

Paper Preview:  
**Measurements of  
 $W^+ / W^-$  cross-section ratio  
in *pp* Collisions at RHIC**

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# Information

- **Title:** Measurements of  $W^+ / W^-$  cross-section ratio in  $pp$  collisions at RHIC
- **Target Journal:** PRD
- **PA:** Jae Nam, Matt Posik, Bernd Surrus
- **Webpage:** [link](#)
- **Analysis Note:** [link](#)
- **Paper draft:** [link](#)
- **Previous Publications:**
  - Z cross section and AN, PLB 854 (2024) 138715
  - Unpol W/Z cross section + ratio, PRD 103 (2021) 1, 012001
  - W-AL, PRD 99 (2019) 5, 051102
  - W/Z-AN, PRL 116 (2016) 13, 132301

## Abstract

We report measurements of  $W^\pm \rightarrow e^\pm \nu$  production cross sections and their ratio in unpolarized  $pp$  collisions at a center-of-mass energy of  $\sqrt{s} = 510$  GeV. The measurements were performed within the fiducial region defined by the charged-lepton pseudorapidity  $-1 < \eta_e < 1.5$ , and transverse energy  $25 \text{ GeV} < E_{T,e} < 50 \text{ GeV}$ . The cross-section ratio,  $\sigma^{\text{fid}}(W^+ \rightarrow e^+ \nu) / \sigma^{\text{fid}}(W^- \rightarrow e^- \nu)$  and differential cross sections,  $d\sigma(W^\pm \rightarrow e^p m \nu) / d\eta_e$ , are sensitive to the unpolarized sea quark ratio,  $\bar{d}/\bar{u}$ , and are measured as a function of the leptonic pseudorapidity within the studied region. The results are based on  $pp$  collision data collected with the STAR detector at RHIC in 2017, which correspond to an integrated luminosity of  $340 \text{ pb}^{-1}$ , and found to be consistent with previous STAR publications. Perturbative predictions at  $\mathcal{O}(\alpha_s^1)$  accuracy based on the ResBos2 framework with the latest global fit results, supplemented with next-to-leading-log corrections and resummation, are compared to the combined STAR data. While the predicted cross-section ratio describes the data well, the predictions for the differential cross sections overestimate the measurements by approximately  $2\sigma$ .

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Close communication with PAs (Pavel Nadolsky, et al.)  
For this presentation, **FEWZ** framework is used (w/o NLL)

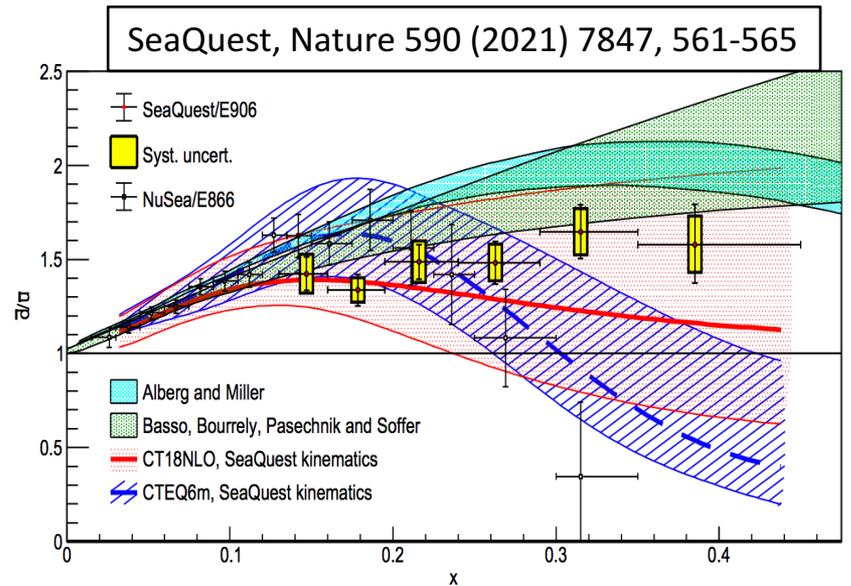
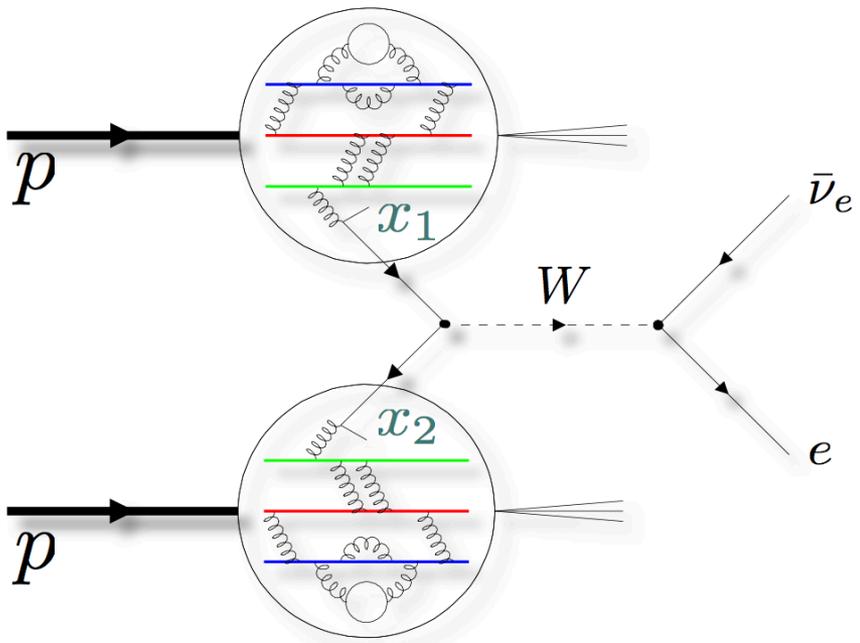
This is also the case for unpol. W/Z Run 2011-13  
Run 17 Z pub. under investigation

# Motivation

- $W$  ( $W \rightarrow ev$ ) in  $pp$  at LO
  - $d\sigma^{W^+} \propto u(x_1)\bar{d}(x_2) + u(x_2)\bar{d}(x_1)$
  - $d\sigma^{W^-} \propto \bar{u}(x_1)d(x_2) + \bar{u}(x_2)d(x_1)$

$$\rightarrow R_W = \frac{\sigma^{W^+}}{\sigma^{W^-}} \sim \frac{u(x_1)\bar{d}(x_2) + u(x_2)\bar{d}(x_1)}{\bar{u}(x_1)d(x_2) + \bar{u}(x_2)d(x_1)}$$

- Sensitive to non-perturbative SU(2)  $\bar{d}/\bar{u}$  asymmetry
- Momentum scale set by the  $W$  mass, complementary to fixed-target experiments
- Provides robustness test of STAR  $W/Z$  reconstruction scheme
- Key features of  $W \rightarrow ev$  in  $pp$ 
  - High  $p_T$  electron.
  - Large imbalance in  $p_T$  due to missing  $\nu$ .

