

Decomposition of the Gamma correlator in 200 GeV Au+Au collisions (run11)

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Outline

- Introduction to CME
- Method
- Result
- Future Work

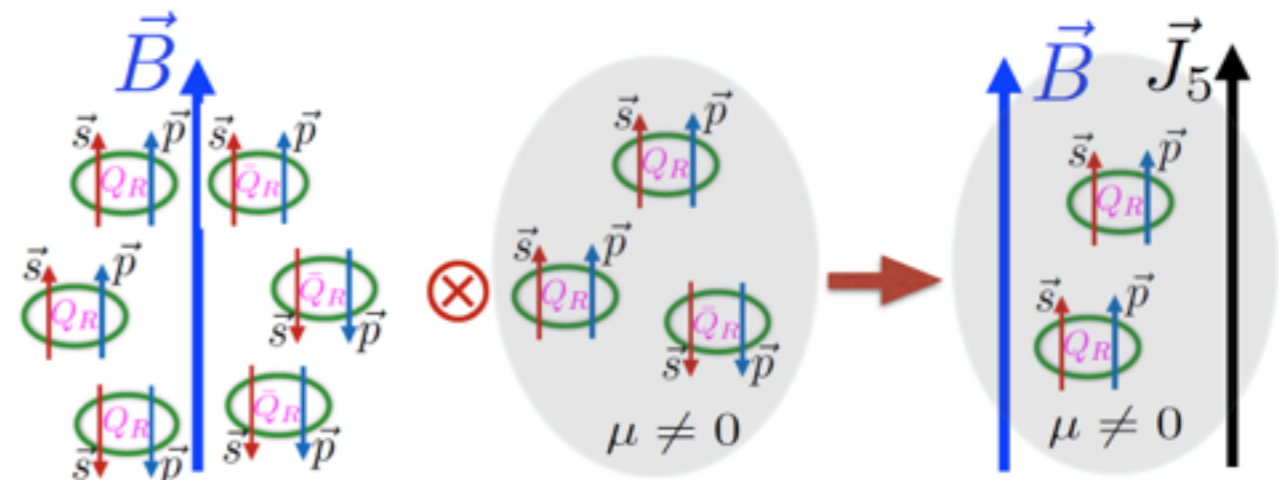
Introduction to CME

- CME physics: In the magnetic field the current will be generated due to the unbalance of chiral quarks(right-handed or left-handed)
- Gamma and delta definition:

$$\gamma \equiv \langle \cos(\phi_1 + \phi_2 - 2\Psi_{RP}) \rangle = \langle \cos^* \cos - \sin^* \sin \rangle$$

$$\delta \equiv \langle \cos(\phi_1 - \phi_2) \rangle = \langle \cos^* \cos + \sin^* \sin \rangle$$

