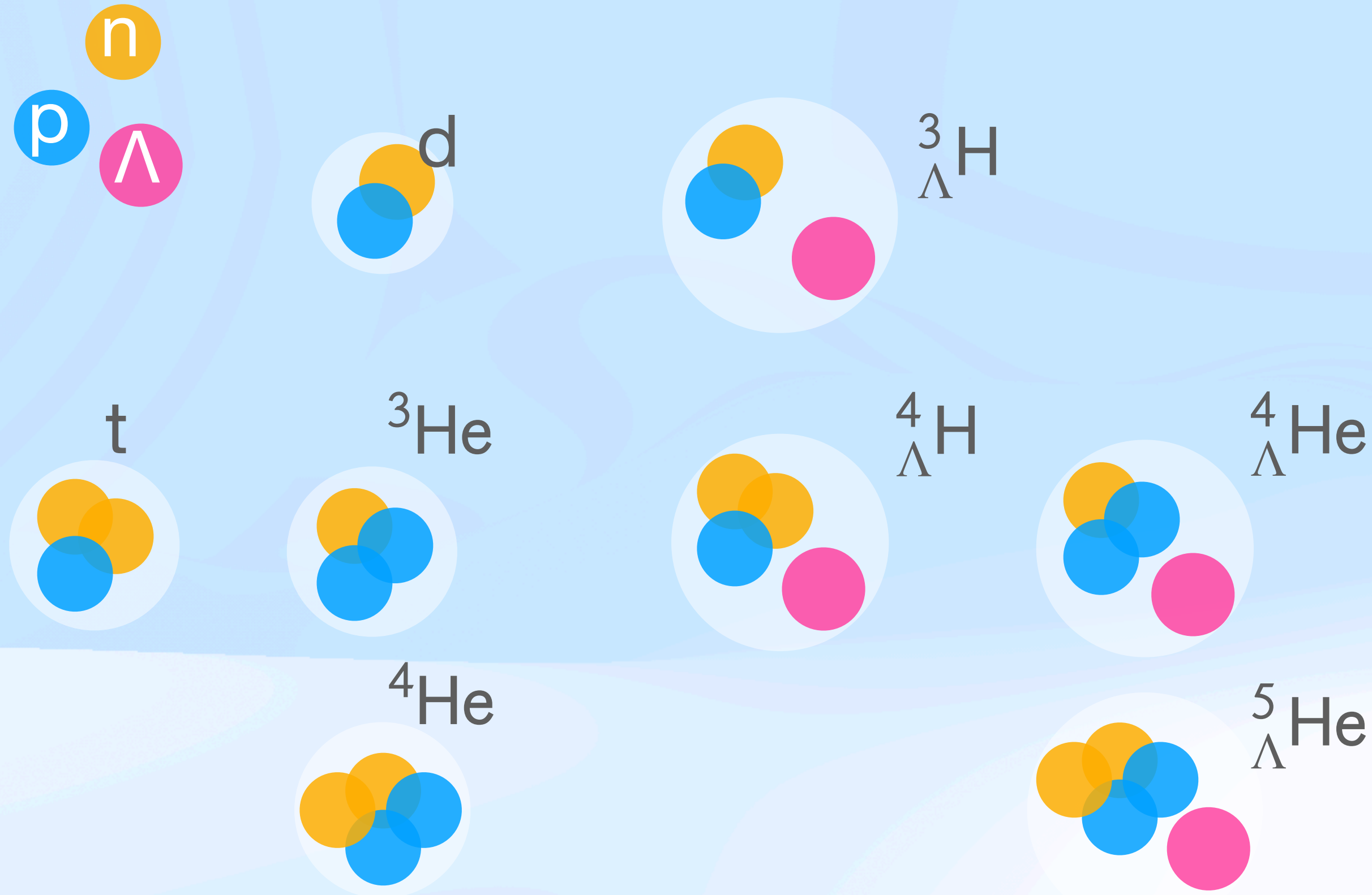




New Hypernuclei Measurements from STAR

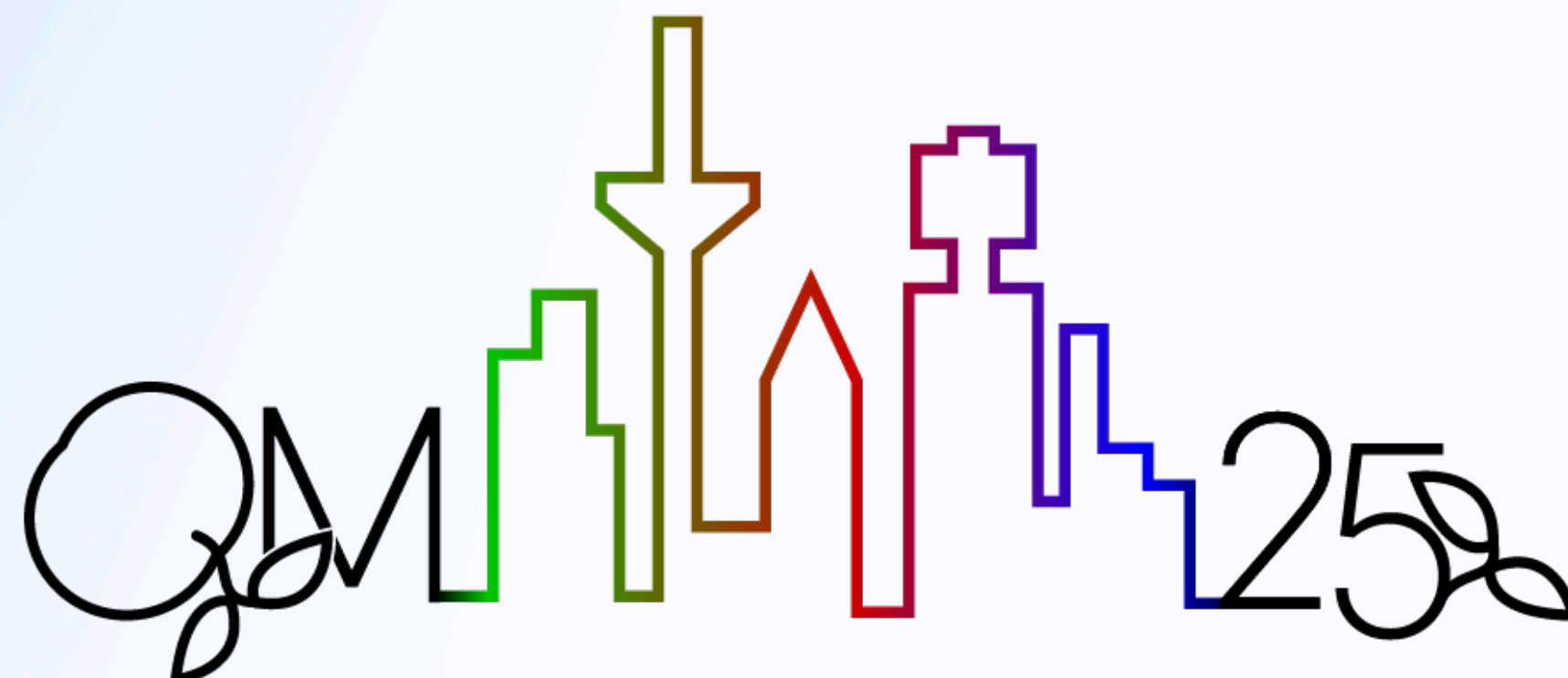


Outline

1. Introduction
2. Yields and Yield Ratios
3. Transverse Momentum Distribution
4. Collective Flow
5. Summary and Outlook

*Yingjie Zhou for the STAR Collaboration
CCNU, GSI*

Quark Matter 2025, Frankfurt, Germany



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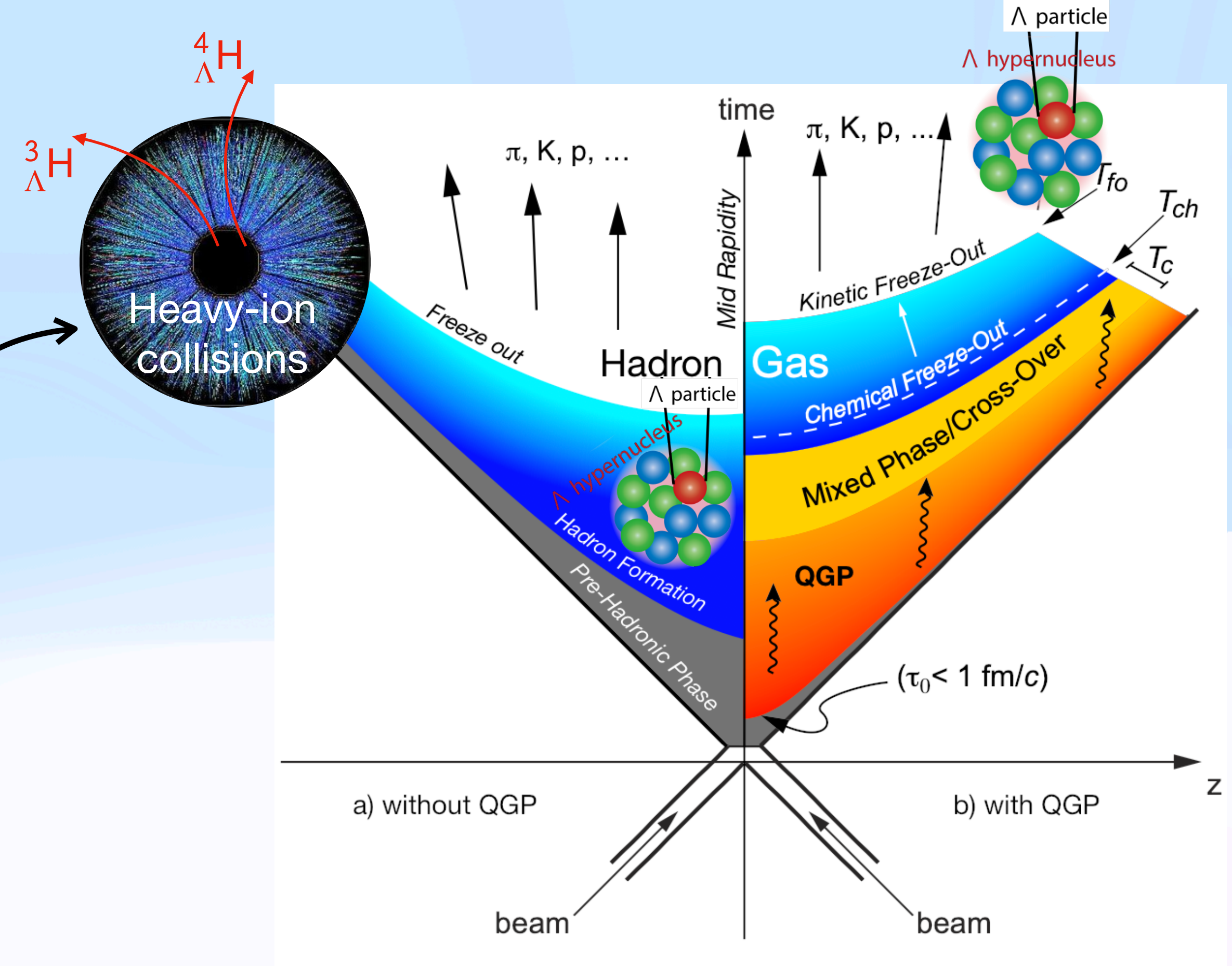
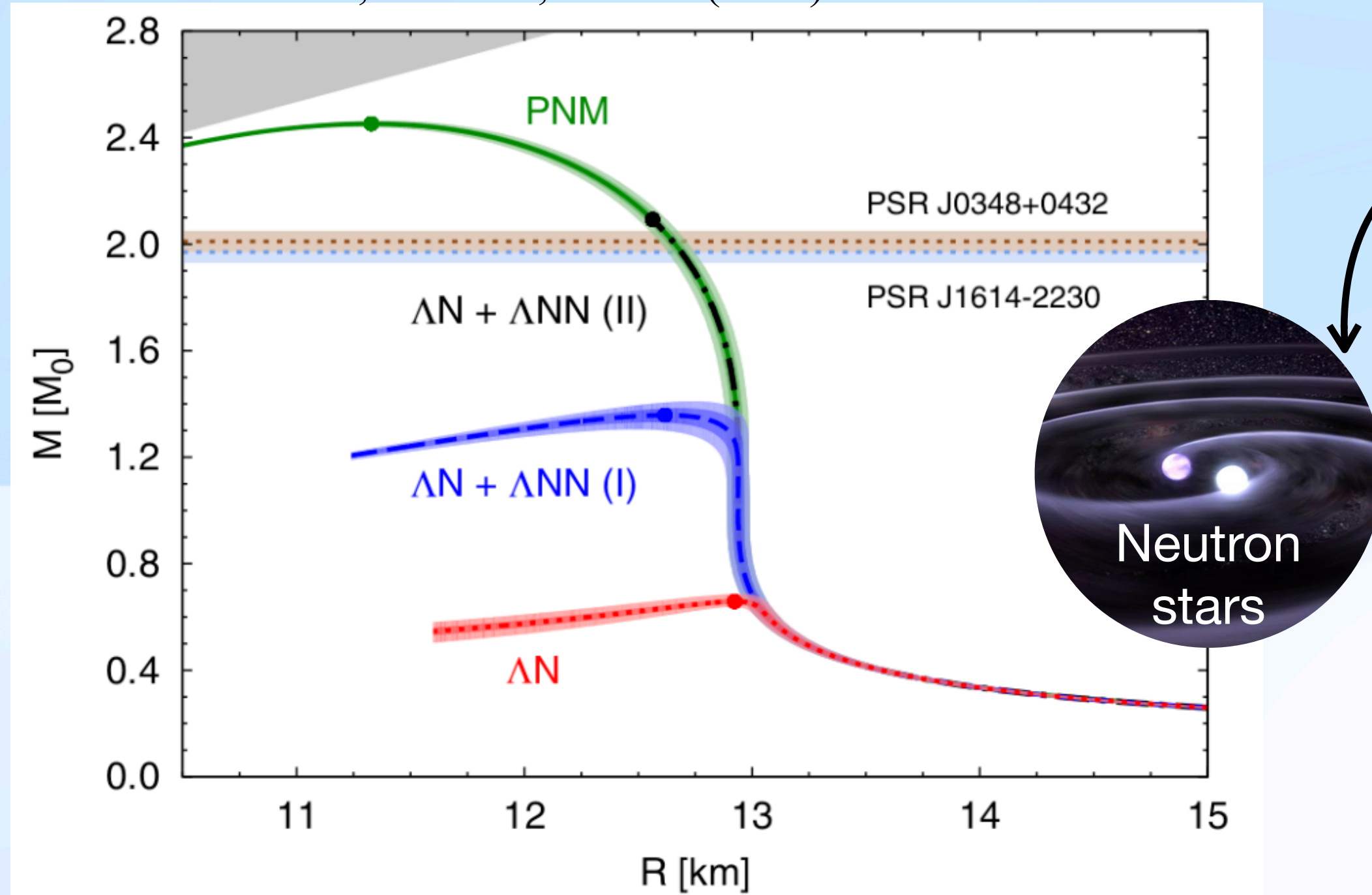
Office of
Science