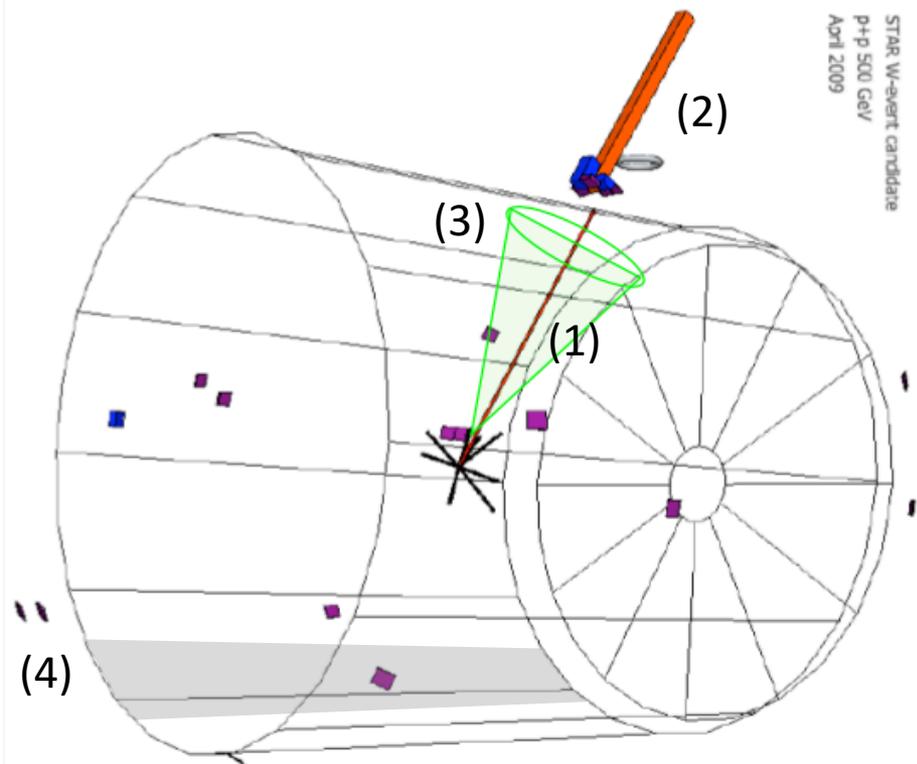


# Data & MC

- Data: STAR Run 2017  $pp \sqrt{s} = 510 \text{ GeV}$
- Trigger set: pp500\_production\_2017
- Production tag: P20ic
- Trigger: L2BW || L2EW
- Run list: /star/u/jaenam/W2017/lists/runnumber.list (2691 runs,  $340 \text{ pb}^{-1}$ )
  - 2.6% correction based on Xiaoxuan's luminosity study
  - $352 \text{ pb}^{-1}$  (Jamie's table)  $-2.6\% = 343 \text{ pb}^{-1}$ ; reporting as  $340 \text{ pb}^{-1}$ )
  - Same has been done for previous publication
- TPC Sec 20 masked out
- Simulation: Request ID 20201502 ([link](#))
  - Pythia 6, CTEQ5L,  
STAR Tune based on Perugia 2012 (minimally affected)

# Event Reconstruction

Event Display of typical W event



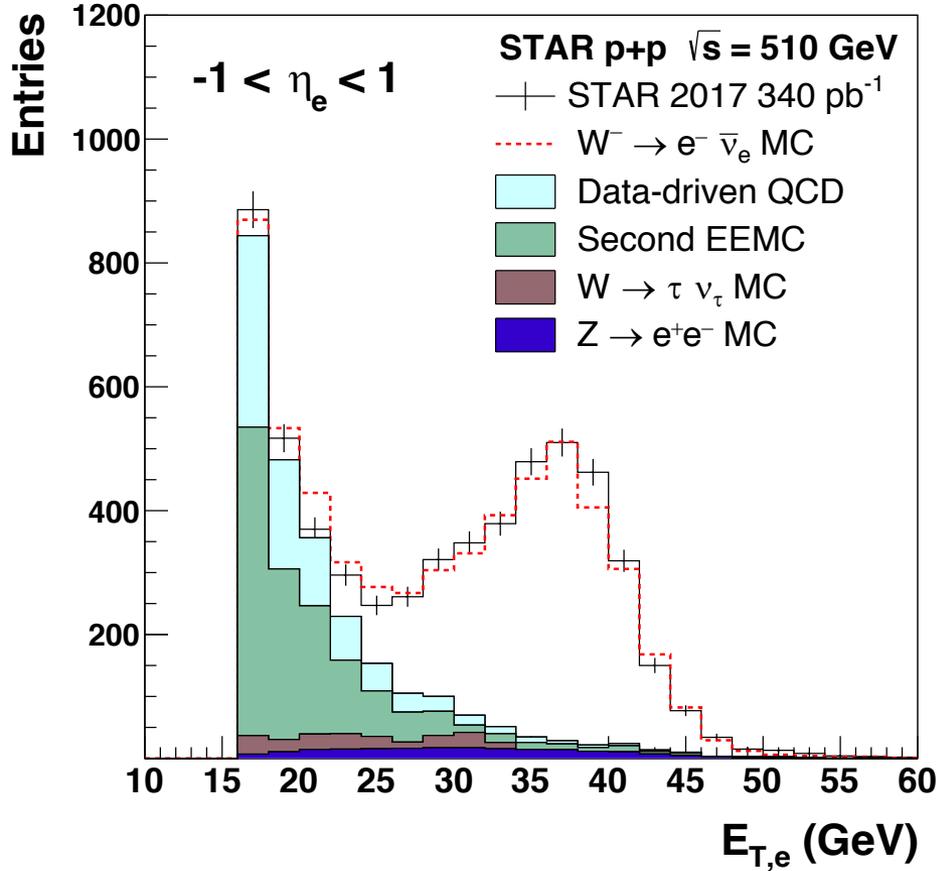
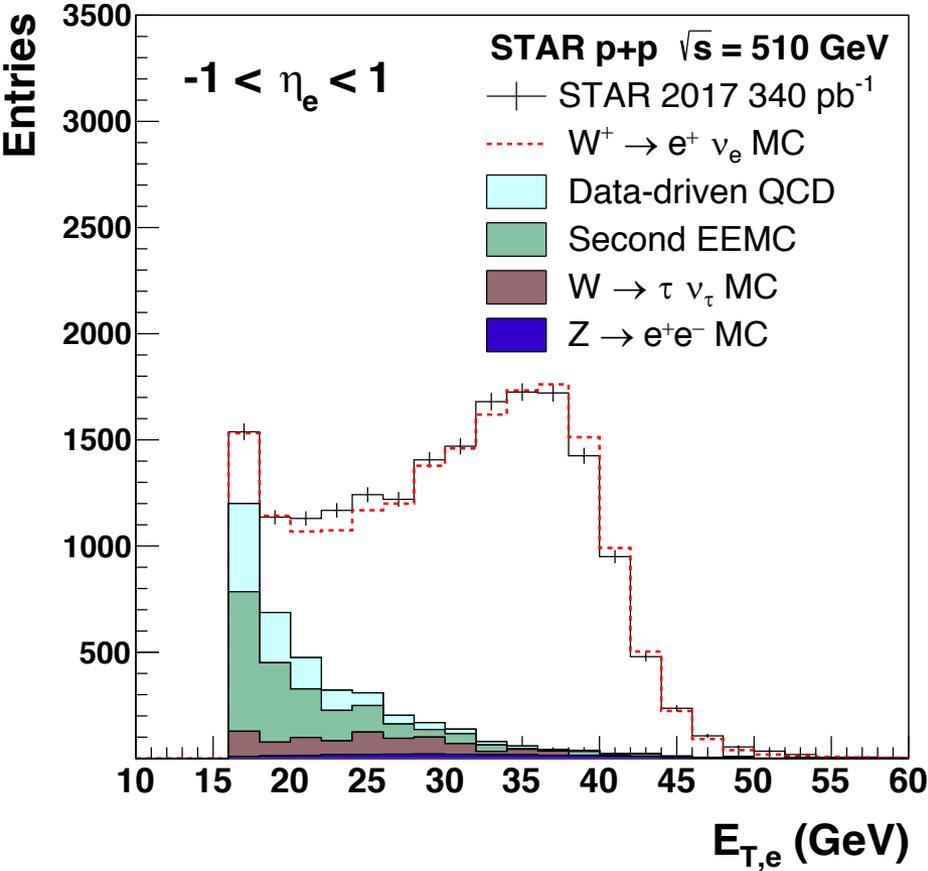
- 1) A high momentum track is identified ( $p_T > 10$  GeV)
- 2) Energy cluster is formed ( $2 \times 2$  towers, each covering  $0.05 \times 0.05$  in  $\eta \times \phi$ ; cluster:  $0.1 \times 0.1$ )
- 3) Isolation requirement ( $E_T^{2 \times 2} / E_T^{\Delta R < 0.7} \sim 1$ ,  $E_T^{2 \times 2} / E_T^{4 \times 4} \sim 1$ )
- 4) Backward (neutrino direction) energy flow requirement ( $E_T^{\Delta \phi \sim \pi} < 11$  GeV)
- 5) Energy imbalance ( $p_{T, bal}$ ) ( $\hat{p}_{T, e} \cdot (\vec{p}_{T, e} + \sum \vec{p}_{T, jet}) > 16$  GeV)

