# Coherent diffractive photoproduction of $\rho^0$ mesons on gold nuclei at RHIC

STAR Collaboration

# 4 Abstract

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The STAR Collaboration reports on the photoproduction of  $\pi^+\pi^-$  pairs in gold-gold collisions at a center of mass energy of 200 GeV/nucleon. These pairs 6 are produced when a nearly-real photon emitted by one ion scatters from the other. The differential cross section  $d\sigma/dt$  clearly exhibits a diffraction pattern, 8 compatible with scattering from a gold nucleus, with 2 dips visible. We fit the  $\pi^+\pi^-$  mass spectrum to a combination of  $\rho^0$  and  $\omega$  resonances, plus a direct 10  $\pi^+\pi^-$  continuum; the ratio of  $\rho^0$ : direct  $\pi^+\pi^-$  is consistent with previous 11 measurements in lighter systems. The  $\omega$  component is comparable with that 12 expected from the measured  $\omega$  photoproduction cross section and  $\omega \to \pi^+ \pi^-$ 13 branching ratio. 14

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#### 17 **1. Introduction**

Relativistic heavy ions are accompanied by high fluxes of nearly-real pho-18 tons, due to their large electric charge and the strongly Lorentz contracted 19 electric fields. In relativistic heavy ion collisions, these fields can produce pho-20 tonuclear interactions. When they collide and interact hadronically, the strong 21 interactions obscure these electromagnetic interactions. However, when they 22 physically miss each other, the photonuclear interactions can be visible; these 23 are referred to as Ultra-Peripheral Collisions (UPCs). The photon flux is well 24 described within the Weizsäcker-Williams formalism [1, 2]. These exchanged 25 photons are nearly real, with virtuality  $(\langle Q^2 \rangle \sim 2 \times 10^{-3} \text{GeV}^2)$ . 26

For photoproduction of  $\rho$  mesons at RHIC near mid-rapidity, the photonnucleon center of mass energy ranges from from 9 to 18 GeV, depending on the final state transverse momentum and rapidity. In this region, the  $\rho^0$  photoproduction cross section increases slowly with energy; the  $\gamma p \rightarrow \rho p$  cross section is well described by the soft-Pomeron model [3]. The Pomeron itself may be described as a gluon ladder [4, 5].

A more detailed model considers the photon as a combination of Fock states: a bare photon with virtual  $q\bar{q}$  pairs, plus higher virtual states. It was succesful at describing many of the Deep Inelastic Scattering (DIS) measurements performed at HERA [6] and is also applicable in the UPC environment.

Many models have been proposed to describe the  $\rho$  photoproduction cross 37 section in ultra-peripheral heavy ion collisions. The first calculation used HERA 38 data on  $\gamma p \rightarrow \rho p$  as input to a classical Glauber calculation to predict the cross 39 section with heavy ions [7]; it successfully predicted the  $\rho$  photoproduction 40 cross section at RHIC energies from 62 GeV/nucleon [8] to 130 [9] and 200 41 GeV/nucleon [10], and up to 2.76 TeV/nucleon at the LHC [11]. A later cal-42 culation used a dipole model, treating the  $q\bar{q}$  pair as a dipole, plus a quantum 43 Glauber calculation, which found a cross section about 50% lower, in disagree-44 ment with the data [12]. Most recently, a modification of the quantum Glauber 45 calculation has been proposed; in this model nuclear shadowing reduces the cal-46 culated  $\rho$  cross section to match the data [13]. Other calculations which include 47 nuclear saturation mechanisms, including the colored glass condensate [14, 15]. 48 Two-photon production of  $\pi^+\pi^-$  pairs, but the cross-section is much smaller 49 than for photonuclear interactions [16]. 50

Because of the high photon flux these UPC events have a high probability to 51 be accompanied by additional photon exchanges that excite one or both of the 52 ions, into a Giant Dipole Resonances (GDR) or higher excitation. The GDRs 53 typically decay by emitting a single neutron, while higher resonances usually 54 decay by emitting two or more neutrons [17]. These neutrons have small mo-55 mentum with respect to their parent ion, so largely retain the beam rapidity. 56 For heavy nuclei, the total cross section for multi-photon interactions nearly 57 factorizes [18], with the combined cross section given by an integral over im-58 pact parameter space:  $\sigma(A_1A_2 \rightarrow A_1^*A_2^*\rho) = \int d^2b P_{0\text{Had}}(b) P_1(A^*) P_2(A^*) P(\rho),$ 59 where  $P_{0\text{Had}}(b)$ ,  $P_1(A^*)$ ,  $P_2(A^*)$  and  $P(\rho)$  are the respective probabilities for 60 not having a hadronic interaction, exciting the two ions and producing a  $\rho$ . 61 Each photon-mediated reaction occurs via independent photon exchange, so all 62 four probabilities are tied together only through a common impact parameter 63 [19]. The individual photon-mediated subreactions have a strong impact param-64 eter dependence, so the combined probability is highest for impact parameters 65  $b > \approx 2R_A$ , where  $R_A$  is the nuclear radius. 66

#### <sup>67</sup> 2. Experimental setup and Analysis

This letter reports on the measurement of exclusive  $\rho$  and  $\omega$  meson and direct 68  $\pi^+\pi^-$  photo-production in UPCs between gold ions using the Solenoidal tracker 69 at RHIC (STAR) detector at a center of mass energy of 200 GeV/nucleon, using 70 data recorded during Run 10. The current data sample is about 100 times larger 71 than in previous measurements [10] at this energy. The improved statistics allow 72 for much higher precision studies, leading to two main new results. The first 73 is the observation of a detailed diffraction pattern, clearly showing the first 74 and second dips, with a possible third. This diffraction pattern can be used 75 76 to determine the distribution of the hadronic matter in gold nuclei. Second, the high-statistics  $\pi\pi$  invariant mass distribution cannot be fit with just  $\rho$  and 77 direct  $\pi\pi$  components; an additional contribution from photoproduction of  $\omega$ , 78 with  $\omega \to \pi^+ \pi^-$  is required for an acceptable fit. 79

Five STAR components were used for triggering and event reconstruction in the analysis: the Time Projection Chamber (TPC), Time of Flight system (TOF), Beam Beam Counters (BBCs) and East and West Zero Degree Calorimeters (ZDCs).