Hypernuclei and Hyperon-Nucleon (Y-N) Interaction

- **Hyperon Puzzle:** difficulty to reconcile the measured masses of neutron stars with the presence of hyperons in their interiors

D. Lonardonì et al., PRL 114, 092301 (2015)



- Density dependent YN, YNN interactions are essential for solving the hyperon puzzle

- Can hypernuclei production be used to constrain the in-medium Y-N interaction?

*Need a solid understanding in **hypernuclei production mechanisms** before we can use them as probes for medium properties*

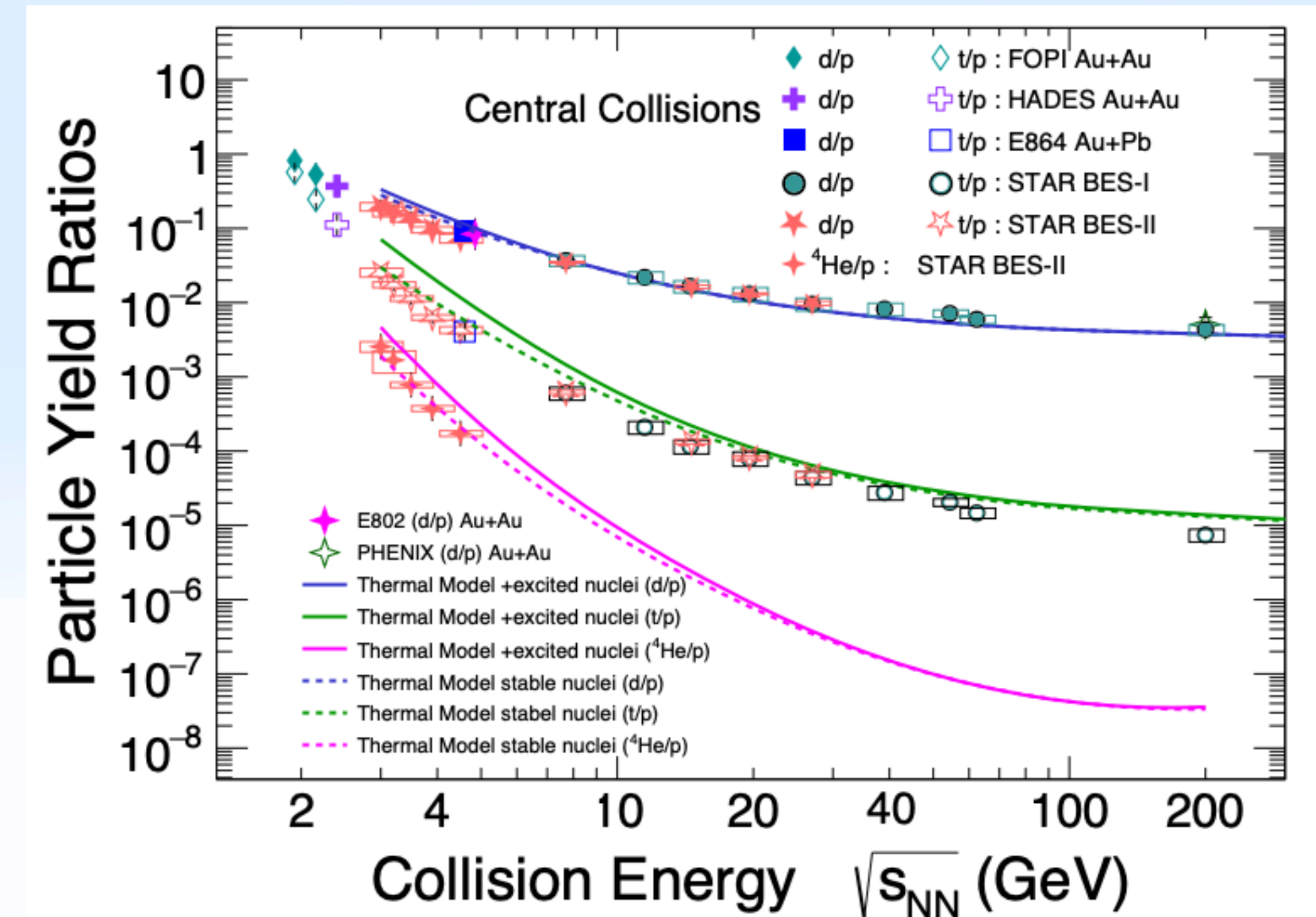
Hypernuclei Production Mechanisms

See poster by Liubing Chen (xx/xx)

When are nuclei produced in a heavy-ion collision?

- Thermal models
 - Hadrons and (hyper-)nuclei are treated equally
 - Yields are predicted with thermal equilibrium assumptions
- Coalescence model
 - (Hyper-)nuclei formation after kinetic freeze-out
 - Nucleon coalescence
 - Wigner function
 - The emission source size and the nuclear radius affect the yields

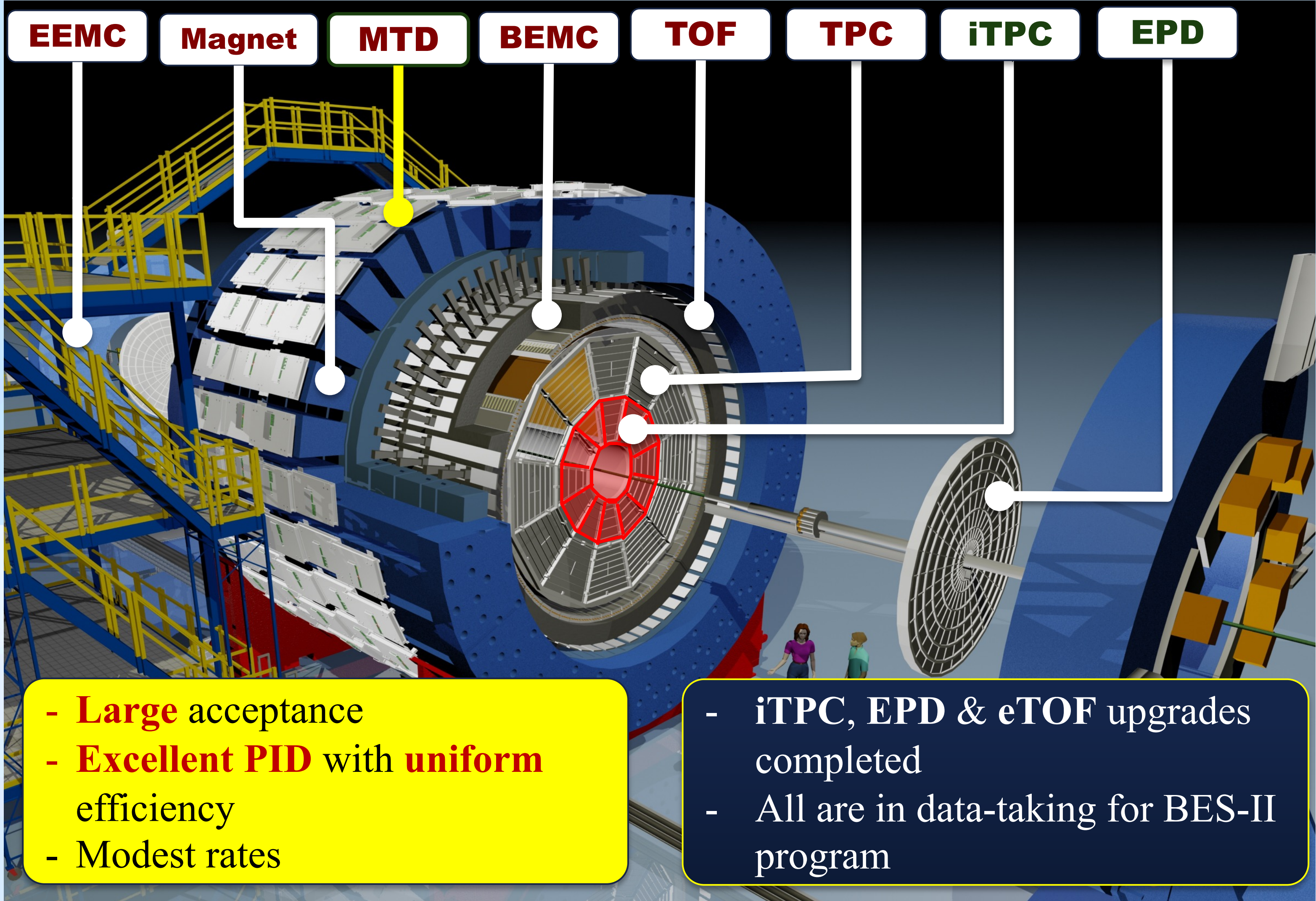
What have we learnt from light nuclei production?