# Selection cuts

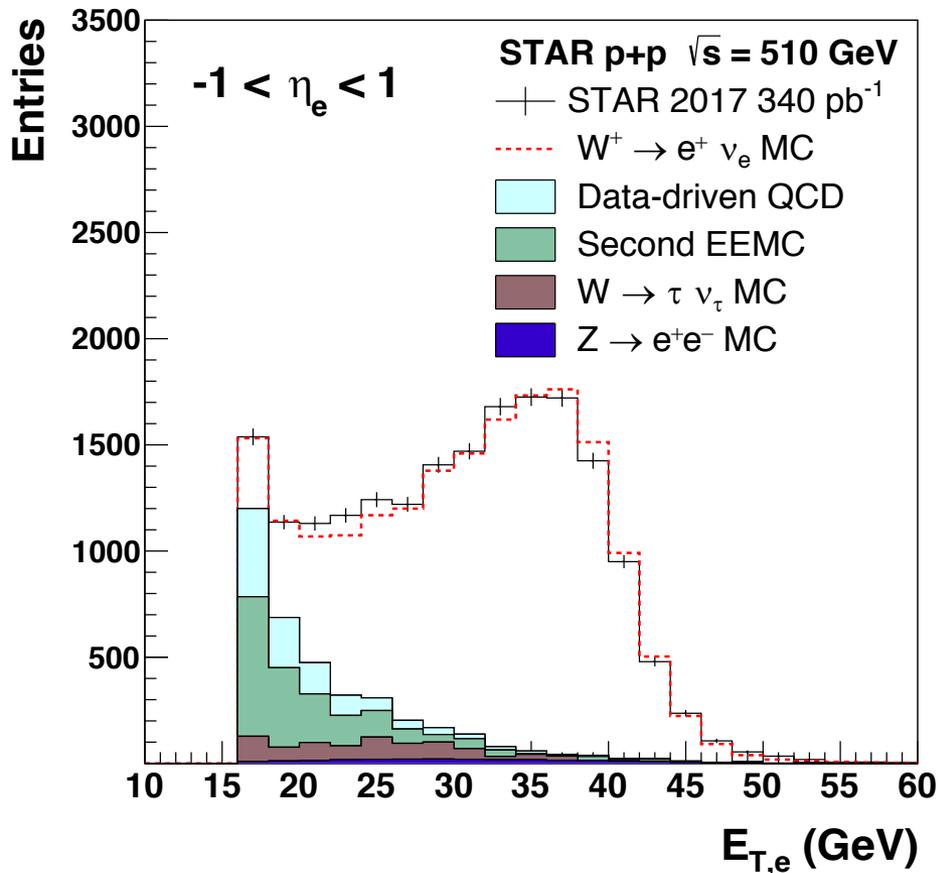
\*same as PRD 103 (2021) 1, 012001

|             | Barrel ( $-1 < \eta < 1$ )   | Endcap ( $1 < \eta < 1.5$ )  |
|-------------|--|--|
| Trigger     | L2BW   | L2EW   |
| Vertex      | Rank $> 0$<br>$ Z_{vtx}  < 100 \text{ cm}$   |  |
| Track       | $p_T > 10 \text{ GeV}$<br>$N_{hit} > 15$<br>$N_{hit}/N_{pos} > 0.51$<br>$R_{TPC,in} < 90 \text{ cm}$<br>$R_{TPC,out} > 160 \text{ cm}$   | $p_T > 7 \text{ GeV}$<br>$N_{hit} > 5$<br>$R_{TPC,in} < 120 \text{ cm}$<br>$R_{TPC,out} > 70 \text{ cm}$   |
| Cluster     | $\Delta R_{trk,cls} < 7 \text{ cm}$<br>$E_T^{2 \times 2} > 16 \text{ GeV}$   | $\Delta R_{trk,cls} < 10 \text{ cm}$<br>$N_{ESMD} > 20$  |
| Isolation   | $E_T^{2 \times 2} / E_T^{2 \times 2} > 0.96$<br>$E_T^{2 \times 2} / E_T^{\Delta R < 0.7} > 0.82$   | $E_T^{2 \times 2} / E_T^{2 \times 2} > 0.97$<br>$E_T^{2 \times 2} / E_T^{\Delta R < 0.7} > 0.88$<br>$R_{ESMD}^{7 \times 7 / 41 \times 41} > 0.6$ |
| W Selection | signed- $p_{T,bal} > 16 \text{ GeV}$<br>$E_T^{\Delta \phi \sim \pi} < 11 \text{ GeV}$<br>$ Q \times E_T p_T  < 3.0$<br>$25 \text{ GeV} < E_{T,e} (\equiv E_T^{2 \times 2}) < 50 \text{ GeV}$ | signed- $p_{T,bal} > 20 \text{ GeV}$<br>$ Q \times E_T p_T  < 2.5$   |

# Fig 1: Signal + BG description



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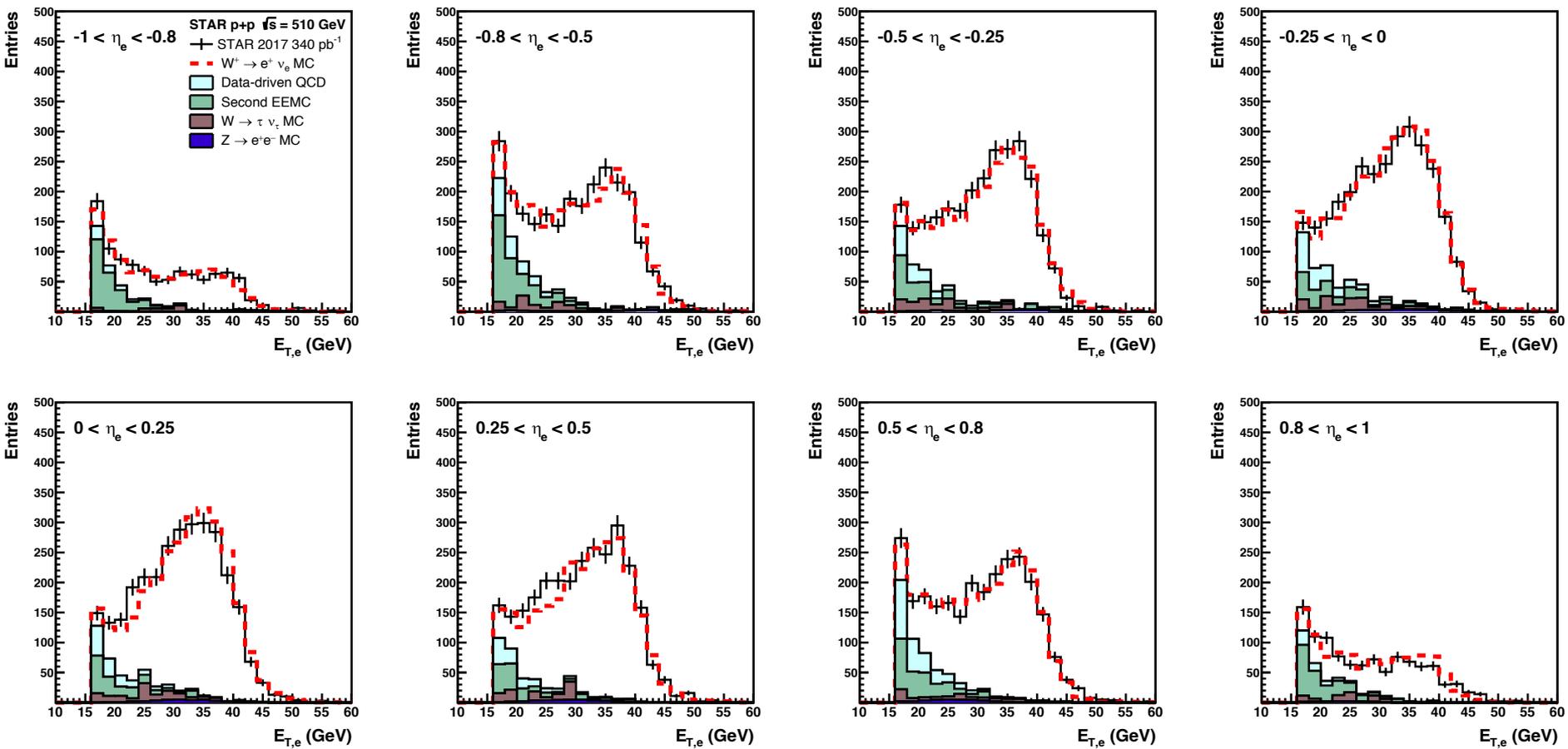