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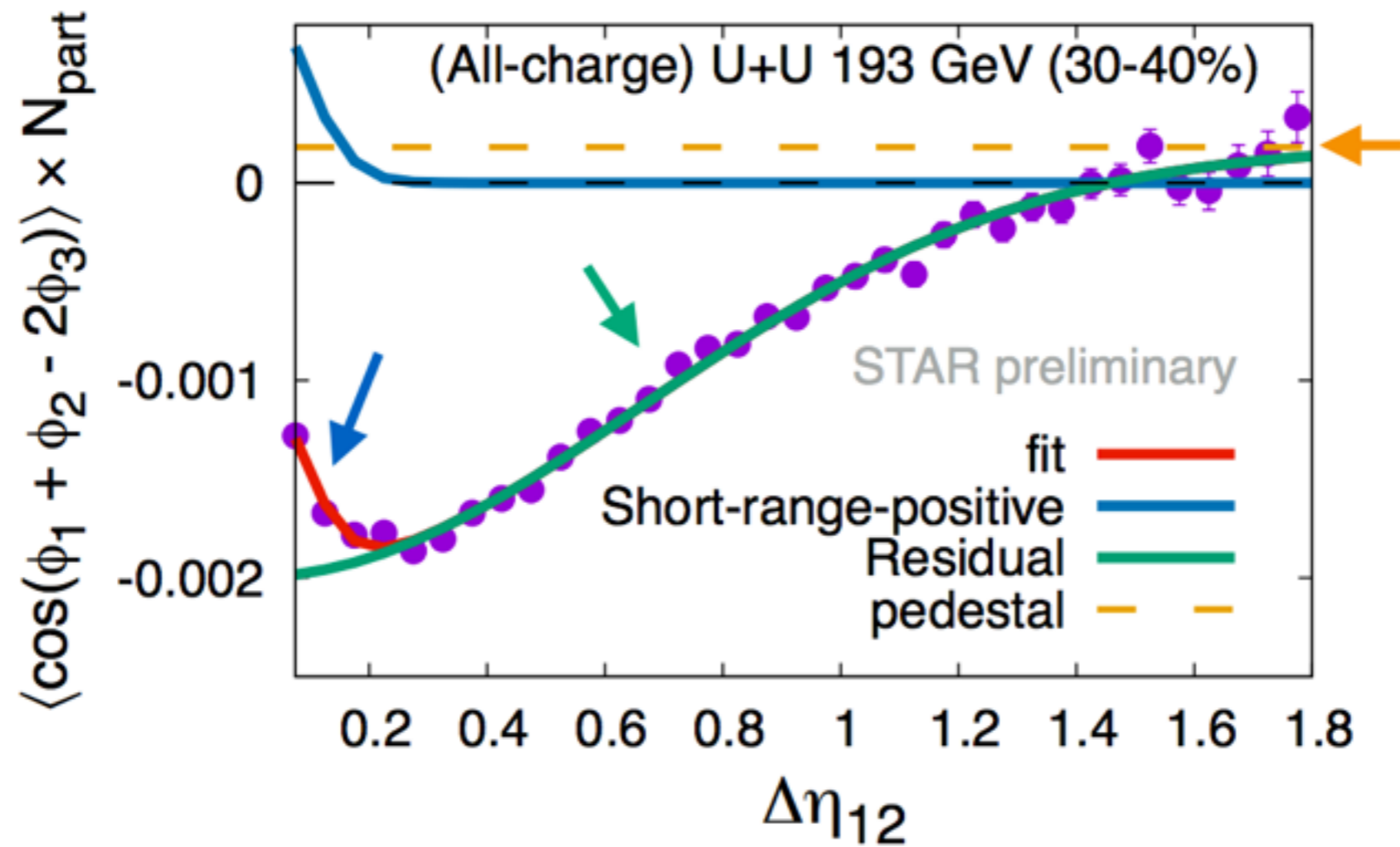


Background Study

- Background: Although there are many contributions to the background, for now, we focus on the short-range correlations
- One method to study and reduce such effect is to fit the data with gaussians and then minus these narrow ones.

$$C_{112}(\Delta\eta_{12}) = \underbrace{A_{SR}^+ e^{-(\Delta\eta)^2/2\sigma_{SR}^2}}_{\text{Short-range-positive}} - \underbrace{A_{IR}^- e^{-(\Delta\eta)^2/2\sigma_{IR}^2} + A_{LR}}_{\text{Residual}} \rightarrow \text{Pedestal}$$