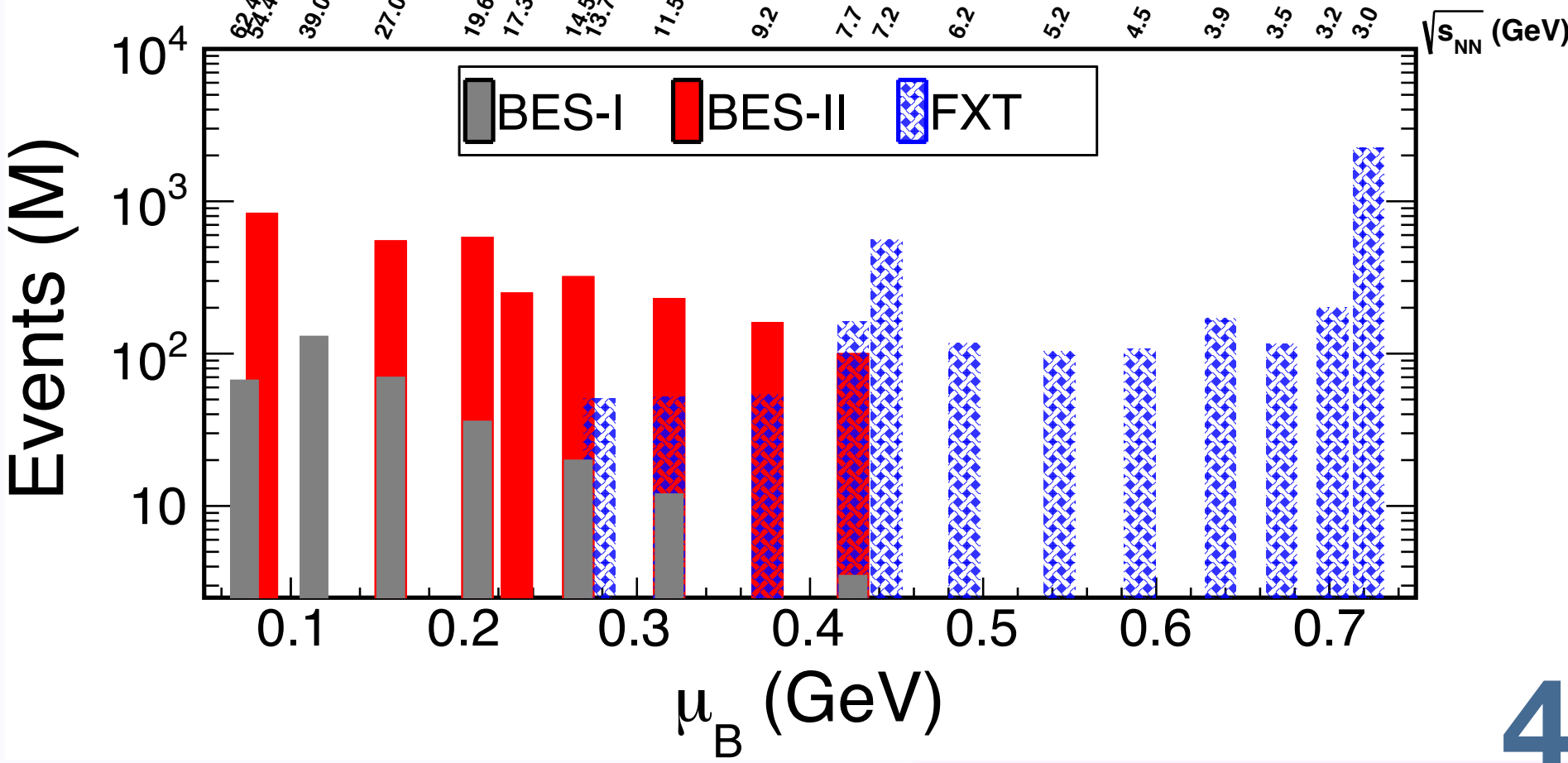
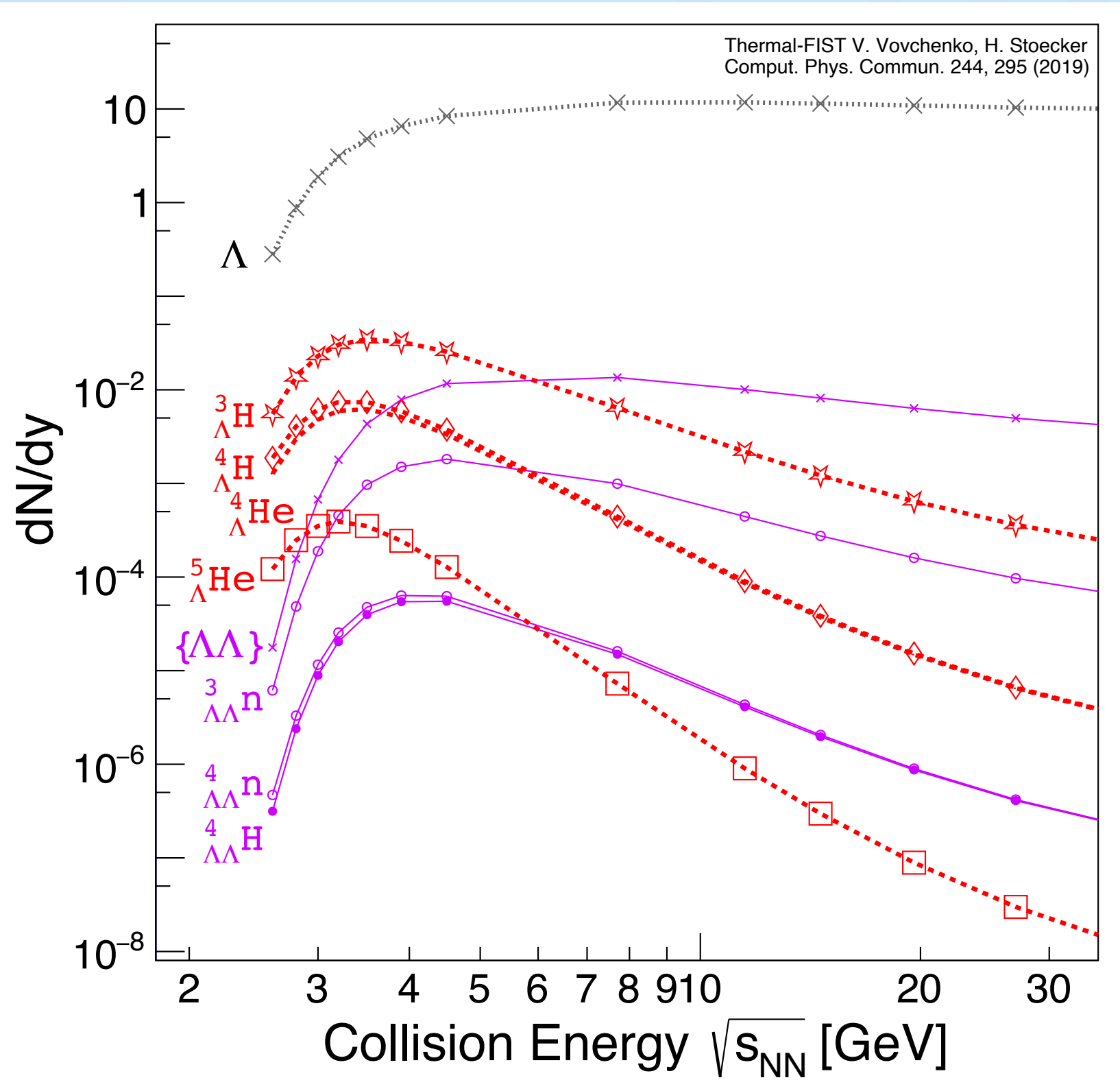
- d/p is fairly well described by thermal model, but t/p , $^4\text{He}/p$ is overestimated