$$N_{sig}^{data} = N_{sel}^{data} - \sum N_{bg,i}$$

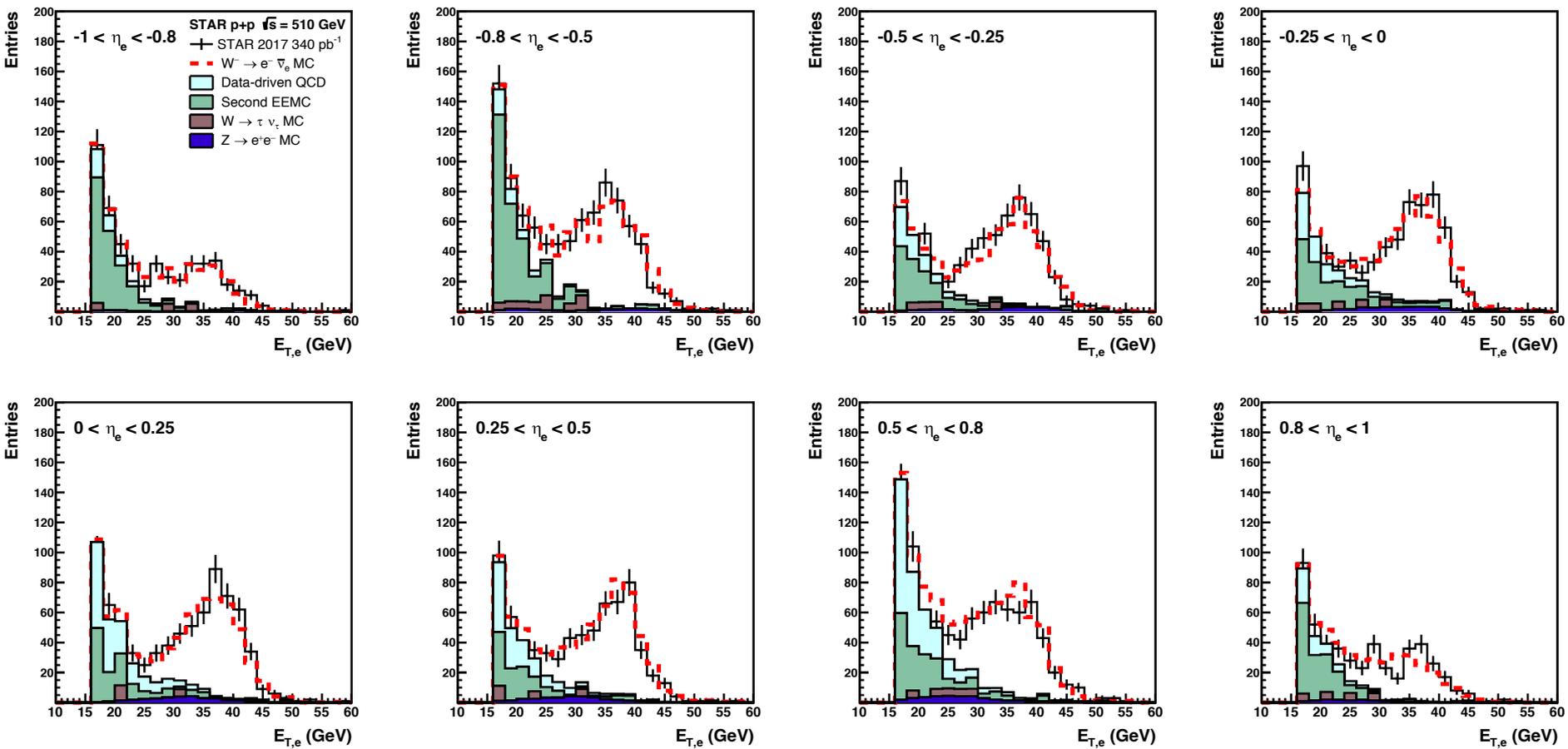
## Remaining background

- $Z \rightarrow e^+e^-$ 
  - with one electron outside STAR acceptance
  - Indistinguishable experimentally  
→ Estimated based on MC
- $W \rightarrow \tau$ 
  - $\tau \rightarrow e$  decay within STAR volume
  - Same as  $Z \rightarrow e^+e^-$  (MC-based)
- QCD processes (dijets)
  - One jet in dijet events outside STAR acceptance
  - Data-driven method  
(signed- $p_{T,bal} < \text{threshold}$ )
  - Using existing EEMC ( $-2 < \eta < -1$ )
  - Systematic uncertainty for data-driven QCD background selection and normalization