The STAR TPC [20] efficiently detects charged tracks with pseudo-rapidities  $|\eta| < 1.5$ , using 45 layers of pad rows in a 2 m long cylinder. In the 0.5 T solenoidal magnetic field, the momentum resolution is  $\Delta p/p = 0.005 + 0.004p$ with p in GeV/c. The TPC can also identify charged particles by their specific ionization energy loss (dE/dx) in the TPC. The dE/dx resolution is 8% for a track that crosses 40 pad rows. This gives good pion/kaon/proton separation up to their respective rest masses.

The other detector components were used solely for triggering. The TOF system that surrounds the TPC in azimuth, with coverage in pseudo-rapidity  $|\eta| < 1$ . At higher rapidities, charged particles are detected using the two BBCs, one on each side of the nominal interaction point. Each is formed with 18 scintillator tiles arranged around the beam pipe, covering a pseudo-rapidity window of  $2 < |\eta| < 5$  [21]. The ZDCs are small hadron calorimeters installed downstream of the collision region to detect neutrons at beam rapidity [22].

The trigger selected events with small multiplicity in the central detector, 98 along with one or more neutrons in each ZDC, along the lines described in [10]. 99 It requires low activity in the TOF detector (at least two and no more than six 100 hits), no charged particles detected in the BBC detectors and finally, showers 101 in both ZDC detectors corresponding to at least the equivalent of one neutron 102 with beam momentum or up to four beam momentum neutrons. The threshold 103 on each ZDC calorimeter was set at 50 ADC channels (the centroid of the one 104 neutron peak sits at 198 channels) making them fully efficient. The integrated 105 trigger luminosity was 1074 ( $\mu b$ )<sup>-1</sup> ( $\pm 10\%$ ); a total of 38 million triggers were 106 recorded. 107

The analysis selected events containing a pair of tracks with opposite charges 108 (like-sign pairs were used as a background measure) that were consistent with 109 originating from a single vertex, located within 50 cm longitudinally of the center 110 of the interaction region. The tracks were required to have at least 14 hits in the 111 TPC (out of a possible 45), and have dE/dx values within  $3\sigma$  of the expected 112 dE/dx for a pion with that trajectory. Both tracks in each pair were required 113 to have a valid hit in the TOF system; this cut rejected events from other beam 114 crossings. It also limited the track acceptance largely to the region  $|\eta| < 1.0$ . 115 The 384,000 events with a  $\pi^+\pi^-$  pair mass in the range  $0.25 < M_{\pi\pi} < 1.5$ 116 GeV were saved for further evalutation. This included photoproduced  $\rho$  and  $\omega$ 117 decays to  $\pi^+\pi^-$  as well as directly photoproduced  $\pi^+\pi^-$  pairs. 118

The largest backgrounds for this analysis were low-multiplicity hadronic interactions (peripheral ion-ion collisions), whith some of their charged particles out of the TPC acceptance. Other backgrounds come from other UPC reactions or from cosmic-rays accompanied by in-time mutual Coulomb exitation. Pure electromagnetic production of  $e^+e^-$  pairs contribute less than 4% to the  $\rho$  peak [9]. The decay  $\omega \to \pi^+\pi^-\pi^0$  produces a  $\pi^+\pi^-$  pair below the  $\rho^0$  peak, but with a larger  $p_T$  than for coherent production; it contributes a few percent (2.7% in



Figure 1: The black histogram shows the distribution of the pion pair transverse momentum. The peak below 100 MeV/c is from the decay of coherently produced  $\pi^+\pi^-$  pairs. The red histogram shows the distribution of the pair momentum for equal sign pion pairs. Both histograms are filled with pairs that come from vertices with only two tracks.

a previous analysis [10]) to the measured incoherent  $\pi^+\pi^-$  pairs. We neglect 126 these minor backgrounds here; they are well within the overall systematic errors. 127 The hadronic backgrounds may be estimated from the like-sign pion pairs. 128 Figure 1 compares the transverse momentum  $(p_T)$  of the  $\pi^+\pi^-$  pair (black his-129 togram) with the corresponding distribution for like-sign pairs (red histogram) 130 in recorded vertices with only two tracks. The signal distribution has a promi-131 nent peak for  $p_T < 100 \text{ MeV/c}$ . This peak is due to coherent photoproduction 132 of pion pairs from the gold nucleus. In this region, the signal to noise ratio is 133 very high; at larger  $p_T$ , the backgrounds are a larger fraction of the signal. 134

The reconstructed events are corrected for acceptance and detection efficiency using a detailed simulation of the STAR detector. A mix of  $\rho$  mesons and non-resonant  $\pi\pi$  events are generated using the STARLight Monte Carlo [7] which reproduces the kinematics of the event, including the mass and  $p_T$ distributions. These events are sent through a complete GEANT simulation of the detector and then embedded in actual 'zero bias' STAR events; this embedding procedure accurately accounts for the detector noise and backgrounds, including overlapping events recorded in the STAR TPC during its sizeable active time windows. Because this single correction includes all components of
the experimental setup as well as the efficiencies of the analysis algorithms, we
refer to it as the "fullChain" efficiency.