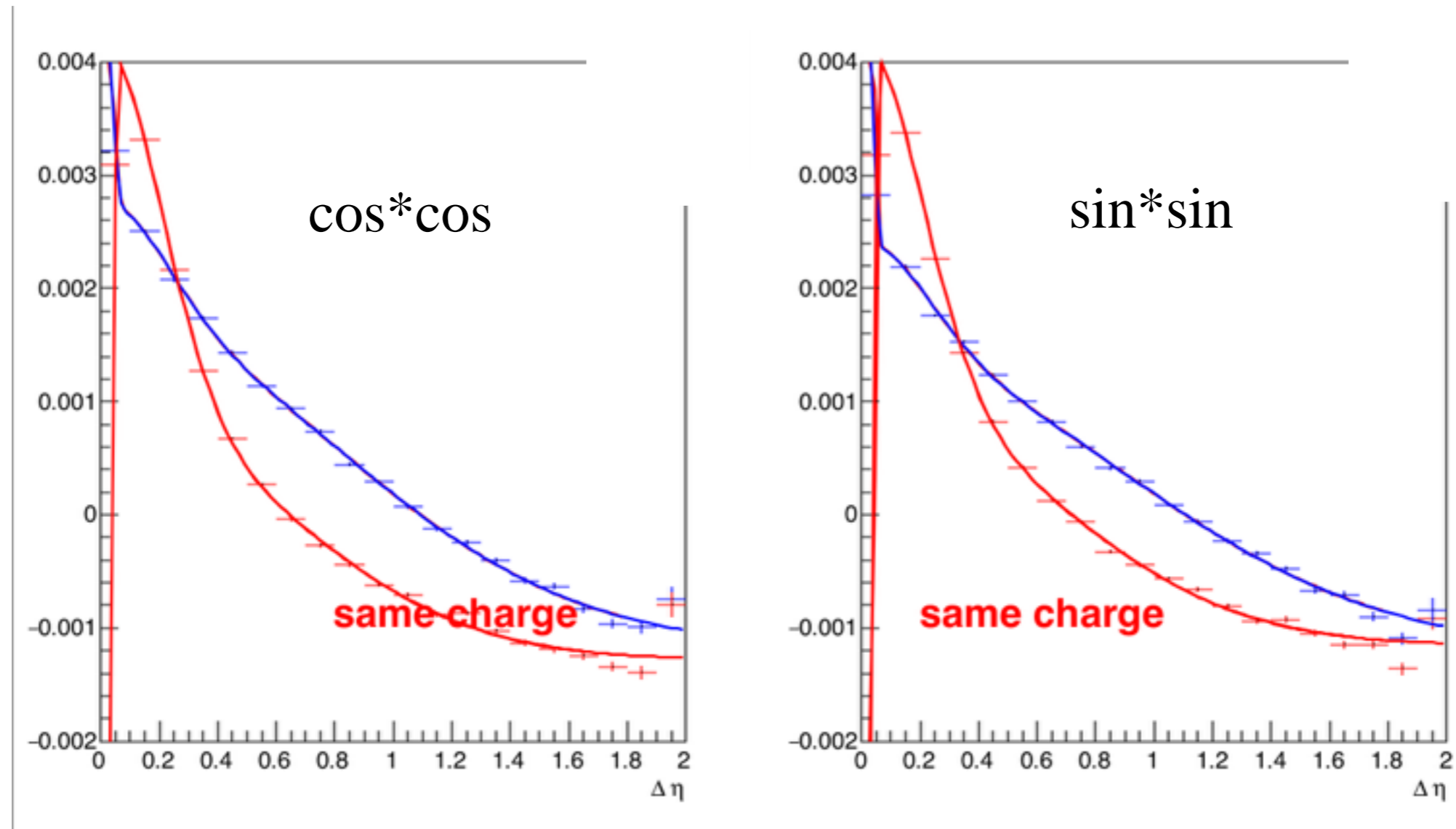
- Previous study from STAR(Prithwish)



Modified method applied to reduce background

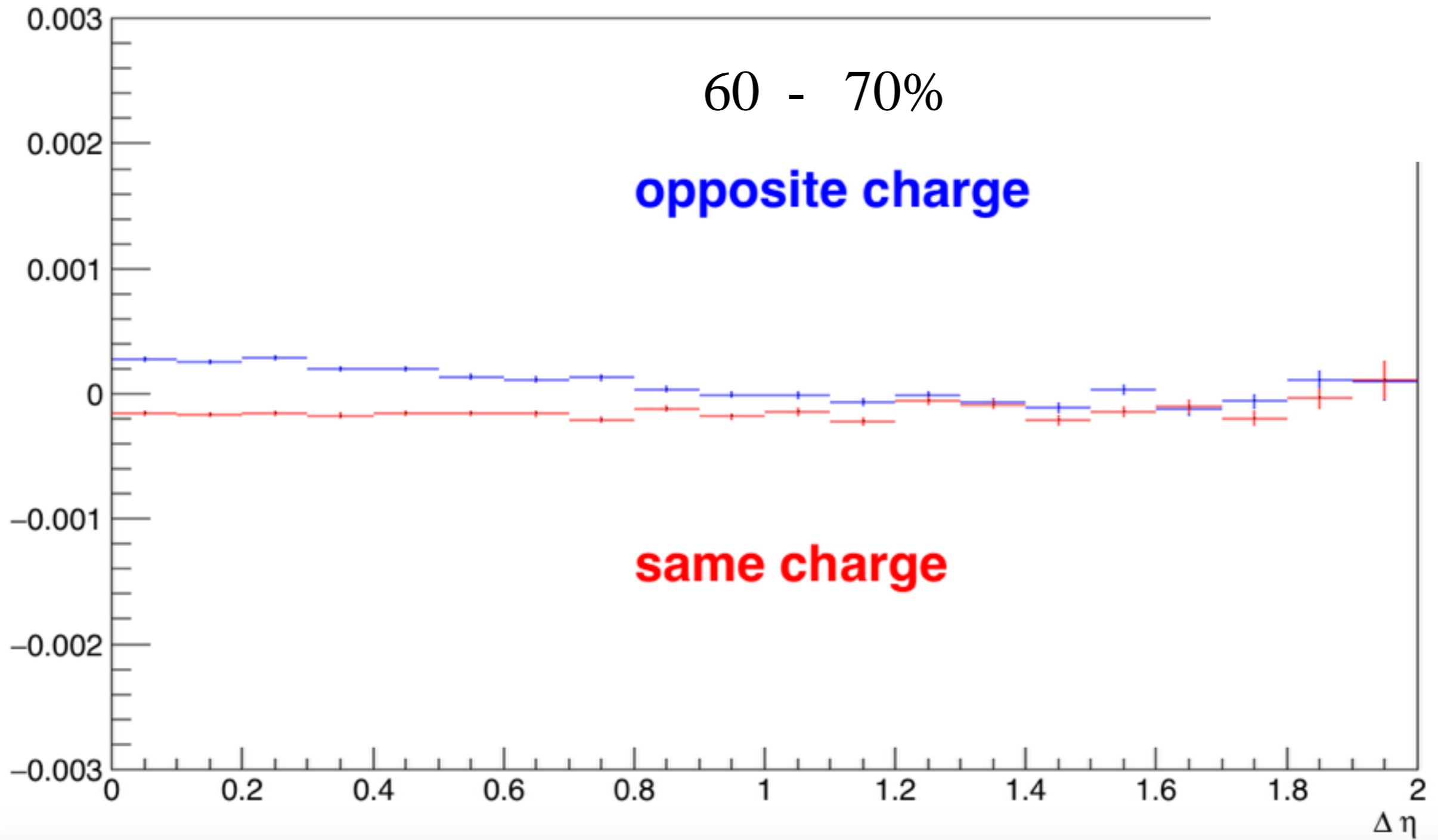
- We fit the data with multiple gaussian functions and then remove the narrow gaussians to eliminate the short-range contribution in the data.
- What's new: Instead of fitting the original signal, I try to fit the data sets by parts, in-plane and out-of-plane separately OS_cosc, SS_cosc, OS_sins, SS_sins then rebuild the signal. All fittings use three gaussians in order to make sure even the very short range effect can be described
- Advantage: Much more smooth fit for data points and clear trend for the overall results.