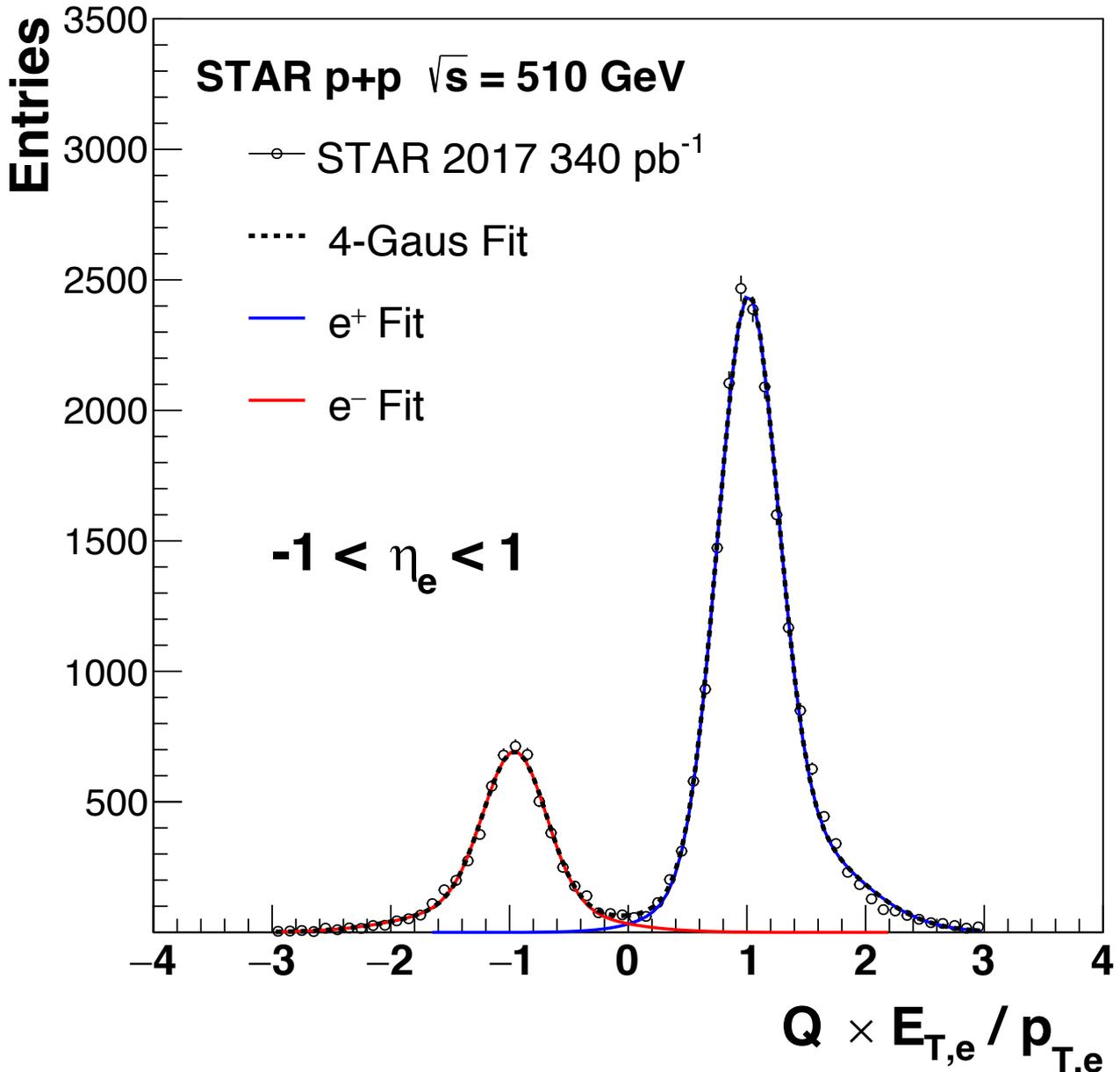
# Fig 2: Signal + BG description (W+)



# Fig 3: Signal + BG description (W-)



# Fig 4: Charge Mis-ID (Barrel, Comb)



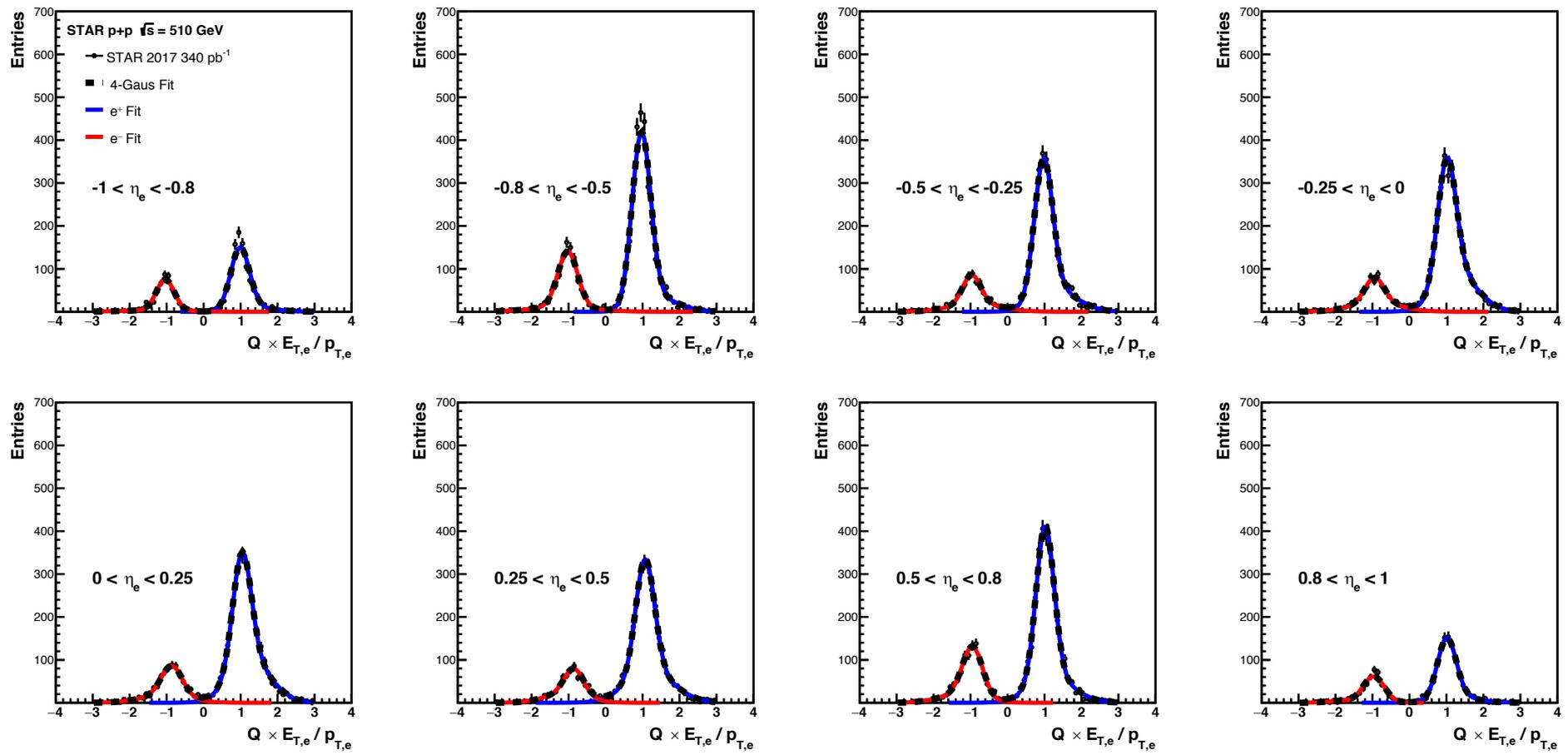
- Charge mis-ID correction based on 4-Gaus fit of  $Q \times E_T / p_T$
- 2-Gaus for each sign to describe signal and background