A major uncertainty in the reconstruction efficiency stems from uncertainties in the actual ('as-built') positions of the TOF slats, which may not be completely accurately reflected in the simulations; this may affect the measured  $d\sigma/dy$ , particularly at large rapidity. The relative acceptance in  $p_T$  (and  $t_{\perp} = p_T^2$ ) and invariant mass varies only slowly with  $p_T$  or mass, and should be insensitive to the positional uncertainties. The "fullChain" efficiency is almost independent of the pion pair  $p_T$ .

The two ZDC calorimeters detect the neutrons emitted by both beam in mutual electromagnetic dissociation with efficiency close to 100% and energy resolution sufficient to separate up to three neutron peaks. Figure 2 shows the ADC distribution from the West ZDC for events that satisfy a cut which selects events with a single neutron in the East ZDC and an exclusive and coherent photoproduced  $\rho^0$  in the rapidity range |y| < 1.

This analysis considers two classes of nuclear breakup: single neutrons (1n), 159 associated with Giant Dipole Resonance), or any number of neutrons (Xn), from 160 a broad range of photonuclear interactions. Unfortunately, the trigger condition, 161 requiring 1 to 4 neutrons, was incompatible with either of these classes. So, we 162 used the 1n1n events to normalize the cross-section, based on the STARlight 163 calculation of the cross-section. We find the ratio of triggered events to those 164 with single neutrons in each ZDC, using the fit results in Table 1, and use the 165 STARlight ratio of XnXn to 1n1n events to normalize the overall cross-section 166 scale. 167

The relative cross-sections in Table 1 decrease slowly with increasing neutron 168 number; for example, the cross-section for the 2n1n + 1n2n (*i.e.* the two direc-169 tional combinations to get 1 neutron in one direction, 2 in the other are very 170 similar to the 1n1n cross-section. This ratio is considerably larger than is seen for 171 mutual Coulomb dissociation, where one calculation has the 2n1n + 1n2n : 1n1n172 ratio around 0.6 [23] and another finds a ratio around 0.4, albeit at a slightly 173 lower beam energy [24]. Some of this difference is because the requirement of 174  $\rho$  photoproduction selects events with smaller impact parameters, where the 175 photon spectrum is harder [18]. 176

Figure 3 shows the efficiency-corrected, background-subtracted invariant-177 mass of the selected pion pairs selected to have  $p_T < 100 \text{ MeV/c}$ . Events with 178 dipion mass  $M_{\pi\pi} > 600 \text{ MeV/c}^2$  were initially fitted with a modified Söding 179 parametrization [25] which included a relativistic Breit-Wigner resonance for 180 the  $\rho^0$  plus a flat direct  $\pi^+\pi^-$  continuum. This 2-component model was a poor 181 fit to the data, so an additional relativistic Breit-Wigner component was added. 182 to account for  $\omega$  photoproduction, followed by its decay to two pions  $\omega \to \pi^+ \pi^-$ . 183 This leads to the following fit function: 184



Figure 2: The shower energy in the West ZDC by neutron produced by mutual dissociation is shown as a distribution of ADC channels. These events had a single neutron detected on the East ZDC. The peaks corresponding to 1 to 4 neutrons are fitted with Gaussian distributions with standard deviations that grow as  $n\sigma$  with n the number of neutrons and  $\sigma$  the standard deviation of the one neutron Gaussian. The red curve is the sum of all Gaussians which are also displayed individually. The quality of fit is given by  $\chi^2/NDF = 498/88$  which is mostly due to the very small statistical errors.

	1n	2n	3n
1n	$1.38\pm0.24$	$0.57\pm0.11$	$0.39\pm0.07$
2n	$0.57\pm0.11$	$0.23\pm0.04$	$0.18\pm0.03$
3n	$\begin{array}{c} 1.38 \pm 0.24 \\ 0.57 \pm 0.11 \\ 0.40 \pm 0.07 \end{array}$	$0.19\pm0.03$	$0.15\pm0.03$

Table 1: Mutual dissociation cross section (in mb) for events with exclusive coherent  $\rho^0$  photoproduction. The row number shows the number of neutrons detected in the East ZDC and the column number lists the number of neutrons detected in the West ZDC. The cross sections listed in the table are an average of two measurements: one of them uses the West ZDC to set the number of neutrons on that beam with ADC channel cuts defined by the dip between the 1 and 2 neutron peaks, and the other measurement has the East ZDC selecting events in similar manner. These two measurements differ in the off diagonal term and the systematic uncertainty on the selection of the number of neutrons in either ZDC is set to be equal to the deviation from the average value. Statistical errors are small (<1%) and are not listed. Systematic errors arising from the cuts used to select the events added were added in quadrature to the sum in quadrature of the relevant common uncertainties listed in table 3 (17%).

$$\frac{d\sigma}{dM_{\pi^+\pi^-}} \propto \left| A_{\rho} \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma_{\rho}}}{M_{\pi\pi}^2 - M_{\rho}^2 + iM_{\rho} \Gamma_{\rho}} + B_{\pi\pi} + C_{\omega} e^{i\phi_{\omega}} \frac{\sqrt{M_{\pi\pi} M_{\omega} \Gamma_{\omega}}}{M_{\pi\pi}^2 - M_{\omega}^2 + iM_{\omega} \Gamma_{\omega}} \right|^2 + f_{\mu}$$
(1)

185 186 187

189

191

where  $A_{\rho}$  gives the  $\rho$  component,  $B_{\pi\pi}$  is for the direct pions and  $C_{\omega}$  is for the  $\omega$ . The momentum-dependent widths are taken to be

$$\Gamma_{\rho} = \Gamma_0 \frac{M_{\rho}}{M_{\pi\pi}} \left( \frac{M_{\pi\pi}^2 - 4m_{\pi}^2}{M_{\rho}^2 - 4m_{\pi}^2} \right)^{3/2} \tag{2}$$

190 and

$$\Gamma_{\omega} = \Gamma_0 \frac{M_{\omega}}{M_{\pi\pi}} \left( \frac{M_{\pi\pi}^2 - 9m_{\pi}^2}{M_{\omega}^2 - 9m_{\pi}^2} \right)^{3/2},\tag{3}$$

where  $\Gamma_0$  is corresponding pole width for each meson. For the  $\omega$ , the  $9m_{\pi}^2$ term is because  $\omega$  decay is dominated by the three-pion channel. Here,  $f_p$  is a quadratic polynomial that describes the remaining remnant background. In this all-rapidity fit, the masses and widths of the  $\rho$  and  $\omega$  were allowed to float, making for a total of ten parameters: four masses/widths, three amplitudes, and three parameters for the polynomial background.