- One fit example for 60%-70% collision

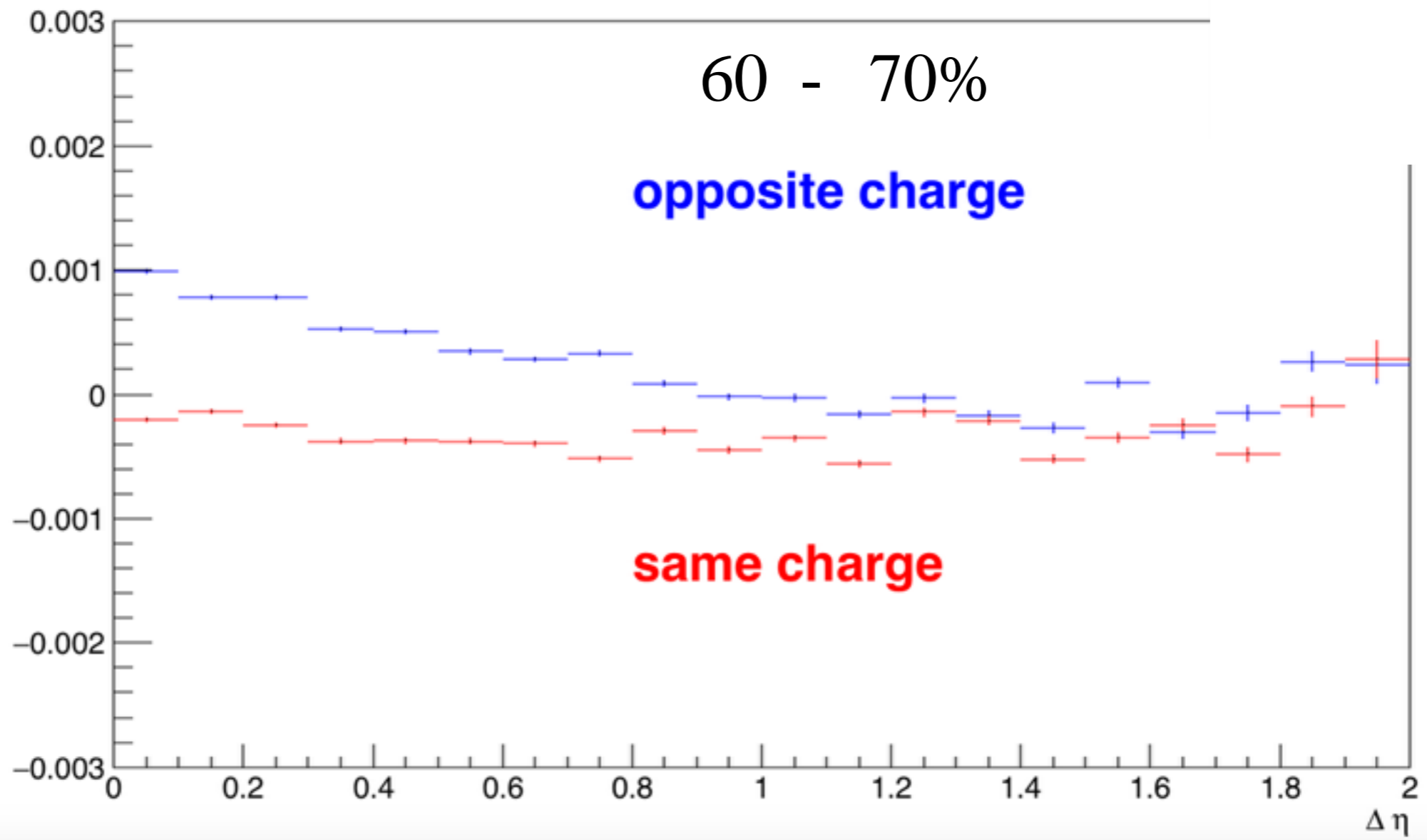


- It is clear to see the short-range contribution on this fit, and since we use three gaussians, the very short-range effect can also be described.

