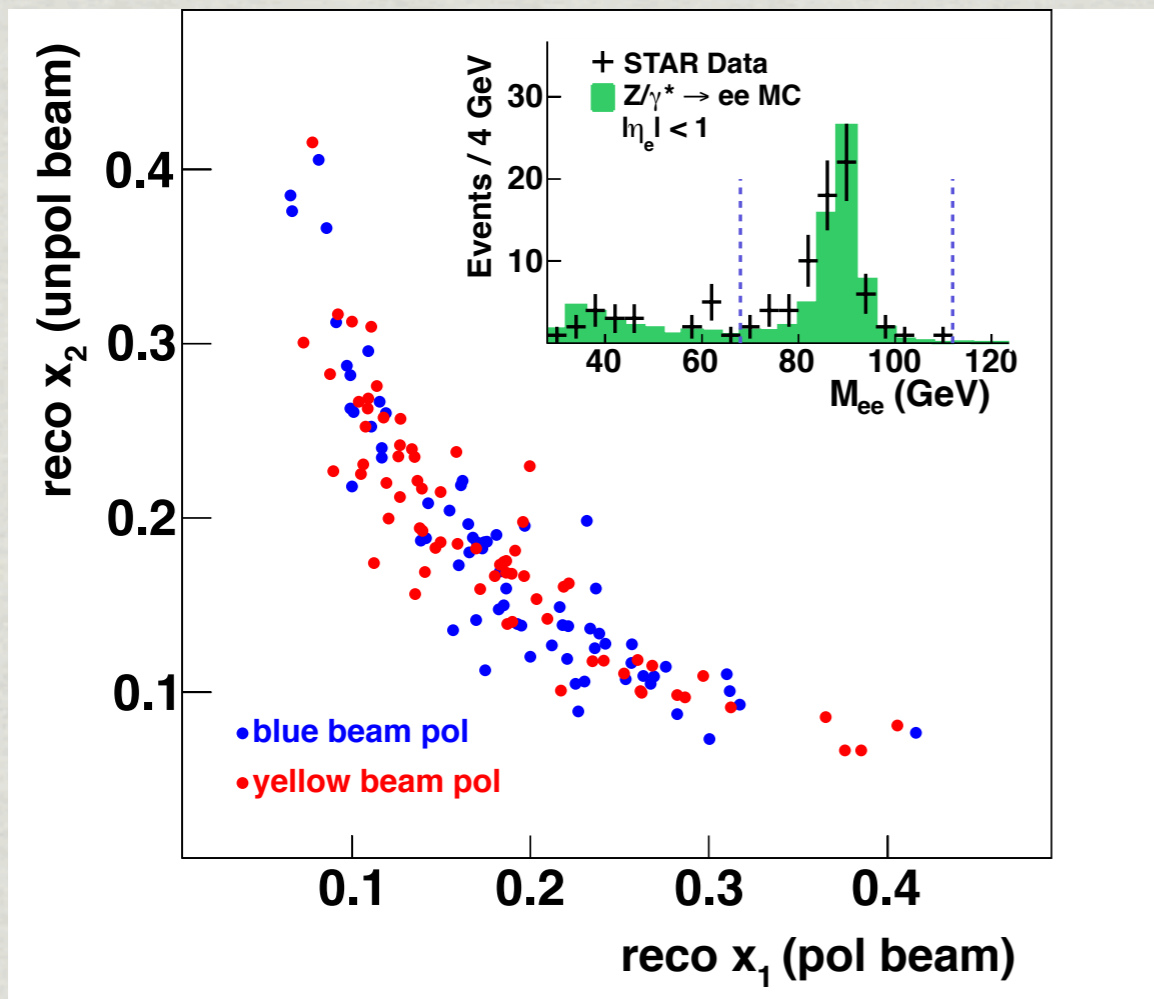
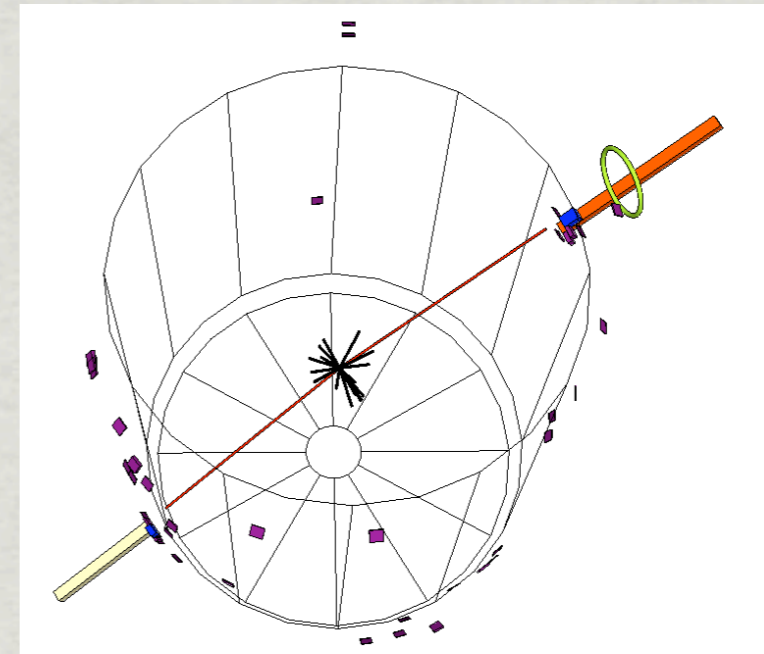
Z → e⁺e⁻ Candidate