In Fig. 3, the fitted  $\rho$  component is shown by the full blue line, with the direct  $\pi\pi$  component shown in dashed black, the dashed blue line shows the interference between the two components. The full red line shows the fitted  $\omega$ component and the dashed red line shows the interference between the  $\rho^0$  and the  $\omega$  components.

Table 2 shows the results of this fit. The particle masses and widths are all in quite good agreement with their generally accepted values [26]. The  $\omega$ mass is considerably wider than the standard value, but it is broadened because of detector resolution; at the  $\omega$  peak, the detector resolution is about 0.0085 GeV/c<sup>2</sup>. The fit  $\chi^2/DOF = 314/297$  is good.

The  $\omega$  amplitude *C* is small, but the  $\omega$  is clearly visible through its interference with the  $\rho$ . This interference produces the small kink in the spectrum just above 800 MeV/c<sup>2</sup>. The  $\omega$  amplitude agrees with a prediction based on the  $\omega$  photoproduction cross-section and the most recent value for the  $\omega$  branching ratio  $(1.5 \pm 0.1\%$  to  $\pi^+\pi^-$ ; this prediction is shown by the solid 'STARlight' lines in Fig. 4.

The DESY-MIT group measured a phase angle of  $1.68 \pm 0.26$ , close to our 1.73  $\pm 0.13(stat.) \pm 0.17(syst.)$ . The systematic error on  $\phi_{\omega}$  is determined by varying the lower range of the fit to the pion pair invariant mass distribution between values of 520 and 600 MeV. This agreement is better than might be expected, since the DESY-MIT experiment used much lower energy photons, in a



Figure 3: The  $\pi^+\pi^-$  invariant-mass distribution for all selected  $\pi\pi$  candidates with  $p_T < 100$  MeV/c. The black markers show the data (in 2.5 MeV/c<sup>2</sup> bins). The black curve is the modified Söding fit to the data in the range  $0.6 < M_{\pi\pi} < 1.3$  GeV. The  $\rho^0$  Breit-Wigner component of the fitted function is shown with a blue curve and the constant non-resonant pion pair component is displayed with a black-dashed one. The interference between non-resonant pion pairs and the  $\rho^0$  meson is shown with a blue-dashed curve and a small third order polynomial shown with a cyan-dashed curve is used to account for a remnant background. The Breit-Wigner distribution for the  $\omega$  mesons is shown with a red curve and the interference between  $\rho^0$  and  $\omega$  is shown with a red-dashed curve.

Fit Parameter	value	units
$M_{\rho}$	$0.7757 \pm 0.0006$	$GeV/c^2$
$\Gamma_{ ho}$	$0.1475 \pm 0.0014$	${ m GeV/c^2}$
$A_{ ho}$	$1.511\pm0.005$	
$B_{\pi\pi}$	$-1.176 \pm 0.016$	$(\mathrm{GeV/c}^2)^{-1/2}$
$C_{\omega}$	$0.0626\pm0.004$	
$M_{\omega}$	$0.7838 \pm 0.0009$	$\mathrm{GeV}/\mathrm{c}^2$
$\Gamma_{\omega}$	$0.0163 \pm 0.0017$	$\mathrm{GeV/c^2}$
$\phi_{\omega}$	$1.73\pm0.13$	radians
$f_p p_0$	$3.566 \pm 0.304$	
$f_p p_1$	$-5.084\pm0.53$	
$f_p p_2$	$1.743\pm0.24$	

Table 2: Parameter values extracted by fitting the invariant mass distribution of selected pion pairs with the parametrization listed in Eq. 1. Three additional parameters, giving the polynomial background, are also listed.

regime where production proceeds via both single meson and Pomeron exchange. 223 Other experiments have studied  $\rho - \omega$  interference using photoproduction to the 224  $e^+e^-$  final state, where the  $\omega$  is more visible, but the cross sections are much 225 smaller, or via the reaction  $e^+e^- \rightarrow \pi^+\pi^-$ , and gotten similar phase angles [28]. 226 Similar fits were performed in five bins of rapidity symmetric about y=0227 and variable widths. The number these bins and their widths is determined 228 by the desire to continue to produce fits as good as the one described above. 229 Each bin had close to 100K pion pairs and the values of  $M_{\omega}$  and  $\Gamma_{\omega}$  were fixed 230 to the values extracted from the fit to the rapidity integrated pion pair mass 231 distribution. Figure 4 shows the ratios |B/A| and C/A at each of the five bins 232 in rapidity. The |B/A| ratio quantifies the fraction of non-resonant pion pairs 233 in these measurements. Both |B/A| and C/A are, within the total errors, flat 234 as rapidity varies. This shows that these ratios do not have a large dependence 235 on the photon energy. 236

The average value of the |B/A| ratio  $0.79\pm0.01~(stat.)\pm0.08~(syst.)~(\text{GeV/c}^2)^{-1/2}$ agrees, within errors, with the value reported in the previous STAR publication [10]:  $(0.89\pm0.08~(stat.)\pm0.09~(syst.))$ . The same ratio has been measured at the higher energy of the LHC (2.76 TeV per nucleon) by the ALICE collaboration [11] which finds a smaller value  $(0.50\pm0.04~(stat.)+0.10-0.04~(syst.))$ . The average |C/A| value extracted from the fits in rapidity bins is  $0.040\pm0.040\pm0.0054~(stat.)\pm0.0048~(syst.)$ .