- $$c_q^+ = \frac{N^{fit+}}{N_{q>0}^{fit+} + N_{q>0}^{fit-}}$$

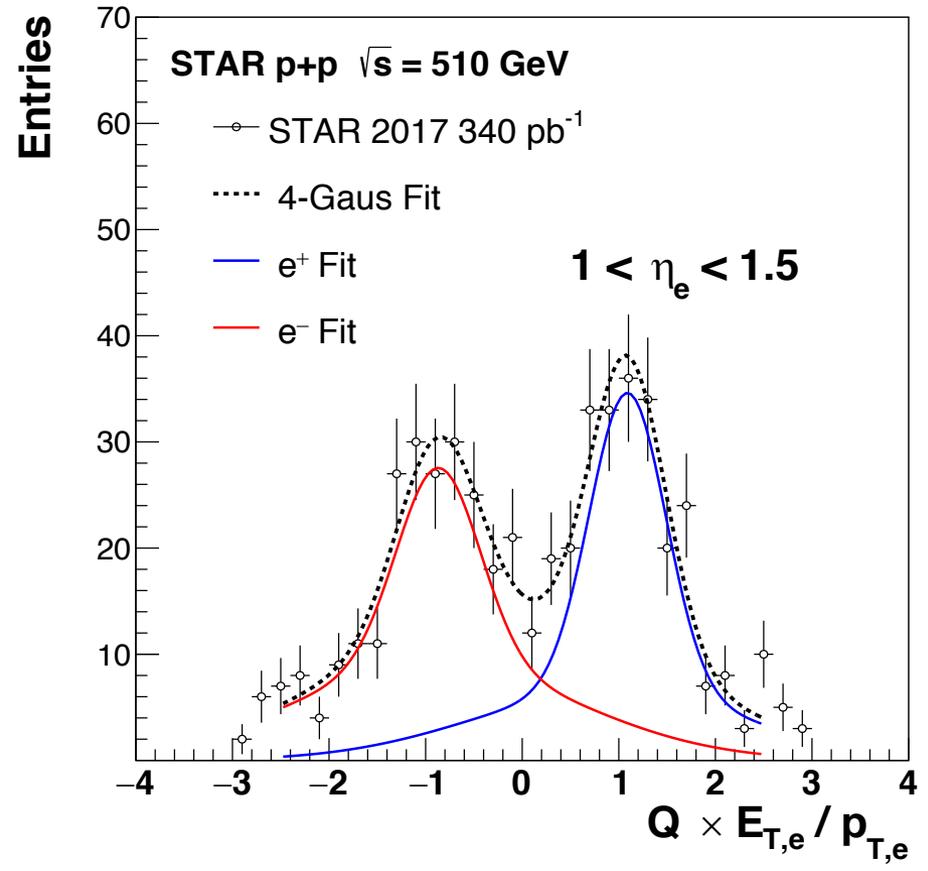
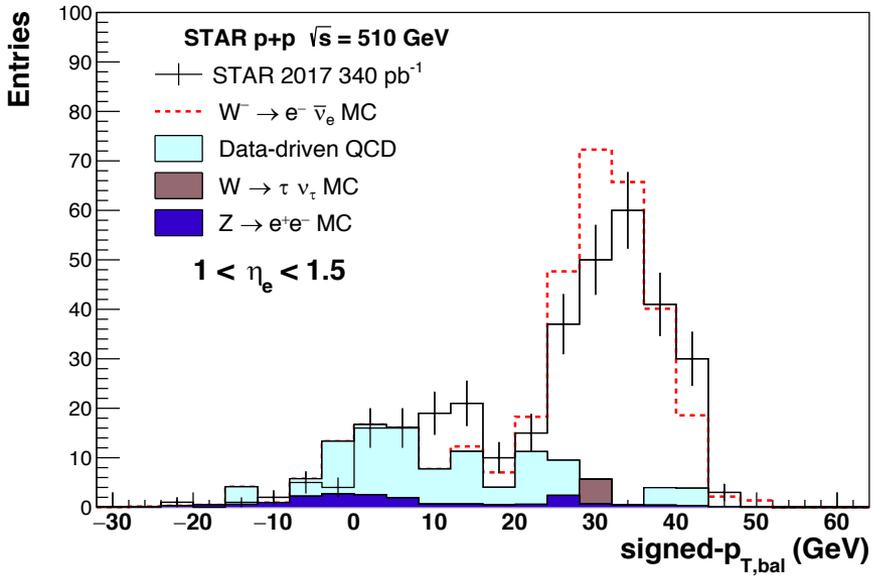
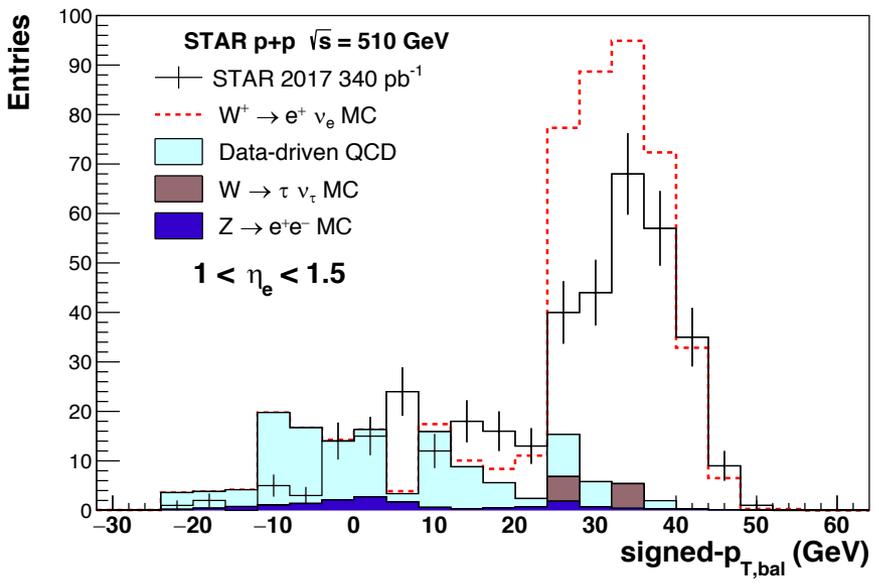
$$N_{sig,cor}^{data} = c_q \times N_{sig}^{data}$$