γ with short-range correlations removed

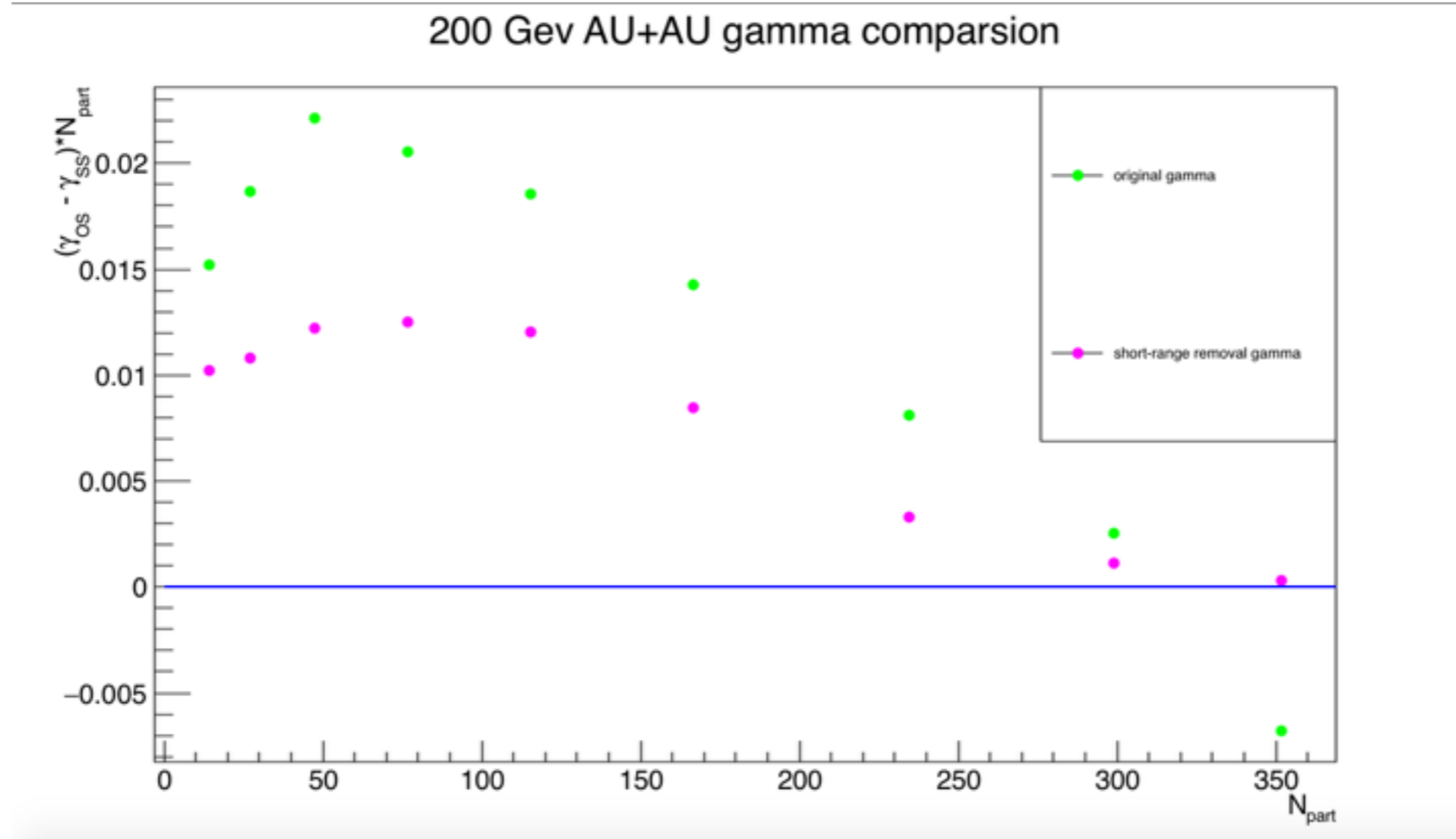


δ with short-range correlations removed



Results

Gamma trend vs centrality
Original vs Short range removal

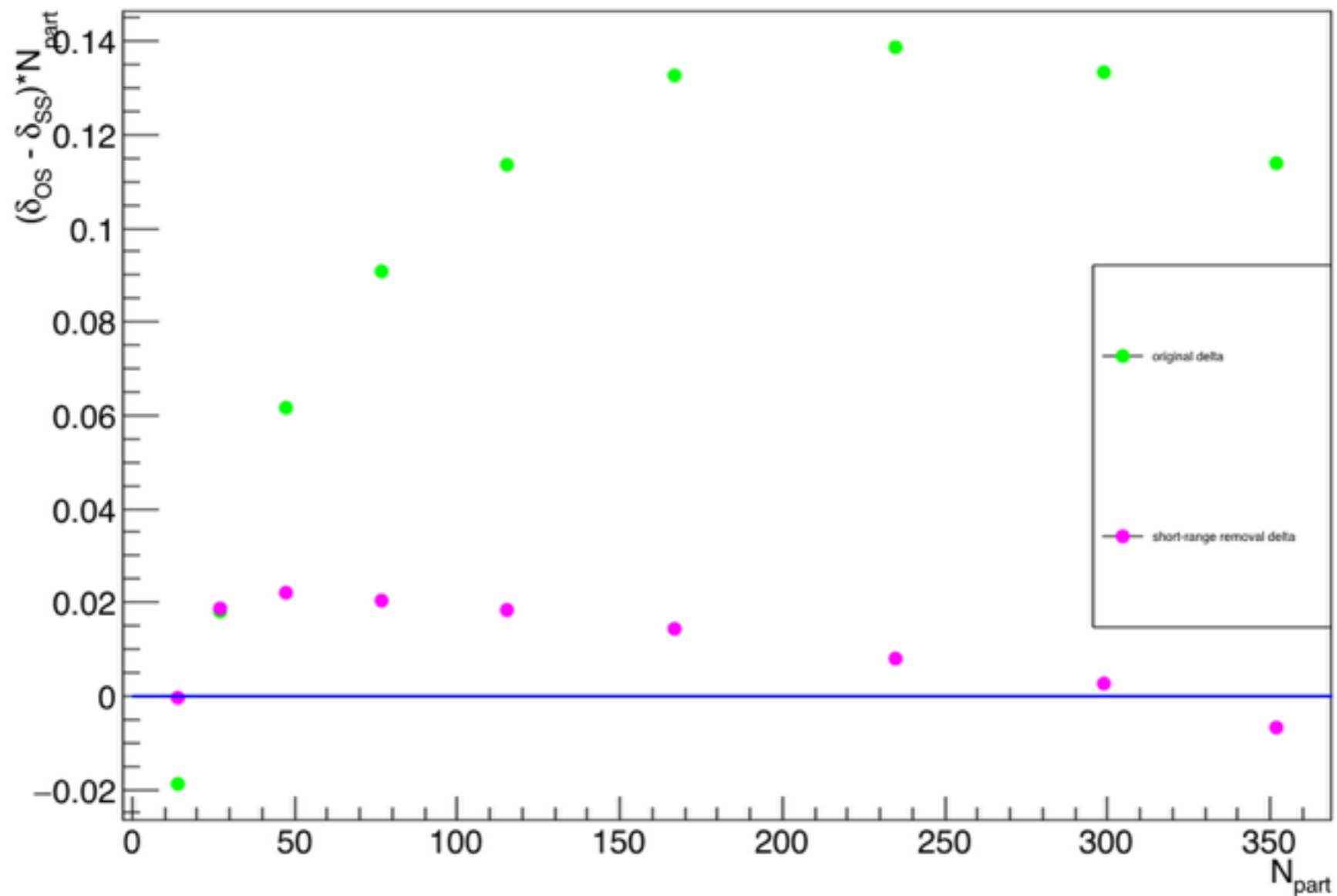


For the most central bin, after short-range removal, the result is more reasonable