Figure 5 shows the acceptance corrected distribution of  $\rho^0$  mesons detected 244 in events with only two tracks out of the triggered vertex The asymmetry be-245 tween positive and negative rapidity gives a measure of the rapidity-dependent 246 systematic uncertainties in the cross section. As noted above, these are likely 247 due to asymmetries in the as-built location of the TOF counters. The magni-248 tude of this uncertainty grows slowly from mid-rapidity to reach a value of 4% at 249 y = 0.7. Since the actual lengths of the TOF system are accurately known, this 250 uncertainty does not apply for the rapidity-integrated measurements presented 251



Figure 4: The left panel shows the ratio |B/A| of amplitudes of non-resonant  $\pi^+\pi^-$  and  $\rho^0$  mesons. The previous STAR results are shown with blue-filled circles. The right panel shows the ratio |C/A| of the  $\omega$  and  $\rho^0$  amplitudes. The data is shown with red markers, while the red band includes the relevant systematic errors. The DESY-MIT |C/A| measurement is shown with a green band. This measurement was at considerably lower photon energies; if converted to rapidity, these measurements would appear at large |y|, outside the current plot. The thin cyan line shows |C/A| calculated using STARlight and the most recent branching ratio for  $\omega \to \pi^+\pi^-$  decay [26].

252 here.

Rapidity is related to photon energy k, with, at low  $p_T$ ,  $k = M_{\pi\pi}/2 \exp(\pm y)$ , 253 with the  $\pm$  sign because of the two-fold ambiguity as to which nucleus emitted 254 the photon, away from y = 0, the cross section is dominated by the lower 255 photon energy. So, one can study how the three components of the mass fit 256 vary with energy by dividing the mass spectrum in rapidity. We do this in 257 five bins: |y| < 0.15, 0.15 < |y| < 0.35, and |y| > 0.35. In each bin, fits to 258 the corresponding invariant mass distributions were performed to separate the 259 resonant pion pairs from the flat non-resonant distribution. The integral of the 260  $\rho$  Breit-Wigner function for masses ranging from  $2M_{\pi}$  and  $M_{\rho} + 5\Gamma_{\rho}$  is then 261 used to extract a correction to the raw counts of  $\rho^0$  candidates. This correction 262 is then applied to all other results presented here. 263

Finally, the resulting distribution has been scaled by the "fullChain" efficiency defined above. The rapidity dependence of this efficiency has a bell shape with a maximum at 13% at  $y \approx 0.1$ . It is slightly asymmetric because of inefficiencies on one of the TPC West (rapidity < 0) sectors.

The systematic uncertainties in these measurements fall into two classes, 268 either an overall scale factor on the cross-section, or point-to-point. The over-269 all scale factor is usually dominant. The scaling from the rapidity distribution 270 extracted from 1n1n events to the previously measured XnXn distribution uses 271 a correction, extracted from the event generator STARLight and introduces 272 a 6% uncertainty related to the uncertainty in the neutron data used as in-273 put to STARlight, squared because we detect neutrons in both beams. This 274 uncertainty applies only to the XnXn results presented in this report. The un-275 certainty in the integrated luminosity is 10%, as with previous measurements 276 [10] mainly driven by the fraction of the total Au+Au cross section accessi-277 ble with the trigger used to collect this data. The selection of the number of 278 neutrons produced in mutual electromagnetic dissociation is based on the ZDC 279 calorimeters response. We allocate a 5% uncertainty to this neutron counting 280 due to small non-linearities in the calorimeters and overlaps between one and 281 many neutron distributions. We also assigned a 7% uncertainty due to mod-282 elling of the TOF system in the simulation. The track reconstruction efficiency 283 for the STAR TPC has a 6% uncertainty [20] while the efficiency of the vertex 284 finder is known with a 5% uncertainty driven by the effect of backgrounds. The 285 uncertainty in how often the BBC detectors will veto good UPC events is due 286 to fluctuating backgrounds. Even with use of embedding techniques, we esti-287 mate that these veto conditions introduce a 2% uncertainty to the results. The 288 same-sign pion pair distributions are the best estimators for the background for 289 these two track events. The background subtraction was done at the level of raw 290 histograms or after a fit to the background to eliminate statistical fluctuations. 291 The relative deviation between those two procedures found in the fully corrected 292 distributions is found to be 1.5%. Adding all these systematic uncertainties in 293 quadrature leads to a 19% overall common uncertainty. This uncertainty is a 294 bit higher than in our comparable previous publication [10], largely because of 295 additional uncertainties associated with the pileup and the more complex trigger 296 and which is required to deal with the higher luminosities. Table 3 summarizes 297

Name	Value	Comment
Luminosity	10.%	
STARLight model	6.%	only for XnXn results
ZDC	5.%	ADC ch. to num. neutrons
TOF geometry modeling	7.%	
TPC tracking efficiency	6.%	STAR standard [20]
Vertex Finder efficiency	5.%	Background driven
BBC veto in trigger	2.%	Background driven
"fullChain" efficiency	7.%	Ev. Gen., Material budget
Background subtraction	1.5%	
Quadrature Sum	18.1%	

Table 3: Summary of all common systematic uncertainties identified in the generation of both the rapidity distribution shown in Fig. 5 and the -t distributions shown in figures 6 and 7. All these uncertainties are presented as a percent fractions of the measured quantities.