# Fig 5: Charge Mis-ID (Barrel)



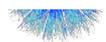
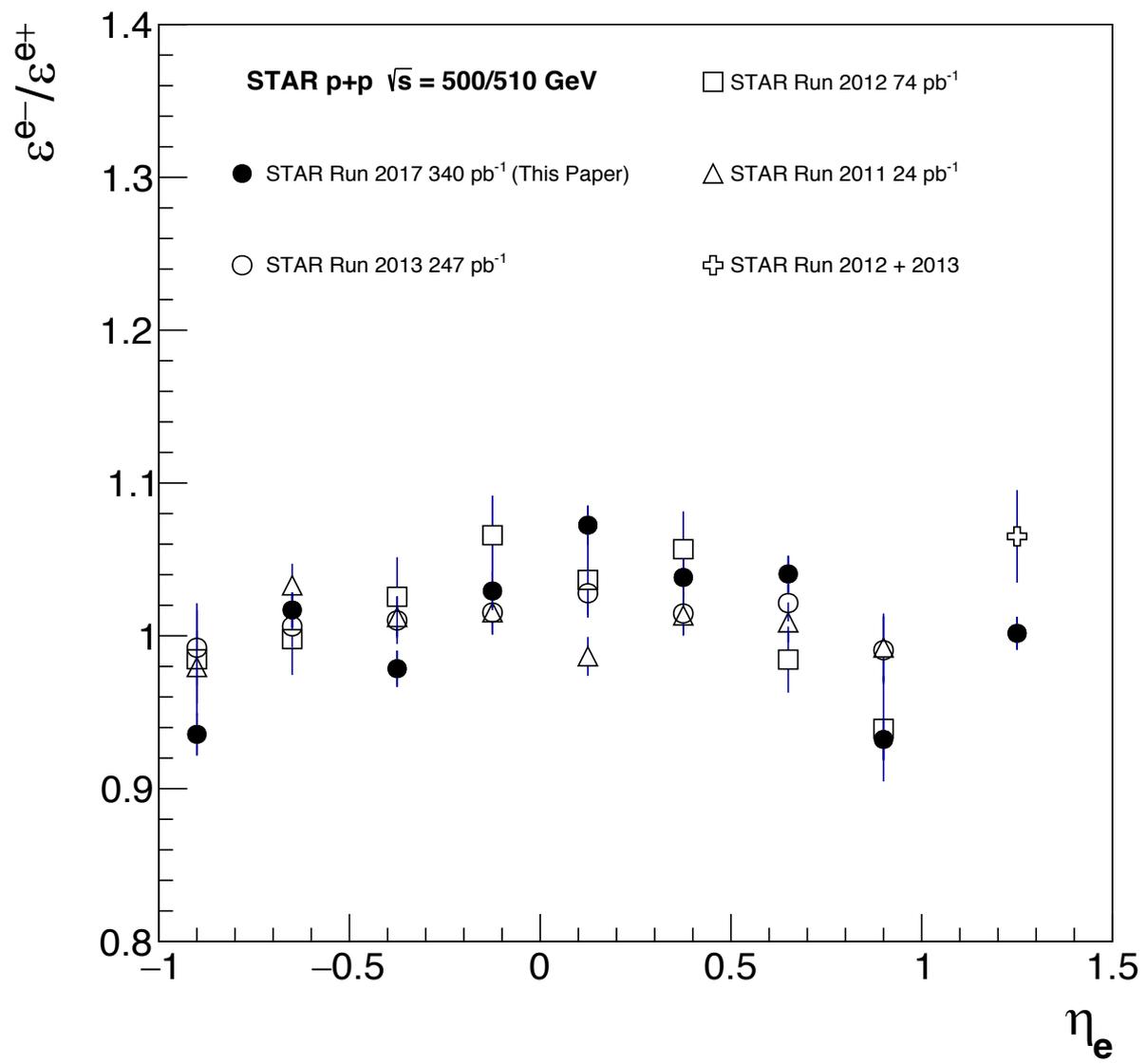
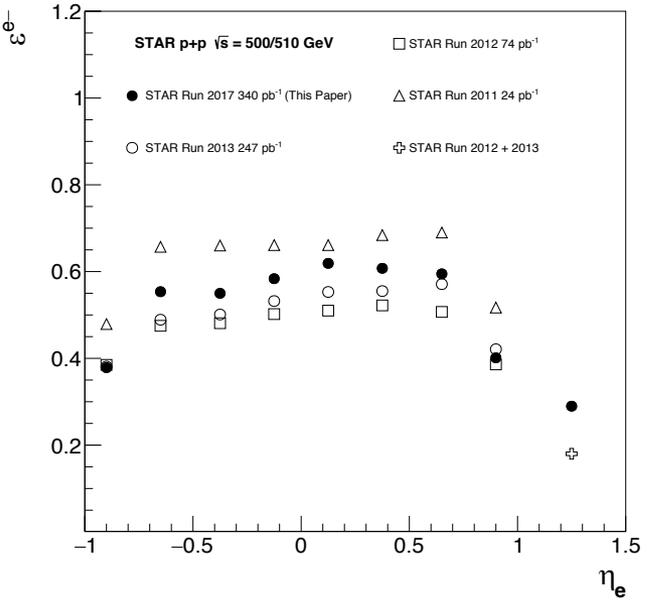
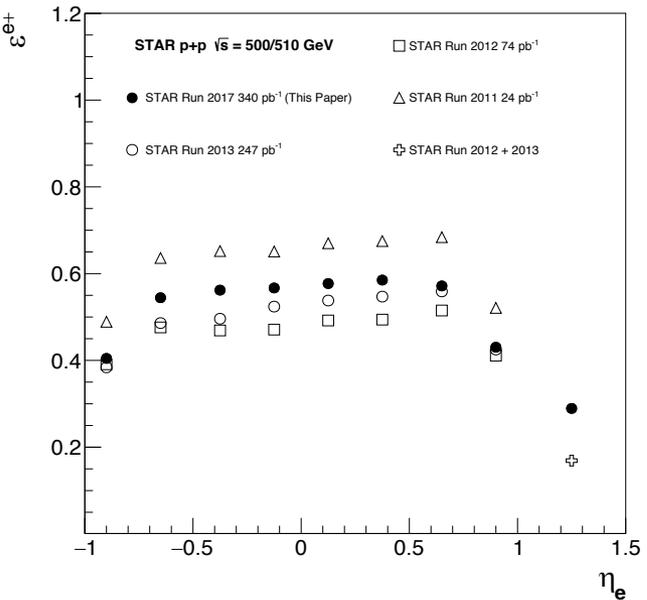
# Fig 6: Signal + Charge (Endcap)



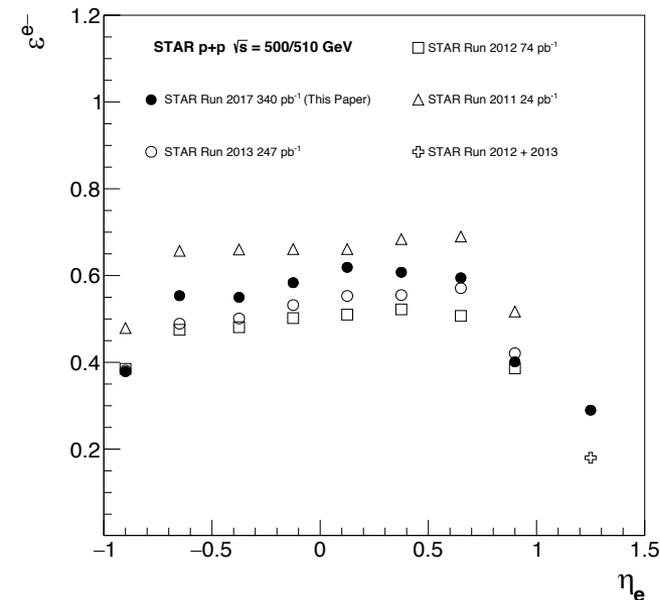
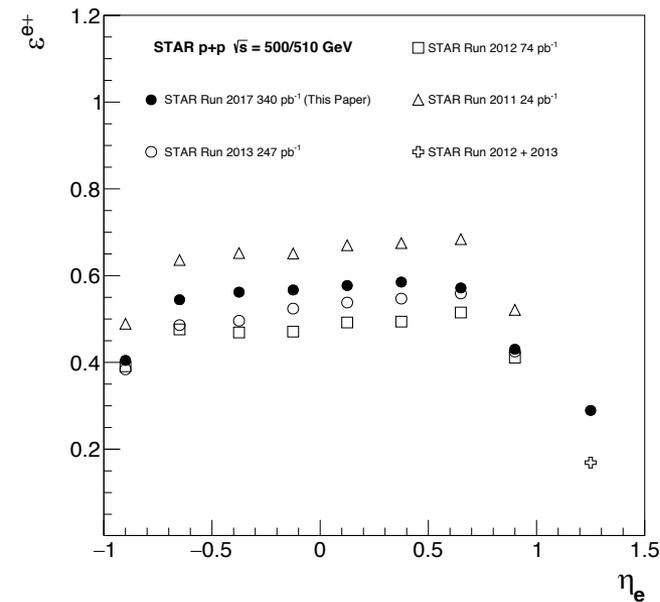
- MC overestimates W+, also seen in previous publication
- Systematics for data-driven QCD (same as barrel) and charge correction