NEW!

STAR 2012 Z A_L

Z → e⁺e⁻ Candidate



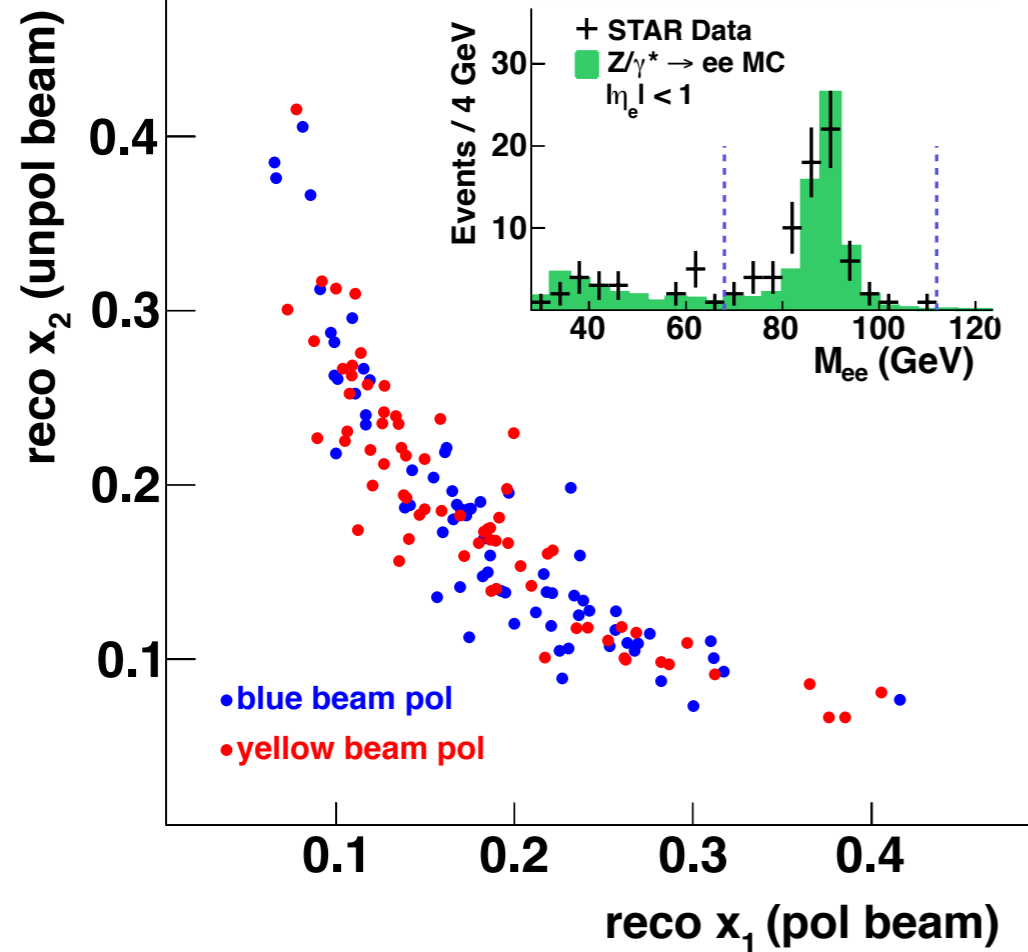
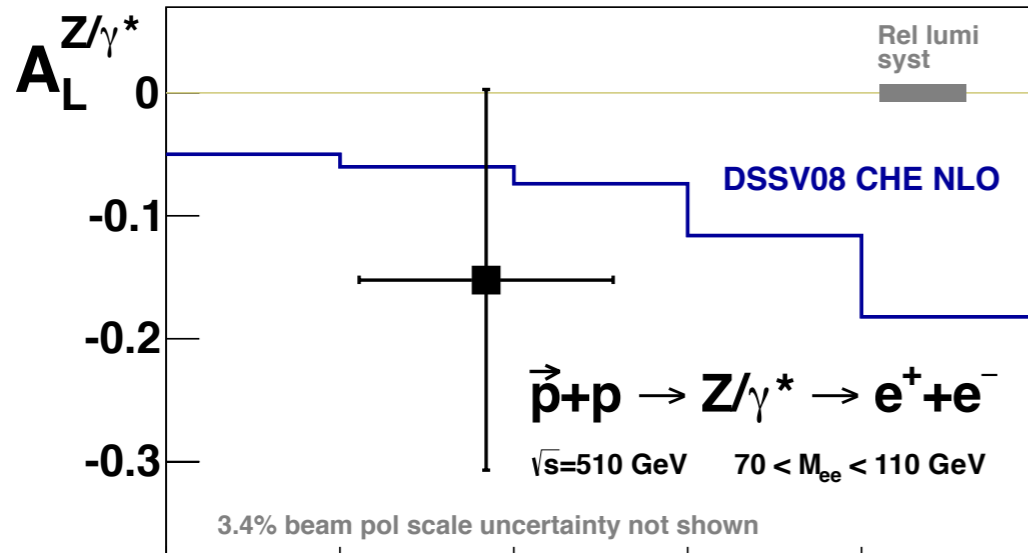
Reconstruct initial state kinematics at leading order:

$$x_{1(2)} = \frac{M_{ee}}{\sqrt{s}} e^{\pm yz}$$

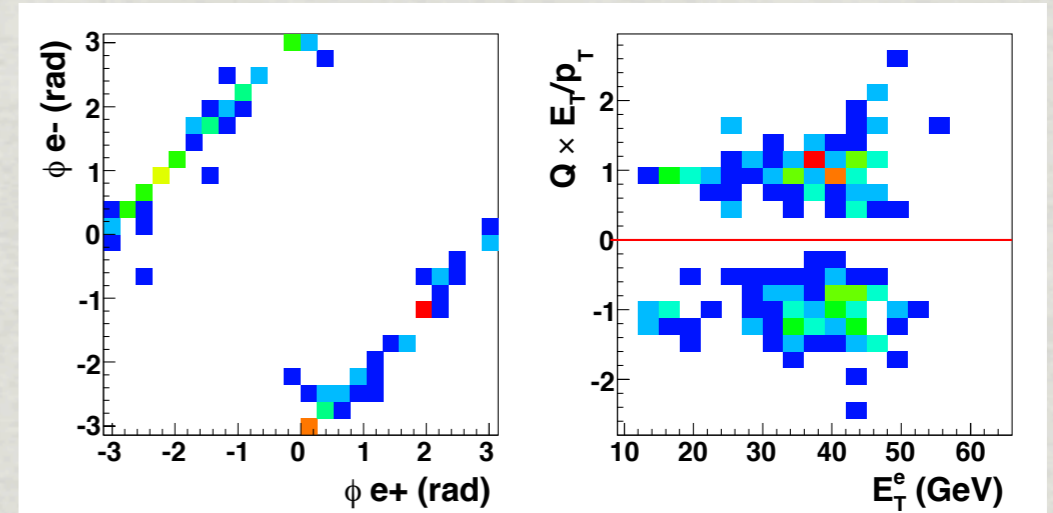
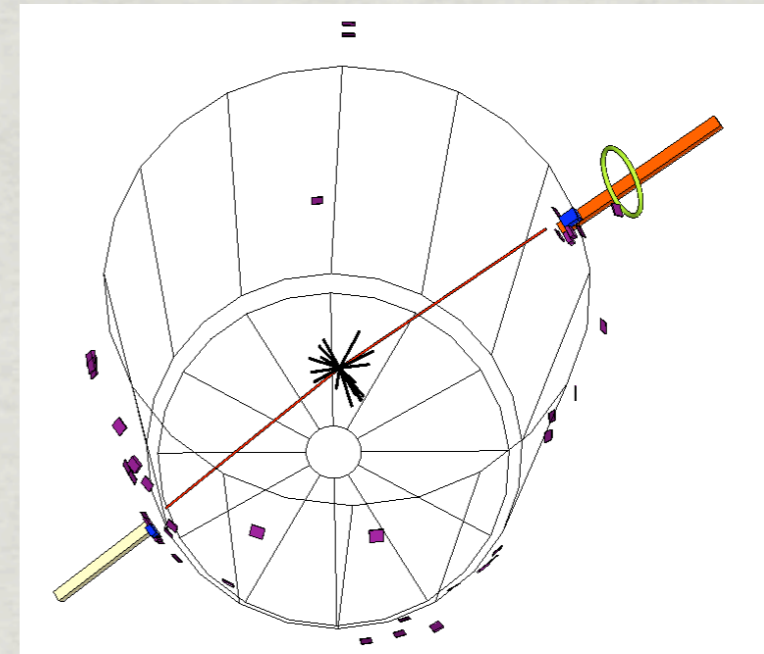
NEW!