Results

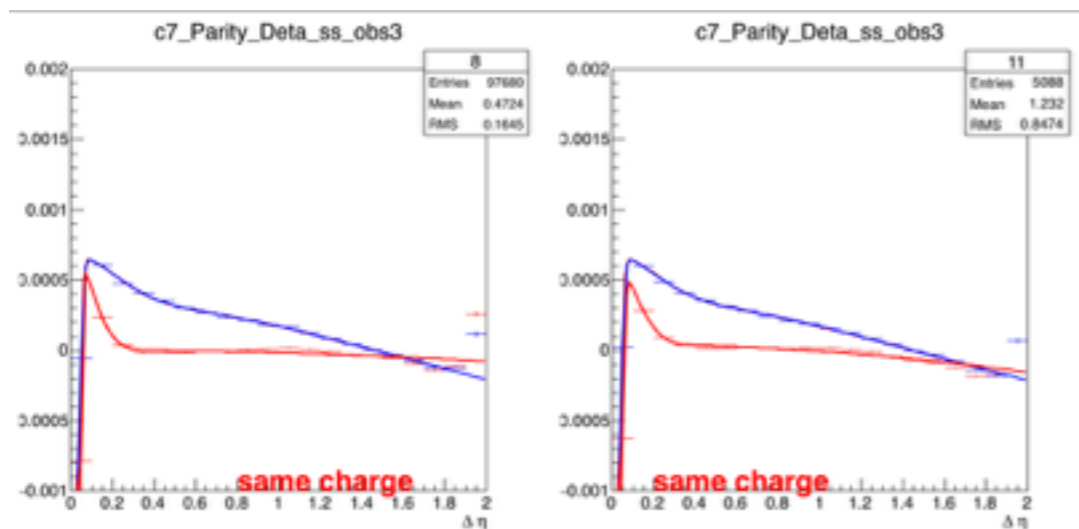
Delta trend vs centrality
Original vs Short range removal

200 Gev AU+AU delta comparison



Systematics

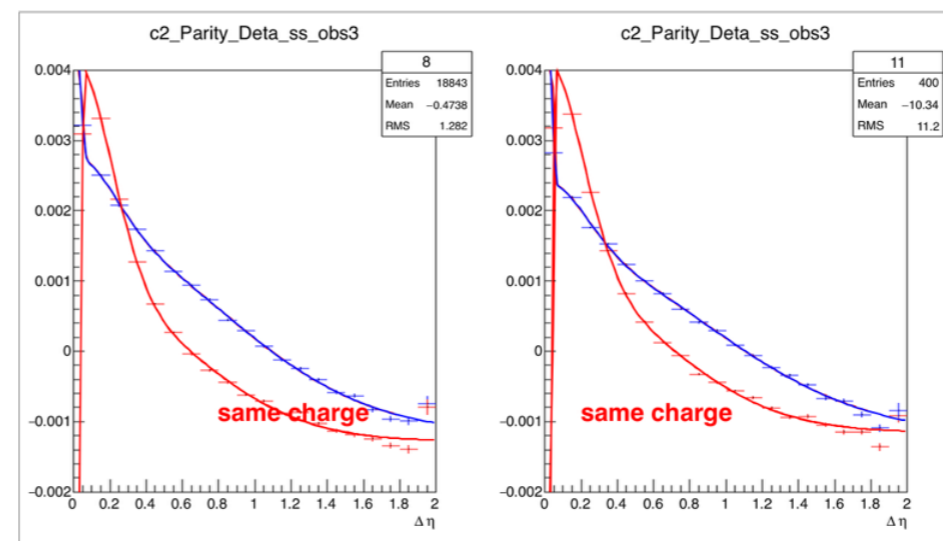
- Two methods are used to check whether the fit results are reasonable: χ^2/ndf and the peak width vs centrality
- Numerically, χ^2/ndf is close to 1 for most cases, but they are also sensitive one or two bad points.