rapidity	PID cut	Fit to eff.	Number of track hits	TOF asymmetry
-0.70, -0.5	8.%	0.25%	0.2%	5%
-0.5, 0.	5.%	0.25%	0.05%	3.6%
0., 0.5	5.%	0.25%	0.05%	3.6%
0.5 - 0.7	8.%	0.25%	0.2%	5%

Table 4: Four point-to-point systematic uncertainties for the rapidity distribution shown in Fig. 5 shown as a percent of the measured cross section in four rapidity ranges. Pions are identified in the TPC their specific energy loss (dE/dx), based on how close (in standard deviations) they are to the calculated energy loss. Truncated distributions are used to improve accuracy [29]. Those cuts were varied simultaneously in the data and simulation to determine the systematic uncertainty due to particle identification. Good tracks are selected based on a minimum number of space points included in their pattern recognition and fits. The minimum number of hits was varied to estimate the systematic uncertainty. The rapidity distribution of photoproduced  $\rho^0$  mesons with symmetric beams should also be a symmetric function around y=0. Relative deviations from the average between measurements at the same absolute rapidity are used to quantify the asymmetry of the measured rapidity distribution.

<sup>298</sup> all the common systematic uncertainties identified in this measurement.

The main factors that introduced point-to-point systematic uncertainties in 299 the rapidity or  $p_T$  distributions were in the track selection and particle iden-300 tification. The systematic uncertainties were evaluated by varying the track 301 quality cuts and PID cuts around their central value in both the data and sim-302 ulation, and seeing how the final results varies. The largest relative deviations 303 listed in wider bins are the point-to-point systematic uncertainties. Table 4 lists 304 the uncertainties in the rapidity distribution. Table 5 list the point-to-point 305 uncertainties allocated for the  $p_T$  distribution. 306

Recently, the ALICE collaboration has also studied  $\rho$  photoproduction, in lead-lead collisions at the Large Hadron Collider (LHC) [11]. They fit their dipion mass distribution in the range from 0.6 to 1.5 GeV<sup>2</sup> to a function like Eq. 1, but without the  $\omega$  component, finding masses and widths consistent with the standard values. Their cross-section values,  $d\sigma/dy$  were about 10% above



Figure 5: The cross section as function of rapidity for exclusively photo-produced  $\rho^0$  mesons in (bottom) events with a single neutron detected on both ZDC detectors (1n1n) and for any number of neutrons XnXn (top). The data are shown with red markers. The statistical errors are smaller than the symbols, the orange band shows the quadrature sum of the point-topoint systematic uncertainties. The red box at  $y \sim -0.9$  shows the quadrature sum of the common systematic uncertainties. The black histograms are the STARLight calculation for  $\rho^0$  mesons with mutual dissociation. The blue markers in the top panel show the previous STAR measurement [10],

312 the STARlight

Figure 6 shows the differential cross section  $d\sigma/dt$  for  $\rho^0$  mesons after like-313 sign background subtraction, with t the Mandelstam variable  $t = t_{\parallel} + t_{\perp}$  with 314  $t_{\parallel} = -M_{\rho}^2/(\gamma^2 e^{\pm y})$  almost negligible and  $t_{\perp} = -(p_T^{pair})^2$ . The number of  $\rho^0$  mesons is obtained from a simple scaling by a common factor of 0.75 extracted 315 316 from comparisons between the number of pion pairs with invariant masses rang-317 ing from 500 MeV/ $c^2$  to 1.5 GeV/ $c^2$  and the integral of the  $\rho^0$  Breit-Wigner 318 function extracted from fits in rapidity and -t bins. In all comparisons, the 319 integrals are performed from  $2M_{\pi}$  to  $M_{\rho} + 5\Gamma_{\rho}$ . The yield of  $\rho^0$  mesons is also 320 corrected for the compounded effects of tracking reconstruction and geometri-321 cal acceptance, vertex finding efficiency and the finite track and TOF detector 322 matching efficiency extracted from the embedded simulations. This correction 323 is flat in t and has an average value of 6.4% over all rapidity values. Finally 324 the distribution is normalized by the luminosity integrated over all data runs 325 used in this analysis. The large peak in  $d\sigma/dt$  for  $|t| < 0.1 \text{GeV}^2$  is expected 326 from coherent photoproduction. At substantially larger |t|, production should 327 be dominated by incoherent interactions with individual nucleons in the target 328 ion. 329

We separate the  $\rho^0$  t spectrum into coherent and incoherent components based on the shape of the distribution in Fig. 6. Because of the neutron requirement in the trigger, and the presence of Coulomb excitation, we cannot use the presence of neutrons from nuclear breakup as an event-by-event sign of incoherence [30].