Recent nuclei measurements poses challenges for thermal model

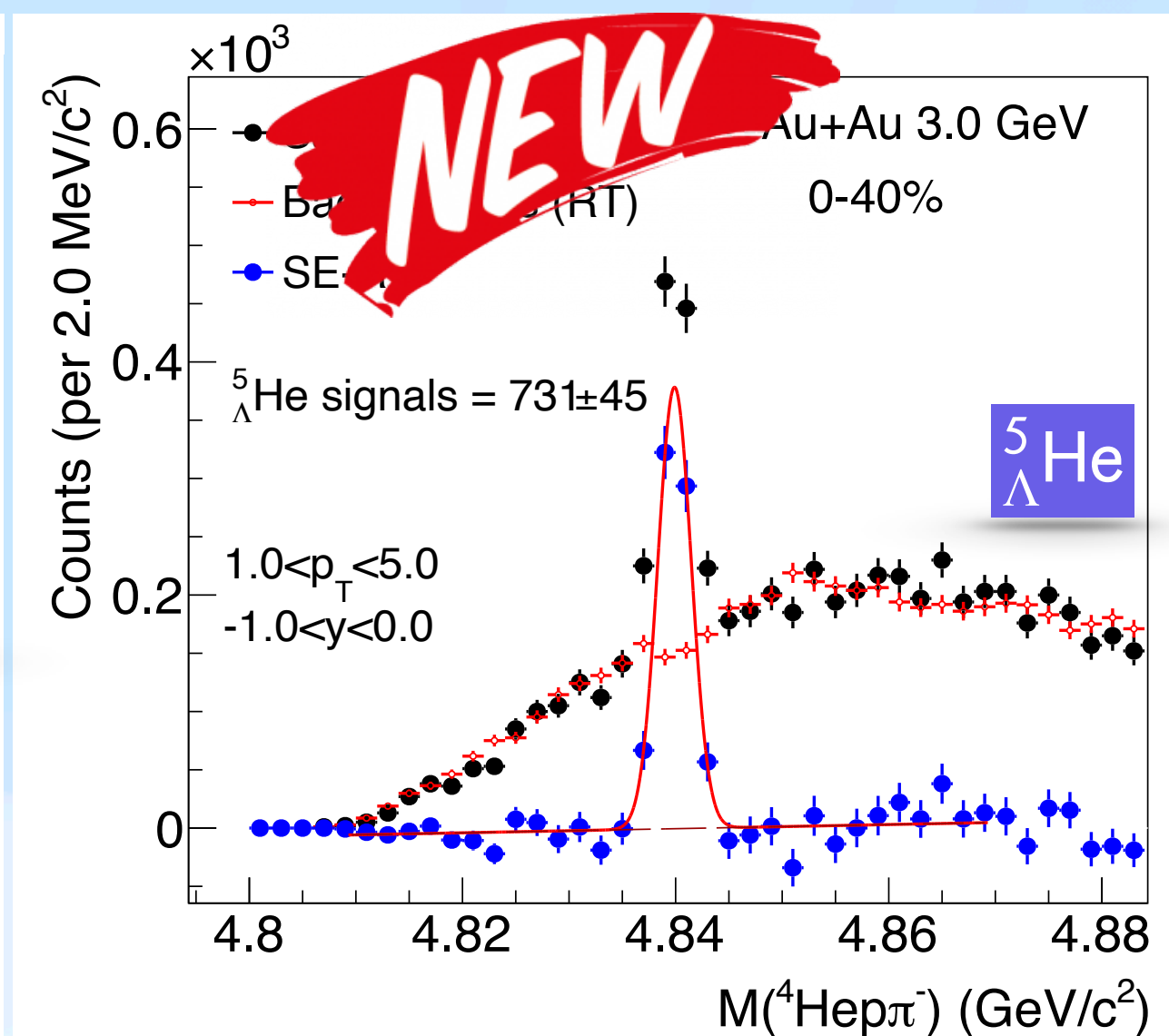
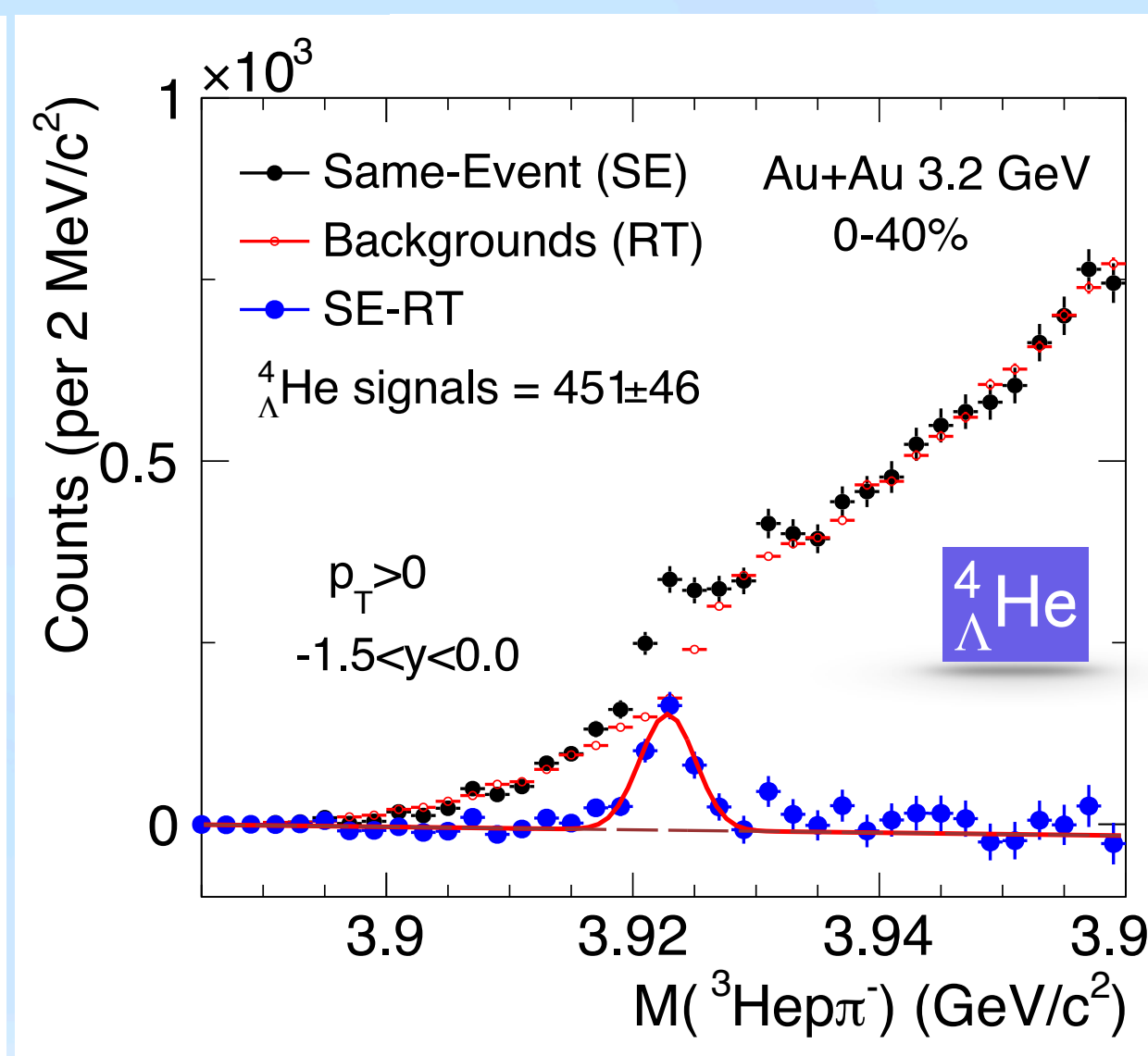
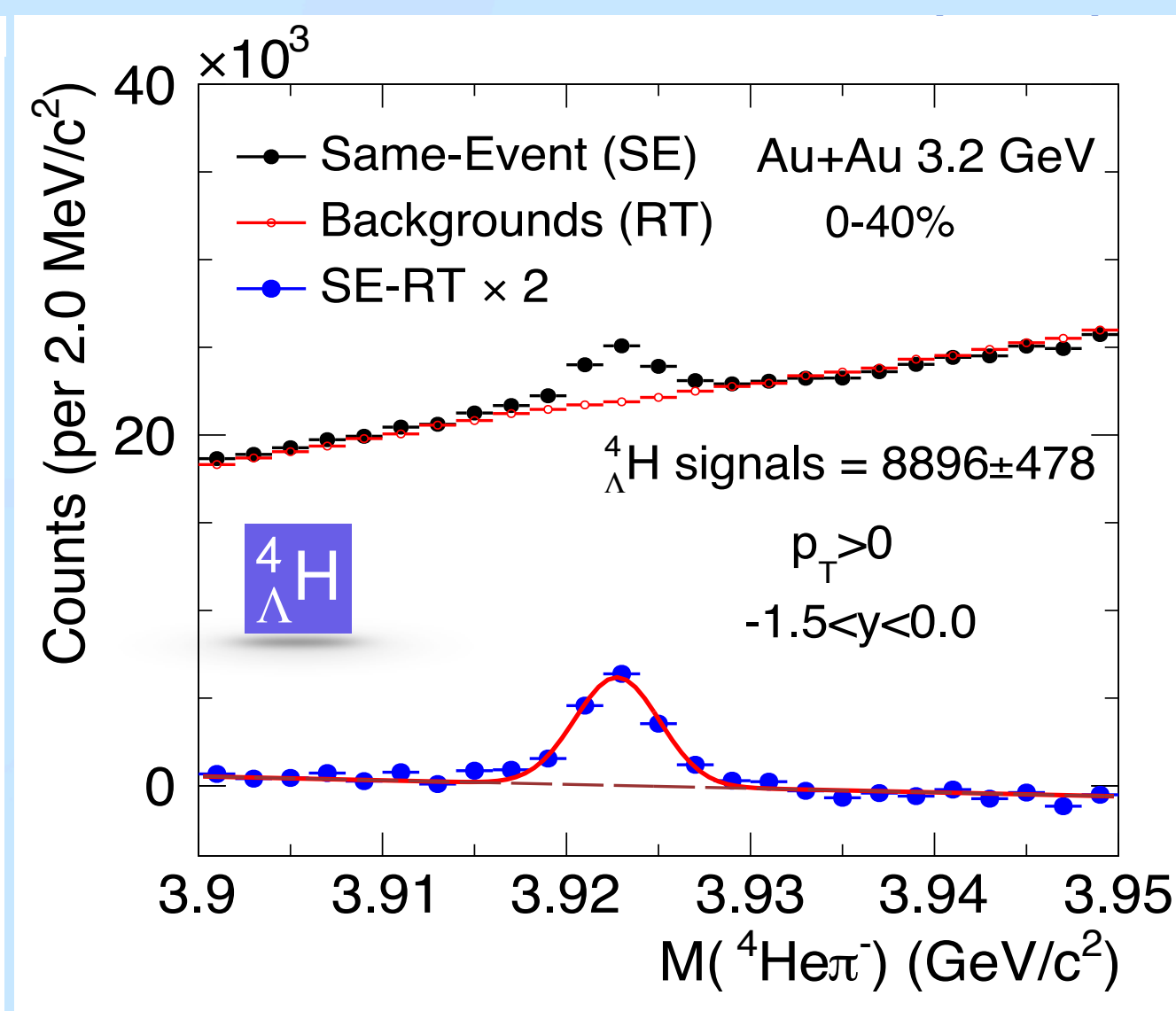
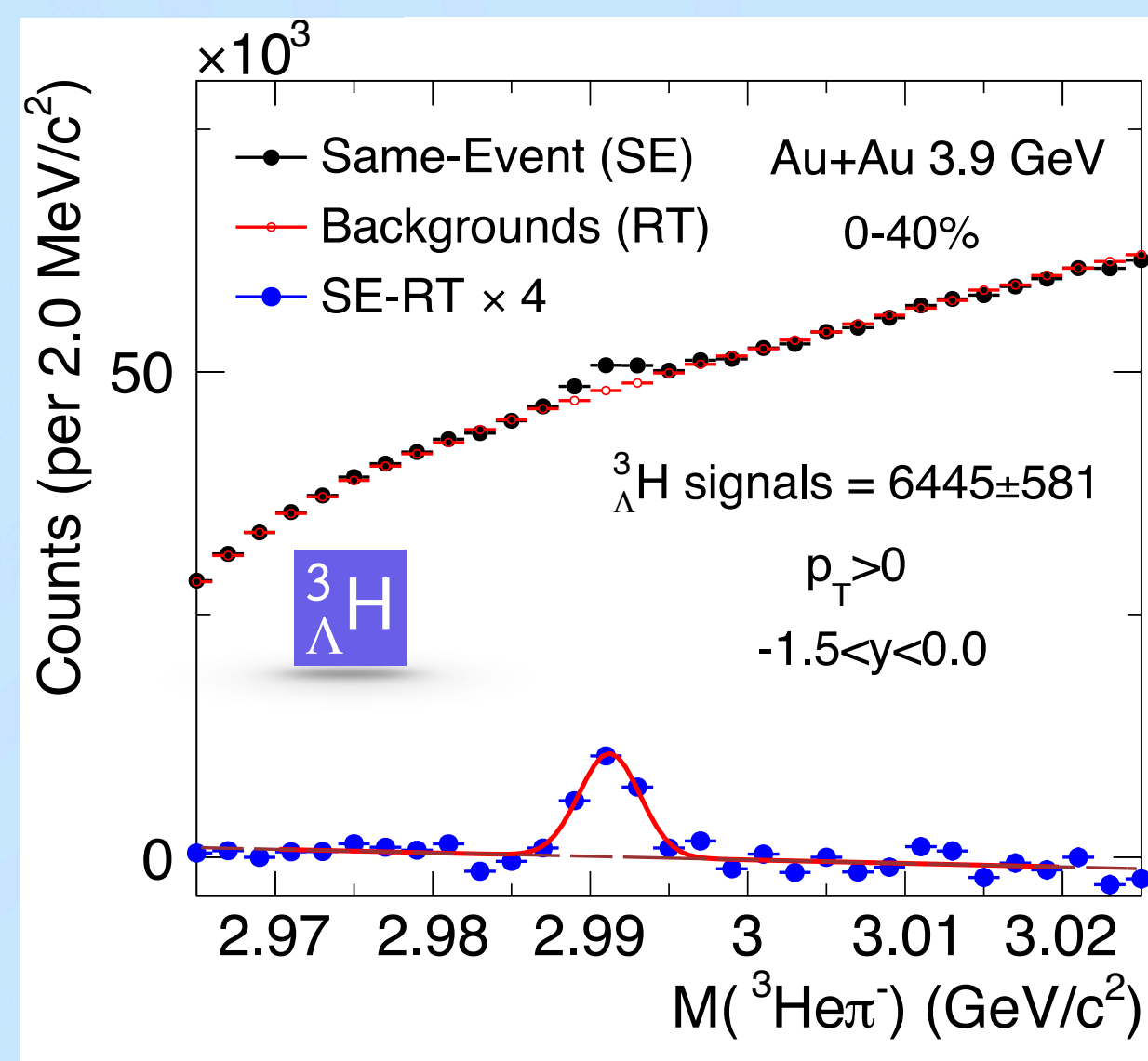
STAR and Beam Energy Scan



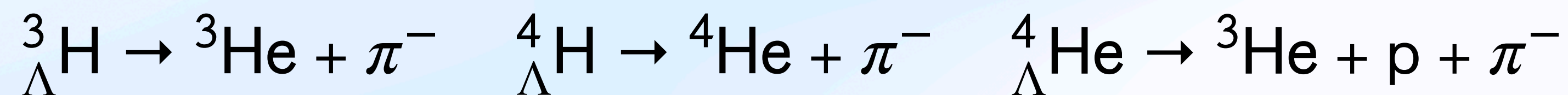
- RHIC BES-II offers great opportunity for hypernuclei measurements