STAR 2012 Z A_L

STAR Preliminary Run 2012



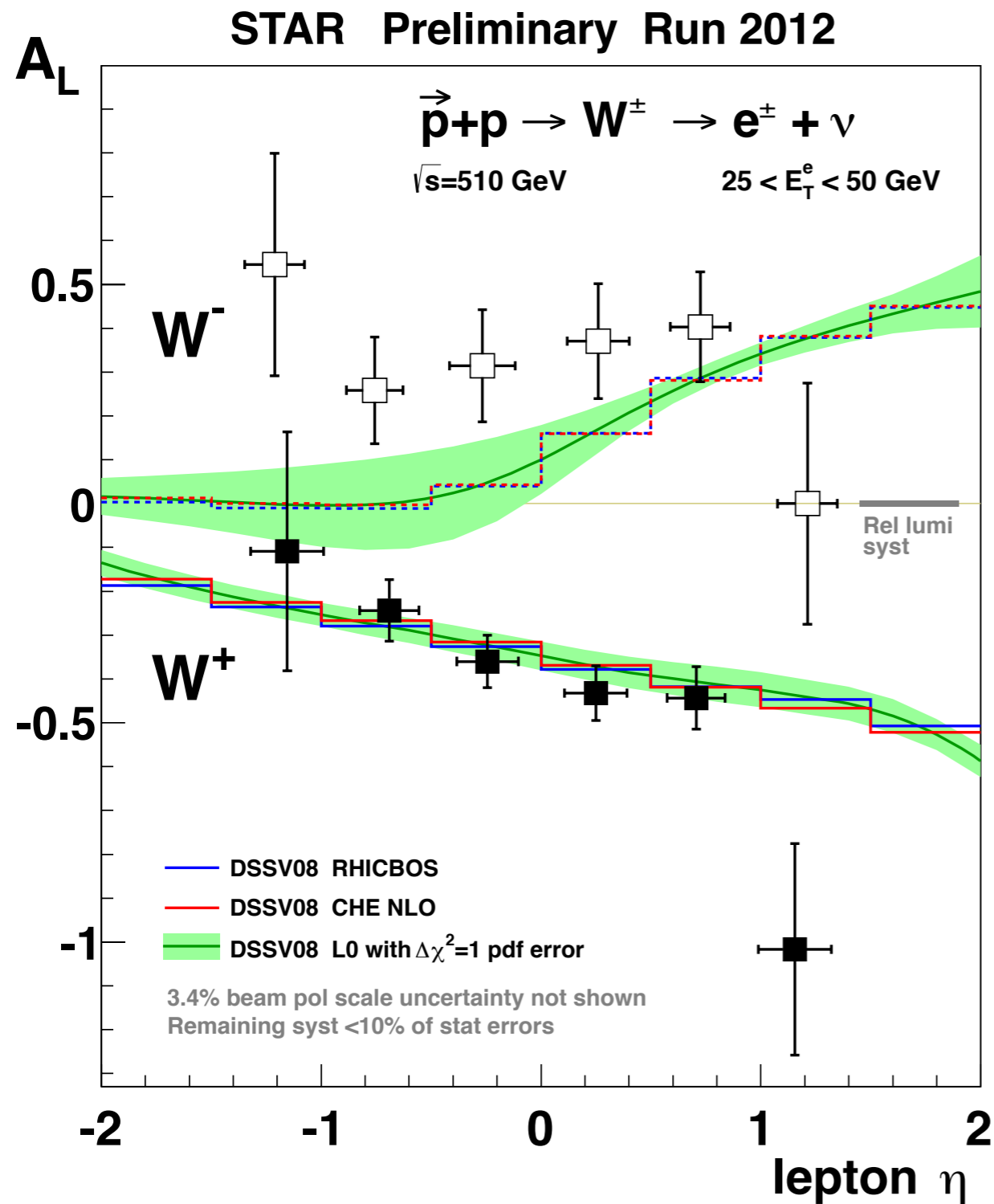
Z → e⁺e⁻ Candidate



Reconstruct initial state kinematics at leading order:

$$x_{1(2)} = \frac{M_{ee}}{\sqrt{s}} e^{\pm y_z}$$

Summary



- * The production of W bosons in polarized p+p collisions provides a new means of studying the spin and flavor asymmetries of the proton sea quark distributions
- * STAR has measured the parity-violating single-spin asymmetry A_L for $|\eta_e| < 1.4$ from Run 12 data, providing the first detailed look at the asymmetry's η_e dependence
- * A_L for Z/γ^* production was also measured, and is consistent with the theoretical predictions