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the cc_os chi^2 is 15.4882
the cc_ss chi^2 is 69.5363
the ss_os chi^2 is 11.005
the ss_ss chi^2 is 53.6317
    
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cen7 fit result and its χ^2/ndf



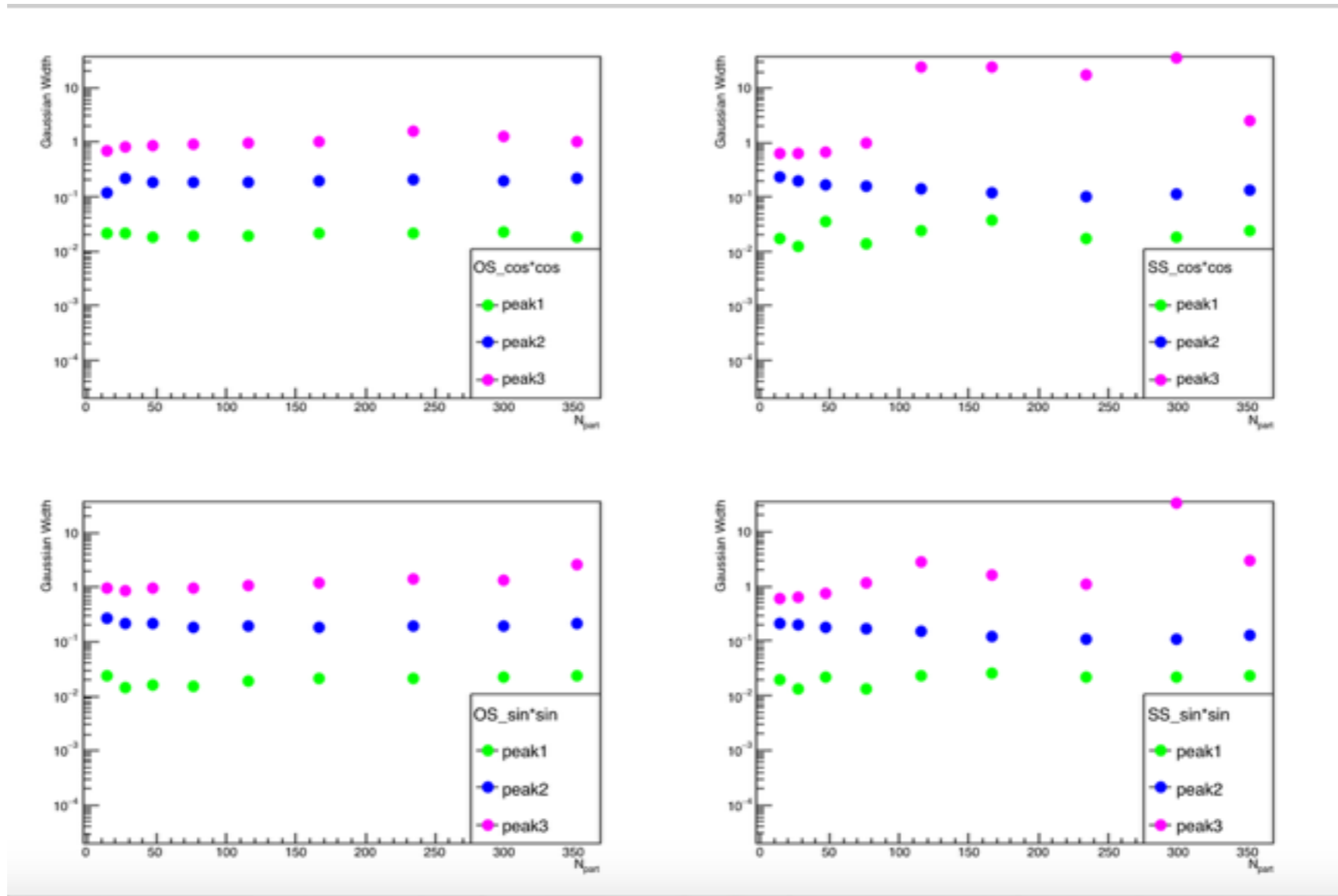
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the cc_os chi^2 is 2.62406
the cc_ss chi^2 is 4.95332
the ss_os chi^2 is 4.30119
the ss_ss chi^2 is 4.86062
    
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cen2 fit result and its χ^2/ndf

Results

Peak width vs centrality



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The three peaks are well separated

Future Work

- For CME physics, better background model is needed to describe the data we have
- For this method, we can apply it to the AMPT model or $p+Au(d+Au)$ data sets to future check whether it can work well.
- I am going to submit an abstract on this topic to DNP REU poster session this year.