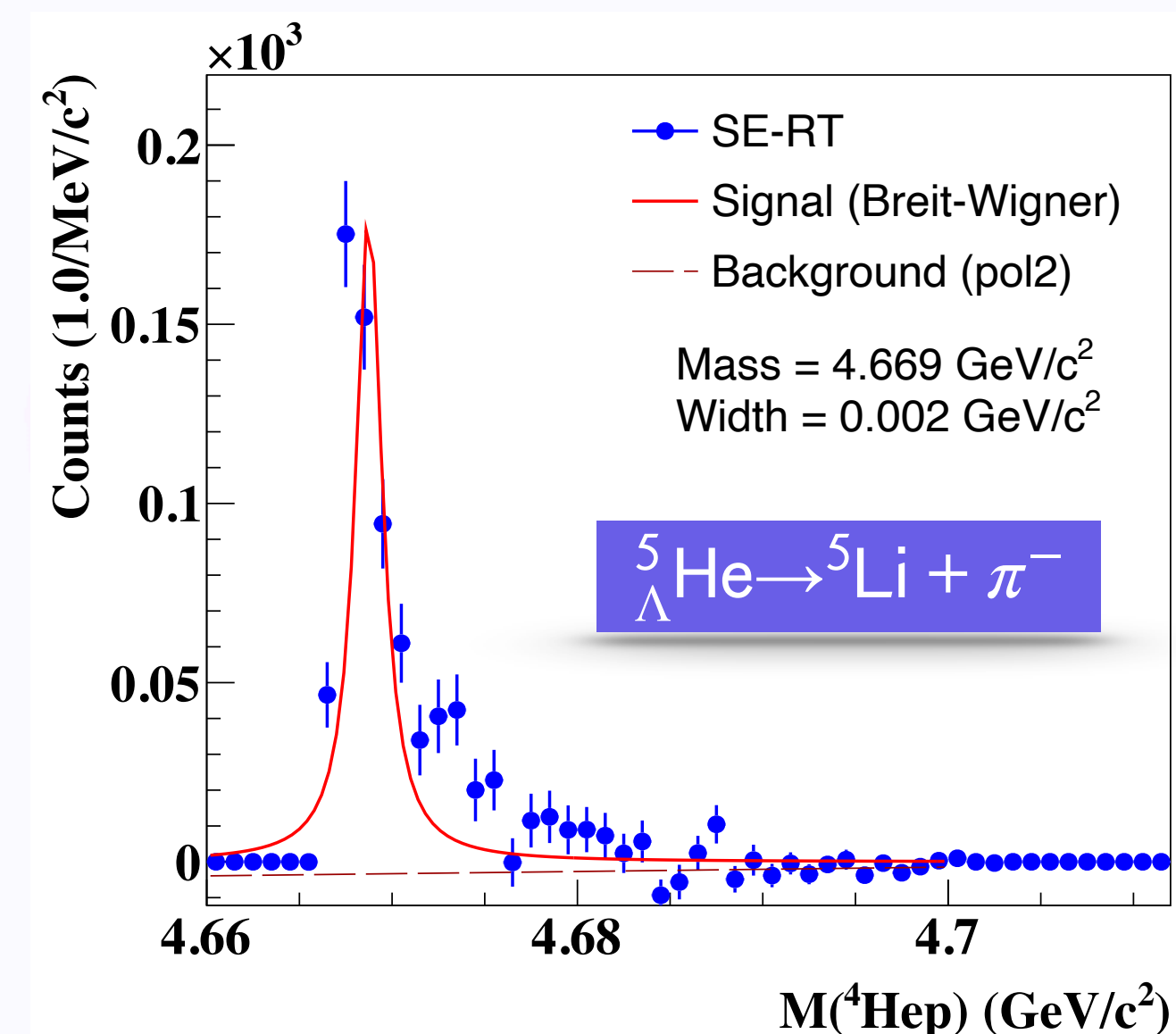
Hypernuclei Reconstruction



- Hypernuclei are reconstructed using the following decay channels:



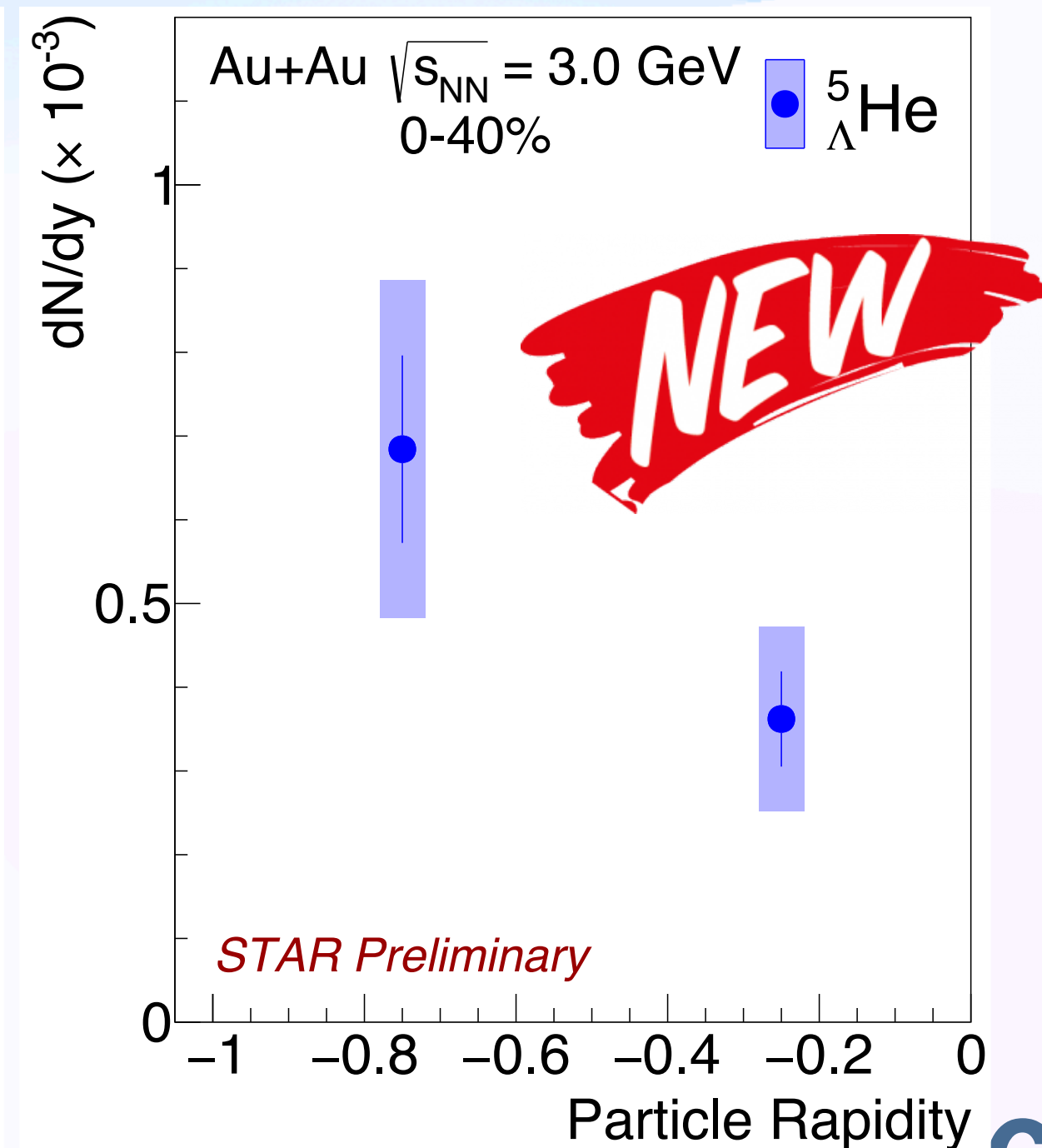
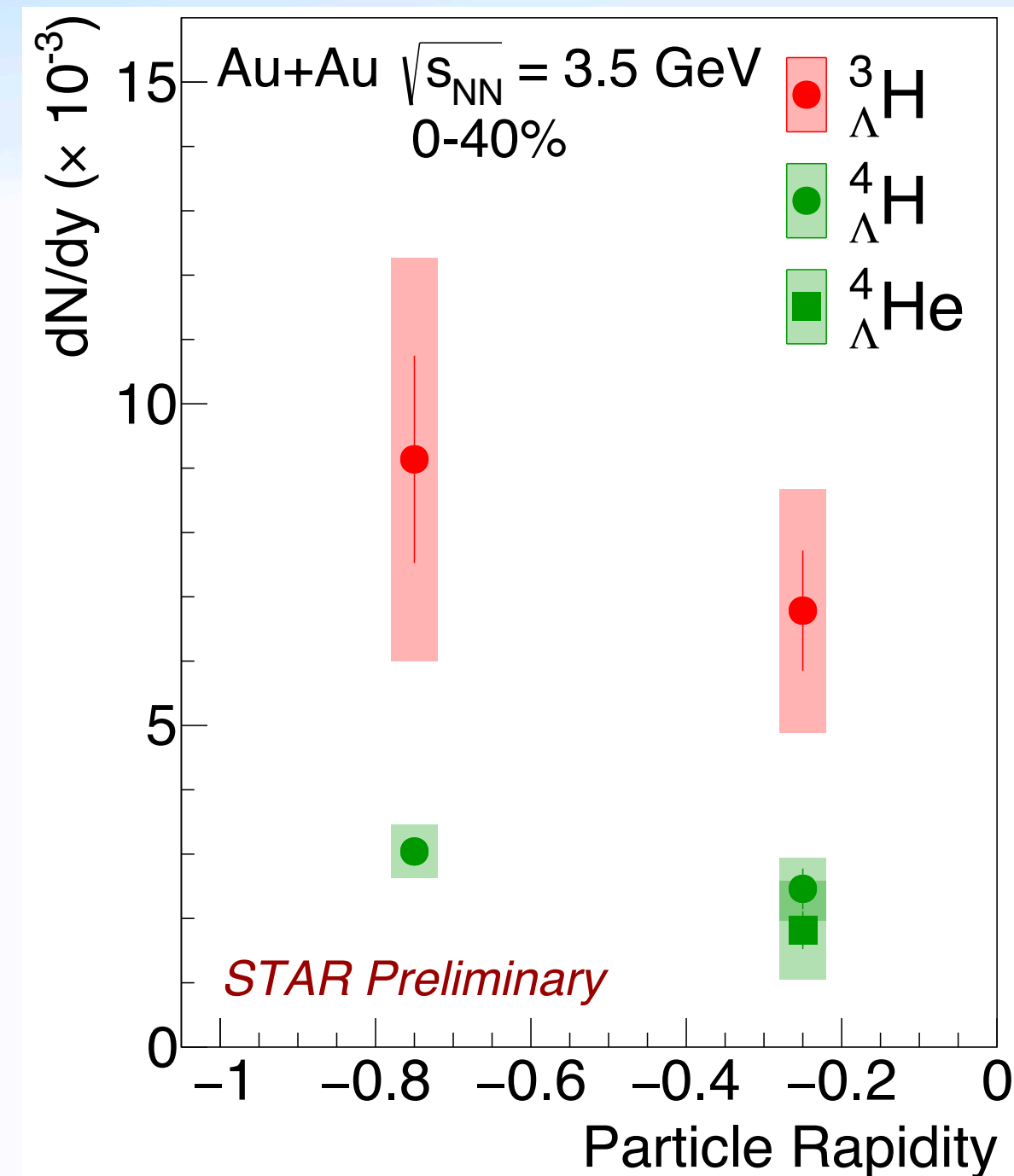
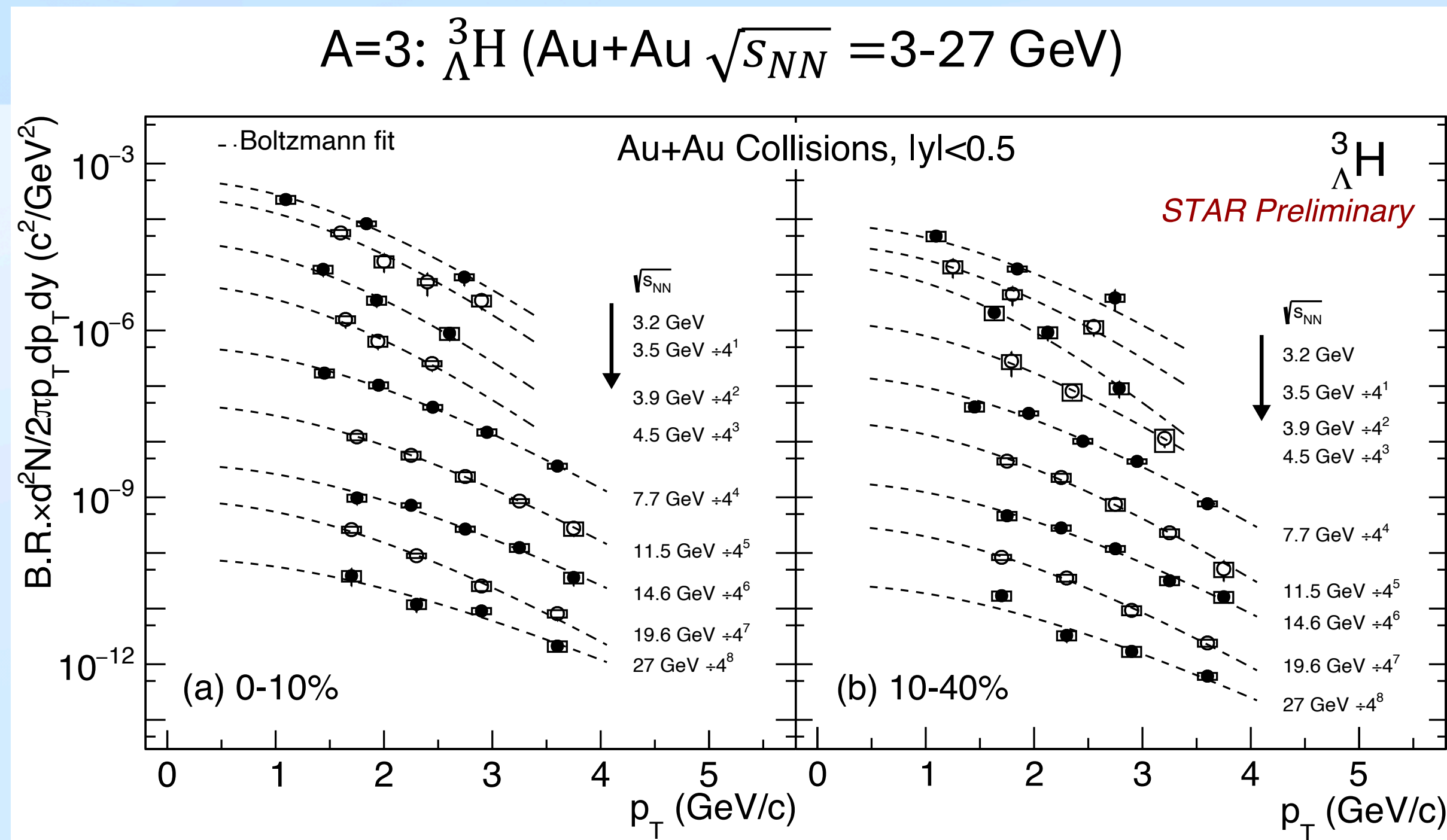
- Combinatorial background estimated via rotating fragments tracks or event mixing
- Efficiency correction using a **data-driven** GEANT simulation
 - To account for the decay kinematics of ${}^4_{\Lambda}\text{He}$, ${}^5_{\Lambda}\text{He}$, the three-body decay phase space is weighted according to a Dalitz plot