The incoherent component is fit with the so called "dipole" form factor

$$\frac{d\sigma}{dt} = \frac{A/Q_0^2}{(1+|t|/Q_0^2)^2}$$

used to describe low  $Q^2$  photon-nucleon interactions [31]. The fit range for the 335 XnXn events starts at  $t = 0.2 \text{ GeV}^2$  (above the coherent production region) and 336 extends to  $-t = 0.45 \text{ GeV}^2$  as shown by the black curve in the figure. The upper 337 limit in t is chosen to reduce the contamination from hadronic interactions. For 338 the events with mutual dissociation into any number of neutrons (XnXn), the fit 339 finds  $A = 3.46 \pm 0.02$ ,  $Q_0^2 = 0.099 \pm 0.015 \,(\text{GeV/c})^2$ , with  $\chi^2/NDF = 19/10$ . For 340 events with mutual dissociation into single neutrons (1n1n), the fit parameters 341 are:  $A = 0.191 \pm 0.003$ ,  $Q_0^2 = 0.099 (fixed) (\text{GeV/c})^2$ , with  $\chi^2 / NDF = 13.7/10$ . 342 The integral of the fit to the incoherent component in the XnXn events results 343 in a value of cross section  $\sigma_{incoh} = 2.89 \pm 0.02 \ (stat.) \pm 0.03 \ (syst.)$  mb in 344 the measured rapidity range |y| < 1. The integral of the coherent component 345 discussed below amounts to  $6.49 \pm 0.01 (stat.) \pm 0.01 (syst.)$  mb. The incoherent 346 component of the distribution extracted from 1n1n events is fitted to the same 347 function as the XnXn distribution. The range of -t and rapidity values is also 348 the same. The integral of the fit to the incoherent component in the 1n1n events 349 results in a value of cross section  $\sigma_{incoh} = 0.162 \pm 0.01 \ (stat.) \pm 0.005 \ (syst.)$  mb. 350 The integral of the 1n1n coherent component amounts to  $0.7696 \pm 0.004 (stat.) \pm$ 351 0.004 (syst.) mb. 352

The corresponding ratios are:

$$\begin{split} \sigma_{incoherent}^{XnXn} / \sigma_{coherent}^{XnXn} &= 0.445 \pm 0.003(stat.) \pm 0.005(syst.) \\ \sigma_{incoherent}^{1n1n} / \sigma_{coherent}^{1n1n} &= 0.233 \pm 0.014(stat.) \pm 0.007(syst.). \end{split}$$

The difference in the ratio found for 1n1n and XnXn collision is somewhat 353 larger than the previous STAR analysis [10], particularly in the XnXn channel. 354 The ratio difference could come from a variety of sources. First, at large |t|, 355 it is possible for a single photon to both produce a  $\rho^0$  and leave the target 356 nucleus excited, breaking the assumed factorization paradigm. The rate has not 357 been calculated for  $\rho^0$ , but for  $J/\psi$ , the cross-section is substantial [32]. The 358 calculated cross-section for vector meson production with excitation is very low 359 for single neutron emission, so this would alter the XnXn cross-section ratio more 360 than the 1n1n. Second, unitarity corrections could play a role here, by changing 361 the distribution of impact parameters between 1n1n and XnXn interactions. 362 These corrections account for the fact that, near grazing incidence, the cost of 363 introducing another low-energy photon into the reaction is small, so four-photon 364 reactions can occur. In these cases, one photon can excite a nucleus, for example 365 to a a GDR, while a second photon can excite the nucleus further, leading to 366 Xn emission rather than 1n [33]. These probabilities are included in STARlight 367 [19]. The additional photon alters the impact parameter distributions for the 368 1n1n and XnXn channels; the XnXn channel will experience a slightly larger 369 reduction at small |t| due to interference from the two production sites; this may 370 lead to slightly different measured slopes and coherent/incoherent ratios. There 371 may also be a larger effect due to a larger non-uniform photon illumination of 372 the target nucleus. 373

The coherent component of the t distribution is then extracted by a sub-374 tracting the power law fit to the incoherent tail. The resulting differential cross 375 section for  $\rho^0$  photoproduction accompanied with mutual dissociation of the nu-376 clei into any number of neutrons (XnXn ) and only one neutron (1n1n ) is shown 377 in Fig. 7 with red and black markers respectively. In both types of events, two 378 well defined peaks can clearly be seen. For both types of events, the location 379 of the first dip is at  $-t = 0.018 \pm 0.005$  (GeV/c)<sup>-2</sup>. A second dip is visible at 380  $0.043 \pm 0.01 ~(\text{GeV/c})^{-2}$ . An exponential function is used to characterize the 381 spectrum below the first peak  $(0.0024 < |t| < 0.0098 \text{ (GeV/c)}^2)$ . The measured 382 slope is  $426.4 \pm 1.8$  (GeV/c)<sup>-2</sup> for the XnXn events and  $407.8 \pm 3.2$  (GeV/c)<sup>-2</sup> 383 for the 1n1n events. 384

Such features of the diffraction pattern indicate, to first order, the coherent diffraction off the same large target for both 1n1n and XnXn events.

At very small t,  $|t| < 10^{-3}$  GeV<sup>2</sup>, both cross section flatten out and turn downward, as can be seen in the insert in Fig. 7. This is expected due to destructive interference between  $\rho$  production on the two nuclear targets [34, 35]. The systematic uncertainties identified for the extraction of this differential cross-section come in two types, the so called common uncertaities which are the same as the ones extracted for the rapidity distribution shown in Fig. 5 and the point-to-point ones described above and listed in Table 5. The red band in



Figure 6: The final -t distribution for exclusive  $\rho^0$  mesons detected in events with mutual dissociation into a single neutron (1n1n) is shown with blue markers, the same distribution constructed from events with mutual dissociation into any number of neutrons (XnXn) is shown with red markers. The high t part of those distributions, which is dominated by the contribution from incoherent interactions is fit to a power law shape described in detail in the text. The upper limit of the fit was selected to avoid contamination from hadronic interactions, which tends to be more pronounced in the XnXn events. The fit result is shown with a solid line within the fit range and with a thinner line outside it. The incoherent component as generated by STARlight for XnXn events is shown with small black markers.

$-t[(\mathrm{GeV}/c)^2]$	track sel.	pion PID	Incoh. comp. sub.
0 - 0.02	0.2%	8%	0.5%
0.02 - 0.04	0.2%	8%	3.0%
0.04 - 0.1	0.2%	8%	8.5%

Table 5: Three point-to-point systematic uncertainties identified in the generation of the -t distribution shown in Fig. 7, as a per-cent of the measured cross section in three -t ranges. The PID and track selection uncertainties are described in the text. The uncertainty in the incoherent component subtraction was estimated by selecting the biggest relative deviation from the default value and cross sections extracted by changing the value of the fit parameters by one standard deviation while the other parameters remain at the default fit value.