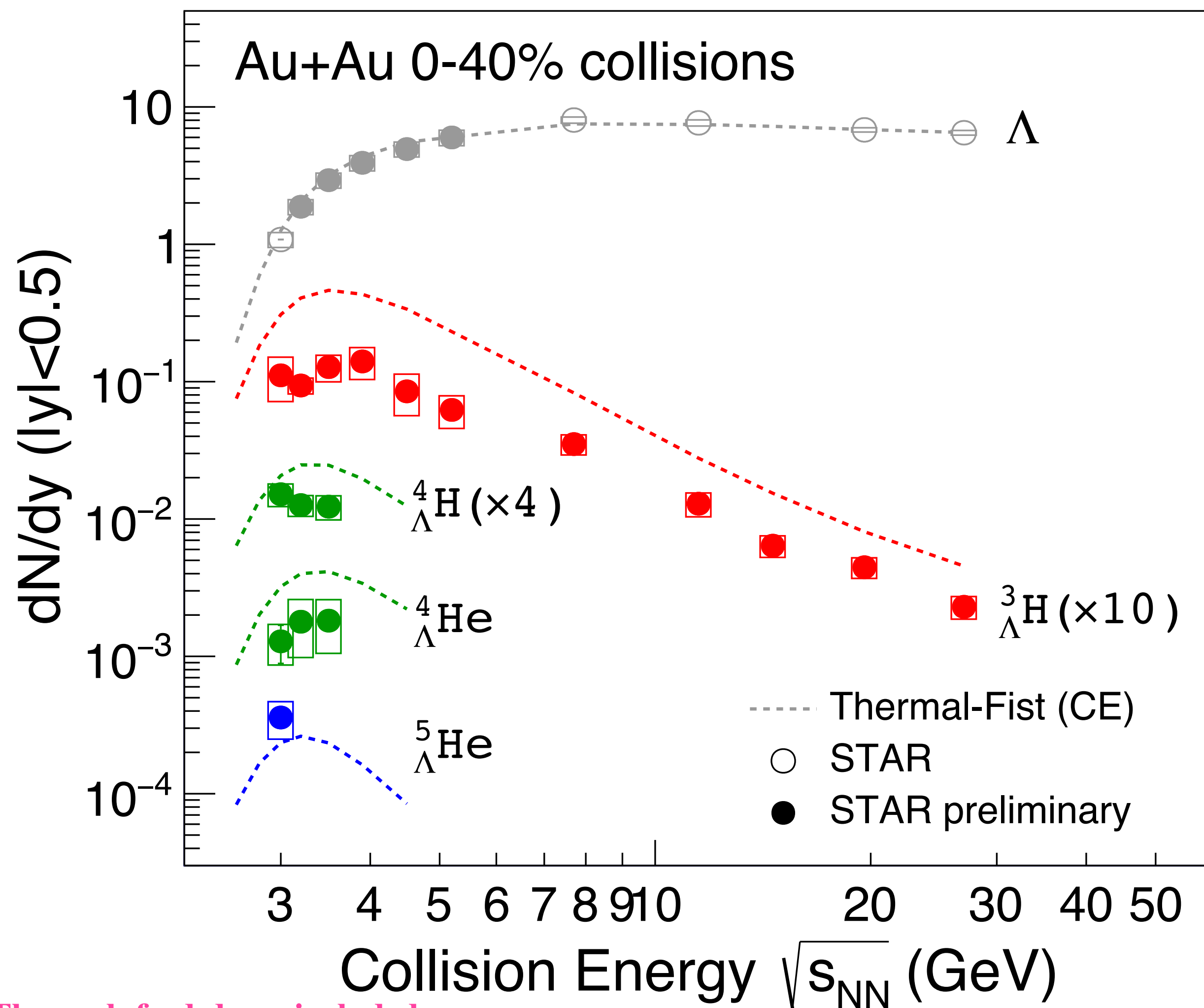
Hypernuclei p_T Spectra and Rapidity

- Measurements cover different energies, different system size!
 - ${}^3_{\Lambda}\text{H}$ in Au+Au collisions at **3-27 GeV**, Au+Au, Zr+Zr, Ru+Ru collisions at **200 GeV**
 - ${}^4_{\Lambda}\text{H}$, ${}^4_{\Lambda}\text{He}$ in Au+Au collisions at **3-3.5 GeV**
 - ${}^5_{\Lambda}\text{He}$ in Au+Au collisions at **3 GeV**
- Production enhanced at target rapidity, more pronounced for heavier nuclei

Spectator matters matter at target rapidity



Excitation Function



Thermal: feed-down included

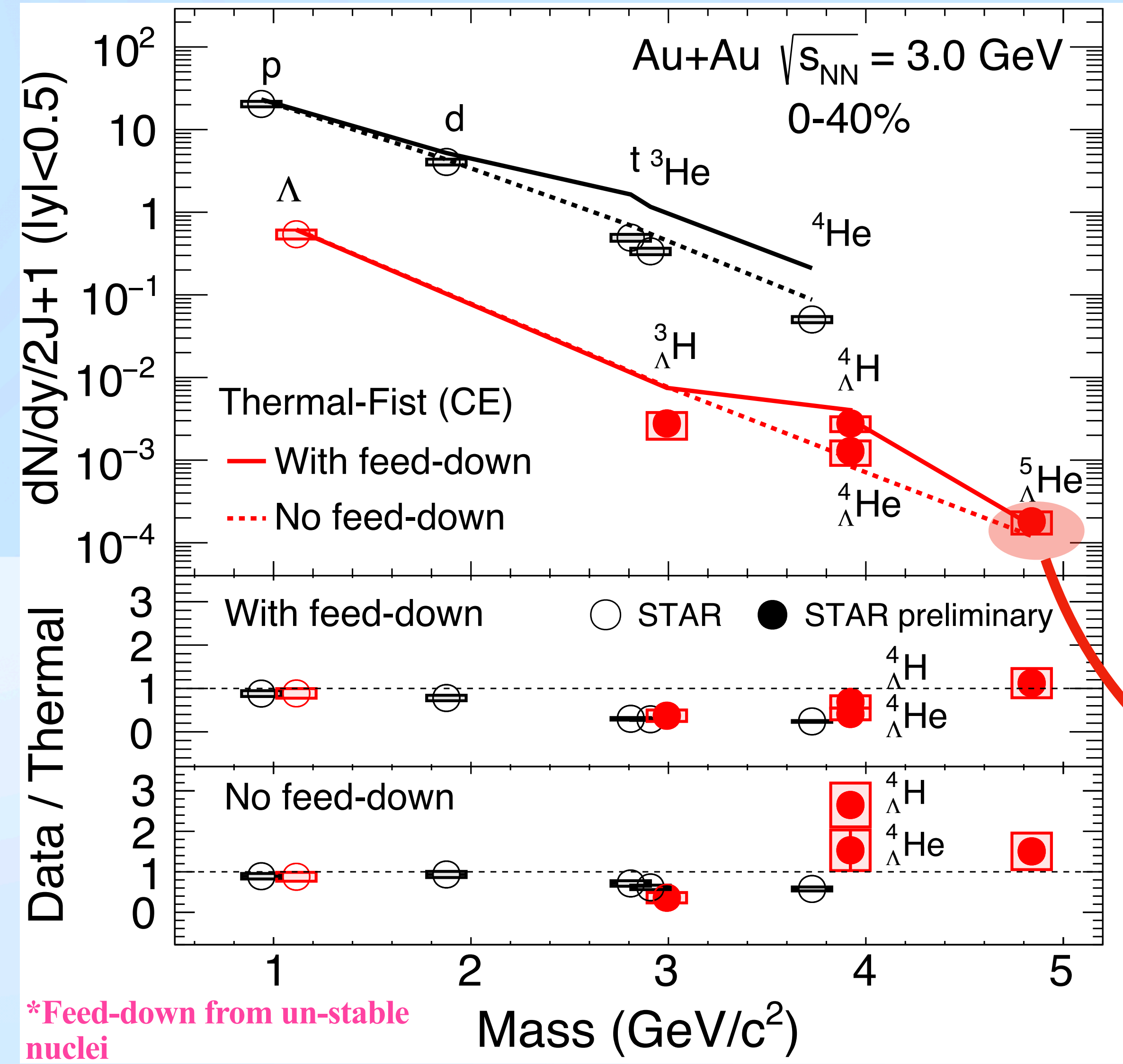
- ${}^3_{\Lambda}H$ show a plateau at $\sqrt{s_{NN}} = 3-4$ GeV
- Similar trend for ${}^4_{\Lambda}H$ and ${}^4_{\Lambda}He$

Interplay between increasing baryon production and stronger strangeness canonical suppression towards low energies

Establishes low energy collision experiments as a promising tool to study exotic strange matter

- Thermal describes Λ , **over-estimate** ${}^3_{\Lambda}H$, ${}^4_{\Lambda}H$, and ${}^4_{\Lambda}He$, slightly **under-estimate** ${}^5_{\Lambda}He$

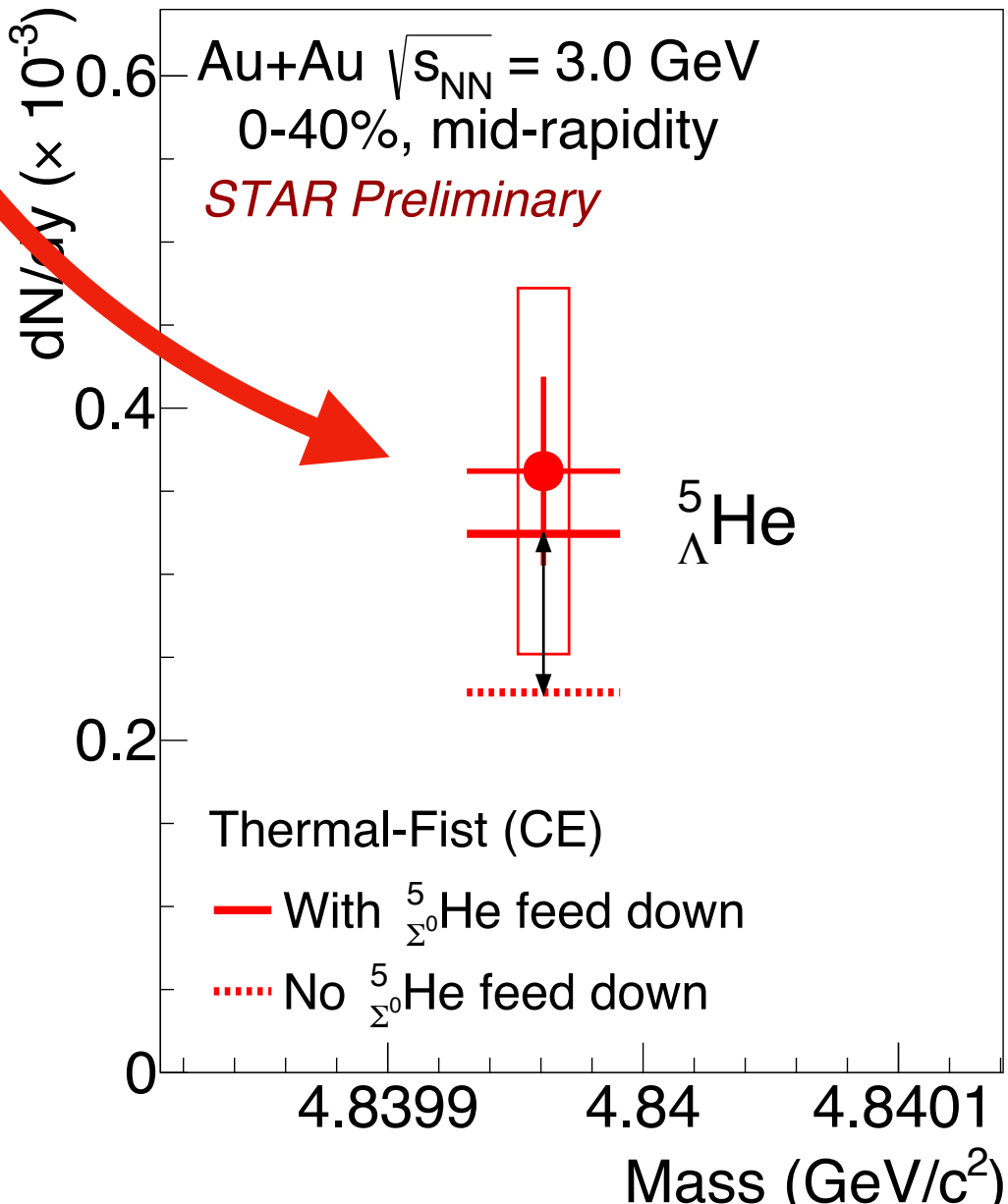
Data to Thermal Model Ratio at 3 GeV



STAR, PRC 110, 054911 (2024)
STAR, PRL 128, 202301 (2022)
STAR, JHEP **2024** (2024)139

- Thermal model predict approxi. exponential dependence of yields/(2J+1) vs A
- Evidence of the formation of ${}^4_\Lambda\text{H}$ and ${}^4_\Lambda\text{He}$ excited states
- Light nuclei overestimated by thermal with feed-down from un-stable nuclei

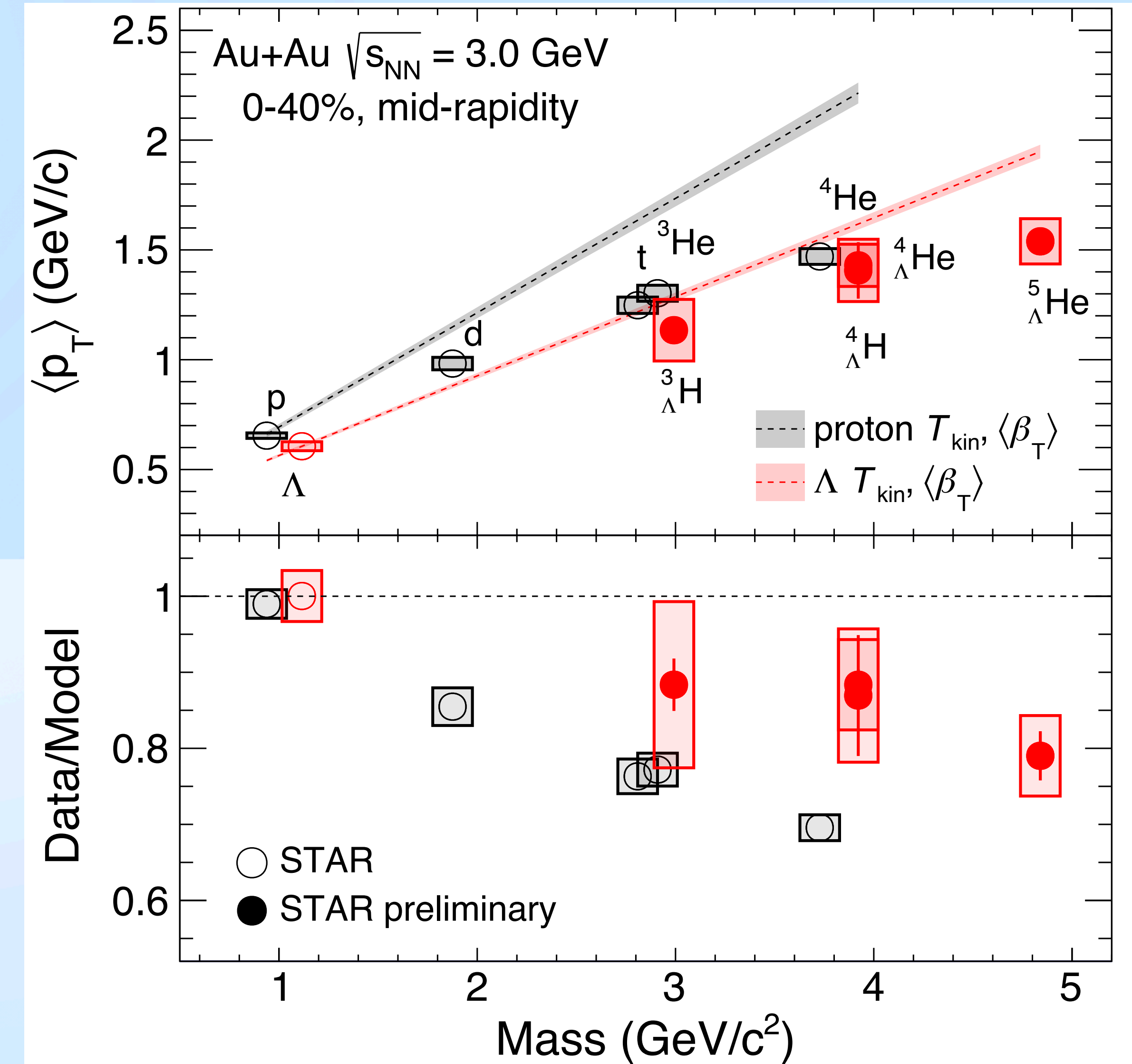
Thermal overestimate feed down fraction?



- Possible feed down from ${}^5_{\Sigma^0}\text{He} \rightarrow {}^5_\Lambda\text{He} + \gamma$

J. Johnstone et al., J. Phys. G 8, L105 (1982)

Mean Transverse Momentum at 3 GeV



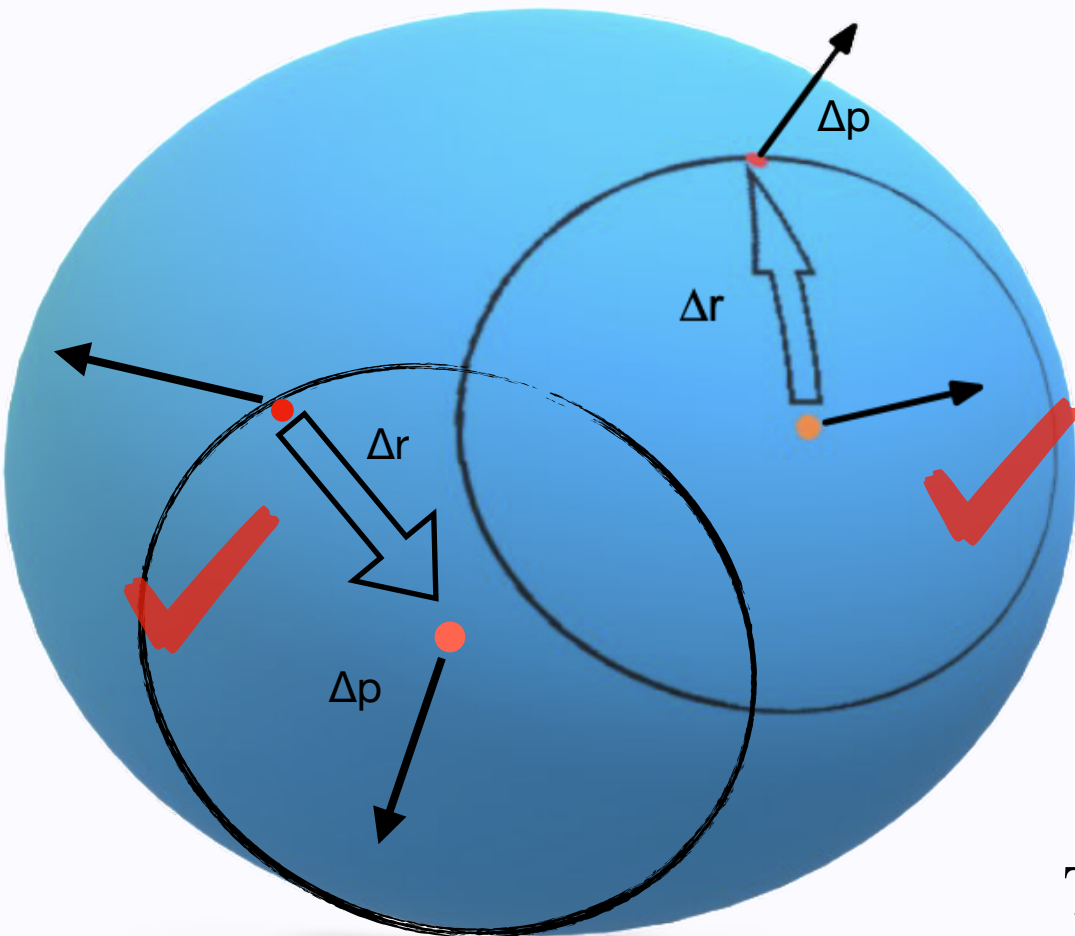
STAR, PRC 110, 054911 (2024)
STAR, PRL 128, 202301 (2022)
STAR, JHEP **2024** (2024)139

Hydrodynamic-inspired Blast-Wave model: assumes particles are emitted thermally from an expanding source with a common flow velocity $\langle \beta_T \rangle$ and kinetic freeze-out temperature T_{kin}

Coalescence scenario: nuclei formed at a later stage after kinetic freeze-out
PLB 794, 50–63 (2019)

- Light nuclei and hypernuclei deviate from the full hydrodynamic picture
 1. **Less correlated nucleons** coalesce to nuclei, leading to **smaller $\langle p_T \rangle$** than perfectly aligned
 2. The heavier nuclei, the larger deviation