Fig. 7 is the sum in quadrature of all systematic uncertainties as well as the statistical errors.

Similar features have been reported at the LHC in Pb + Pb at 2.76 TeV per nucleon by the ALICE collaboration [11]. They have extracted the transverse momentum distribution of the  $\rho^0$  mesons and find a prominent coherent peak and a well defined dip around 0.12 GeV/c while this analysis finds it at -t =0.018 ± 0.005 (GeV/c)<sup>-2</sup>. The sartre event generator run in UPC mode at RHIC energies [36] produces a Au nuclei recoil after  $\rho^0$  elastic scattering with a remarkable agreement with the  $\rho^0$  t distribution presented in this report.

The shape of  $d\sigma/dt$  for coherent photoproduction is determined by the po-403 sition of the interaction sites within the target, and one can, in principle, deter-404 mine the density distribution of the gold nucleus via a two dimensional Fourier 405 transform of  $d\sigma/dt$ . The beam energies at RHIC are high enough so that, for 406  $\rho$  photoproduction at mid-rapidity, the longitudinal density distribution may 407 be neglected and the ions may be treated as discs. They are also azimuthally 408 symmetric, so the radial distribution may be determined with a one-dimensional 409 Fourier-Bessel (Hankel) transformation: 410

411

$$F(b) \propto \frac{1}{2\pi} \int_0^\infty dp_T p_T J_0(bp_T) \sqrt{\frac{d\sigma}{dt}}$$
(4)

Here, it is calculated numerically over Figure 8 shows the result of this transform, 412 taken over the the region  $|t| < 0.06 \text{ GeV}^2$ . The tails of the b distribution 413 may suffer from inteference effects [35]. We varied the upper limit used in the 414 transform in the range 0.05 to 0.09 GeV<sup>2</sup>; this led to substantial variation at 415 small b, shown by the cyan region in Fig. 8. The origin of this variation is 416 not completely clear, but it may be related to aliasing due to the lack of a 417 windowing function [37], or because of the limited statistics at large t. There 418 is much less variation at the edges of the distribution. This leads us to believe 419 that the transform can be used in the region where b ranges from  $\sim 4-7$ 420 fm. In this region, we determine the full-width half-maximum (FWHM) of the 421 distribution to be  $2 \times (6.17 \pm 0.12)$  fm. This FWHM is a measure of the hadronic 422 size of the gold nucleus. With theoretical input, it could be compared with 423 the electromagnetic (proton) radius of gold, as determined by electromagnetic 424 scattering. The difference would be a measure of the neutron skin thickness of 425



Figure 7: Fully normalized coherent diffraction patterns for  $\rho^0$  mesons detected in exclusive XnXn events is shown with red markers. The same distribution but extracted from 1n1n events is shown with balck markers. The filled bands shows the sum in quadrature of all systematic uncertainties listed in table 4 and the statistical errors, which are shown as vertical lines. The insert shows, with finer binning at low  $p_T$ , the effects of the destructive interference between photoproduction with the photon emitted by any of the two ions.



Figure 8: The histogram shows the normalized nucleon distribution in the transverse plane, the result of a two-dimensional Fourier transform (Hankel transform) of the XnXn and 1n1n diffraction patterns shown in Fig. 7. The integration is limited to a region where data is available; in the range  $0 < |t| < 0.06 \text{ GeV}^2$ . The cyan error band shows the effect of changing the maximum t to 0.05, 0.07 and 0.09 GeV<sup>2</sup>. In order to highlight the similarity of both results at their falling edges, the resulting histograms are scaled by their integrals from -12 to 12 fm. The FWHM of both transforms is  $2 \times (6.17 \pm 0.12)$  fm consistent with the coherent diffraction of  $\rho^0$  mesons off an object as big as the Au nuclei.

<sup>426</sup> gold, something that is difficult to measure [38], [39].

Because of the possibility of  $\rho$  absorption the  $p_T$  introduced by the photon, the non-uniformity of the photon field (it is stronger on the 'near' side of the nucleus) and the effect of interference between the two production sites, corrections are required, and care must be used in interpreting the transform.

## 431 **3. Summary and conclusion**

<sup>432</sup> In conclusion, STAR has made a high-precision study of  $\rho$ ,  $\omega$  and direct <sup>433</sup>  $\pi^+\pi^-$  photoproduction in 200 GeV/nucleon gold-on-gold ultra-peripheral col-<sup>434</sup> lisions, using 394,000  $\pi^+\pi^-$  pairs. We measure the cross section  $d\sigma/dt$  over a <sup>435</sup> wide range. The incoherent cross section has a very similar shape to the HERA <sup>436</sup> data on  $\gamma p \to \rho^0 p$ , while the coherent contribution can be used to map out the

density distribution of the gold target nuclei. We also fit the invariant mass 437 spectrum to a mixture of  $\rho$ ,  $\omega$  and direct  $\pi^+\pi^-$  (including interference terms). 438 The ratio of  $\rho$  to direct  $\pi\pi$  is similar to that in previous measurements, while the 439 newly measured  $\omega$  contribution is comparable with predictions based on on the 440 previously measured  $\gamma p \to \omega p$  cross section and the  $\omega \to \pi^+ \pi^-$  branching ratio. 441 The relative fractions of  $\rho$ ,  $\omega$  and direct  $\pi^+\pi^-$  do not vary significantly with 442 rapidity, indicating that they all have a similar dependence on photon energy. 443 This measurement provides a nice lead-in to future studies of photo- and 444

electro- production at an electron-ion collider (EIC) [40], where nuclei may be probed with photons at a wide range of  $Q^2$  [41].

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