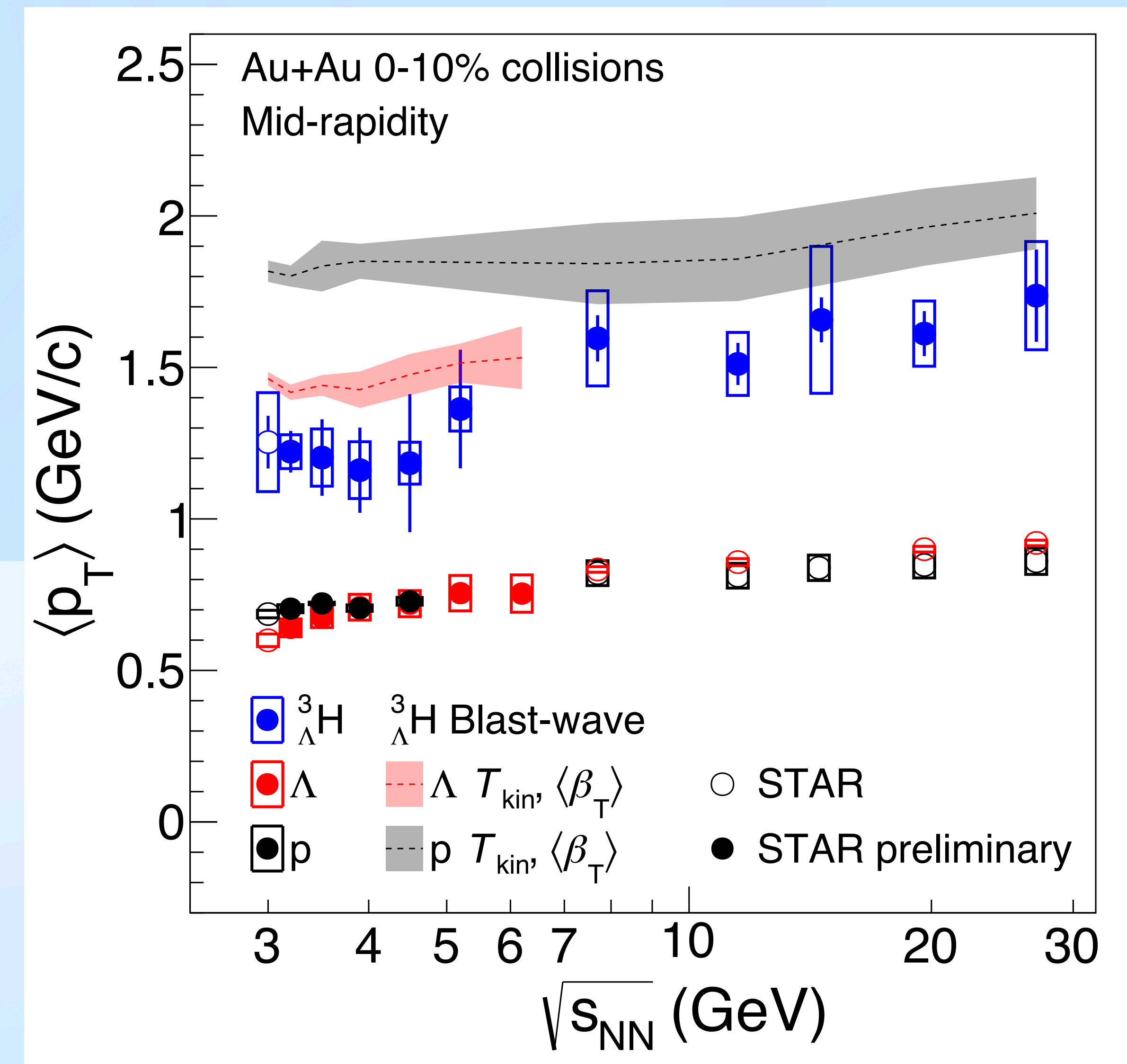
A. I. Sheikh et al., PRC 106, 054907 (2022)



Small effective volume nucleon's r and p are less correlated

T. Reichert, SQM 2022, slides (2022)

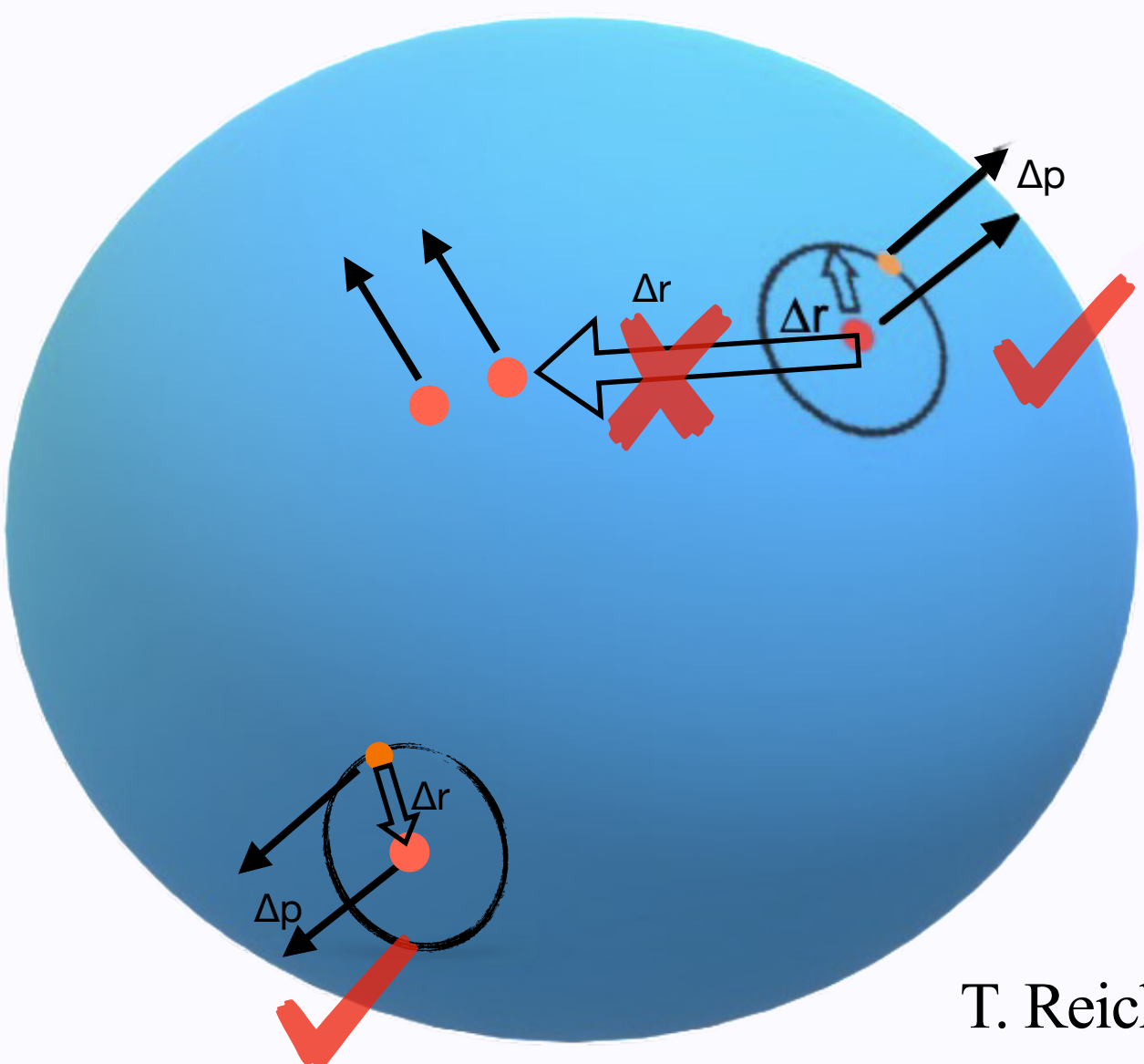
Mean Transverse Momentum v.s. Collision Energy



- At $\sqrt{s_{NN}} \geq 7.7$ GeV, ${}^3_{\Lambda}\text{H}$ $\langle p_T \rangle$ tends to approaching that of protons
- Likely due to increasing effective volume $\leftrightarrow$ decreasing coalescence parameter B_A , where nucleons that eventually coalesce are more likely to be aligned in space and momentum

See poster by Yixuan Jin (xx/xx)

Suggests a change in the effective volume across energies



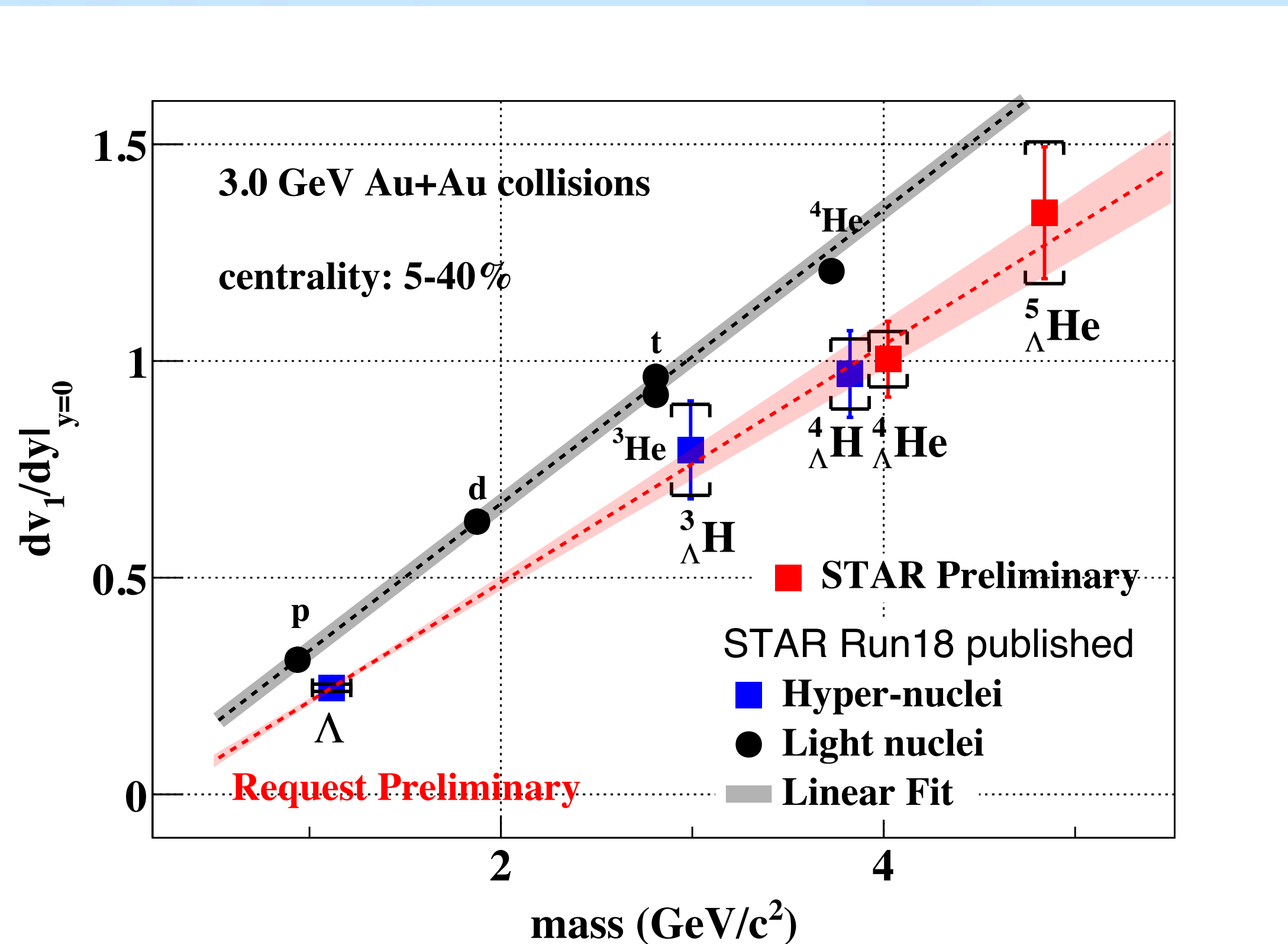
Large effective volume nucleon's r and p are correlated

STAR, PRC 110, 054911 (2024)
STAR, PRL 128, 202301 (2022)
STAR, JHEP **2024** (2024)139

T. Reichert, SQM 2022, slides (2022)

Collective Flow at 3 GeV

See poster by Junyi Han (xx/xx)



- Directed flow of hypernuclei follows mass scaling
- Qualitatively consistent with coalescence formation of hypernuclei

Summary and Outlook

- Hypernuclei measurement from STAR BES-II
 1. **First measurement of $A = 5$ hypernuclei yield and direct flow in Au+Au collisions at $\sqrt{s_{NN}} = 3$ GeV**
 2. Mid-rapidity yields overestimated by thermal model
 3. Mean transverse momentum tends to lower than p (Λ) hydrodynamic-inspired blast-wave model at $\sqrt{s_{NN}} < 7.7$ GeV
Consistent with coalescence picture: weaker space and momentum correlation among coalescing nucleons in a smaller effective volume at low energy
 4. Collective flow qualitatively consistent with coalescence model

Outlook:

- High statistics 3 GeV FXT and RHIC top energy data: more precise measurement, and search for heavier hypernuclei ($A > 5$), double- Λ hypernuclei
 - ➡ Further constrain production mechanism, and YN, YY, YNN interactions