# Fig 7: Efficiency



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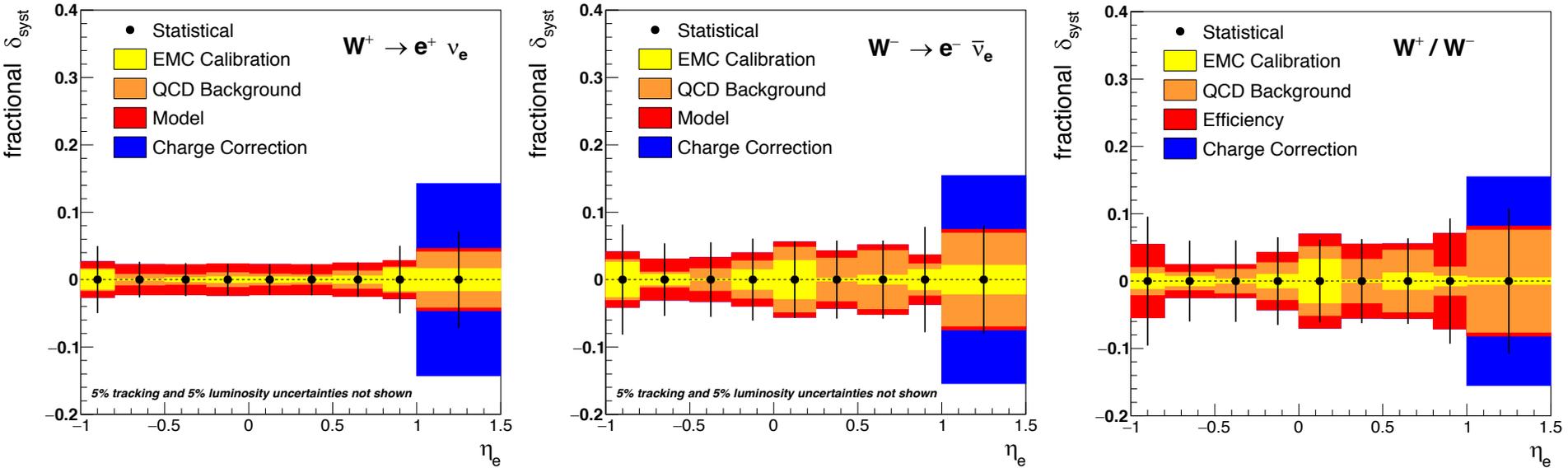


- Run 2011
  - Low material budget
  - Low instantaneous luminosity
- Run 2012 & 2013
  - High material budget (HFT, FGT, support)
  - High instantaneous luminosity
  - STI track algo Run 11, 12 vs STICA Run 13
- Run 2017
  - Low material budget
  - High instantaneous luminosity
  - STICA algo
- Asymmetric in  $\eta_e$  due to TPC Sec 20
  - Sec 20 issue addressed in Run 18 (→ Run 22)

$$\left(\frac{d\sigma}{d\eta}\right)_i = \frac{1}{\epsilon_i} \frac{N_{sig,cor,i}}{\delta\eta_{e,i}L}$$



# Fig 8: Systematics



- EMC Calibration

- BEMC/EEMC gains varied by reported uncertainties in calibration

- QCD Background

- Data-driven QCD background identification and normalization

- Charge Correction (Endcap only)

- Fit function and range

- Model

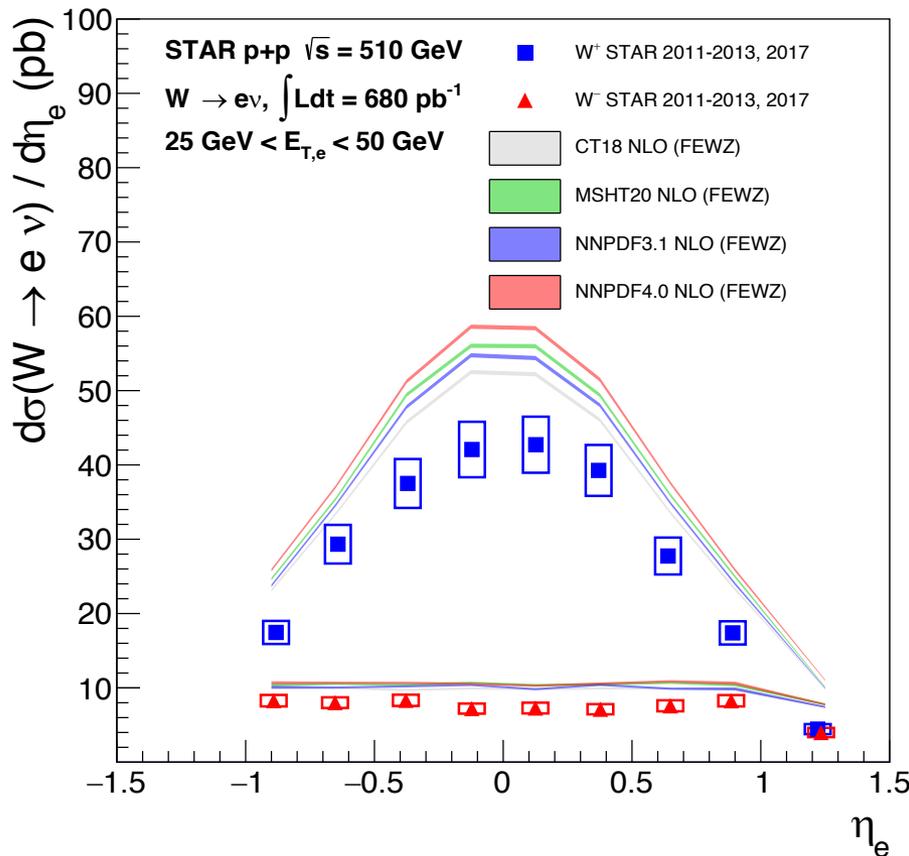
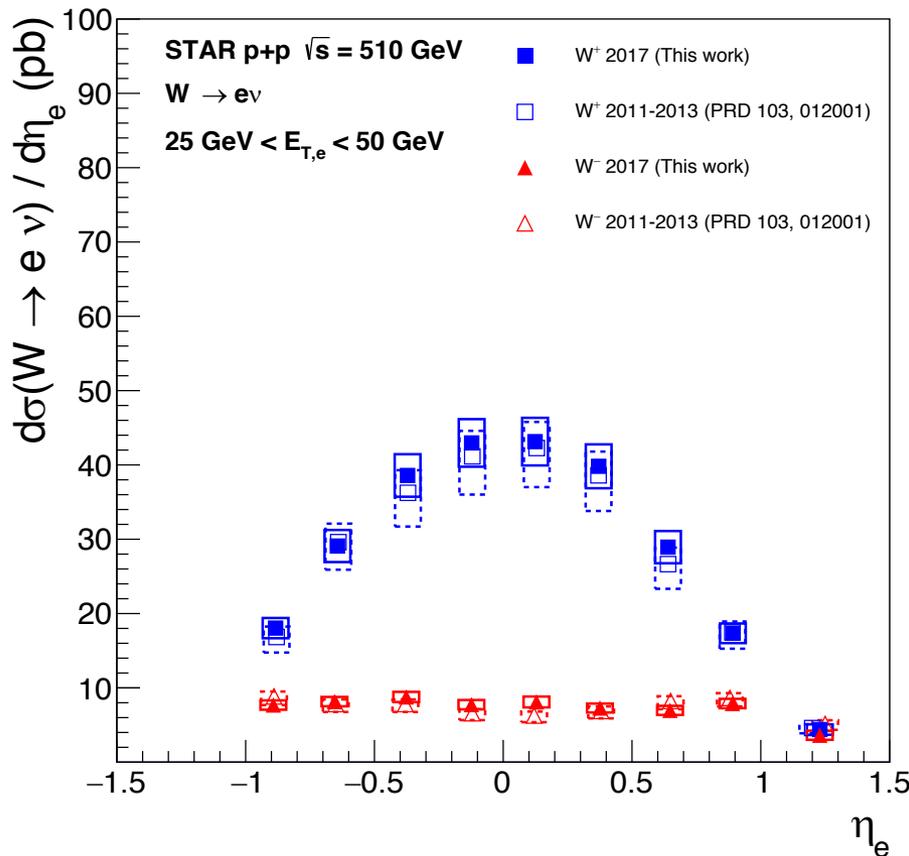
- Mis-estimated  $E_T^{\Delta R < 0.7}$ ,  $E_T^{\Delta \phi \sim \pi}$ , and signed- $p_{T,bal}$  selection efficiency
- $\sim 2.5\%$  correction and 2.5% uncertainty

- Efficiency (ratio only): deviation from 1

- Luminosity (cross section only)  $\sim 5\%$

- Tracking (cross section only)  $\sim 5\%$

# Fig 9 Cross Section

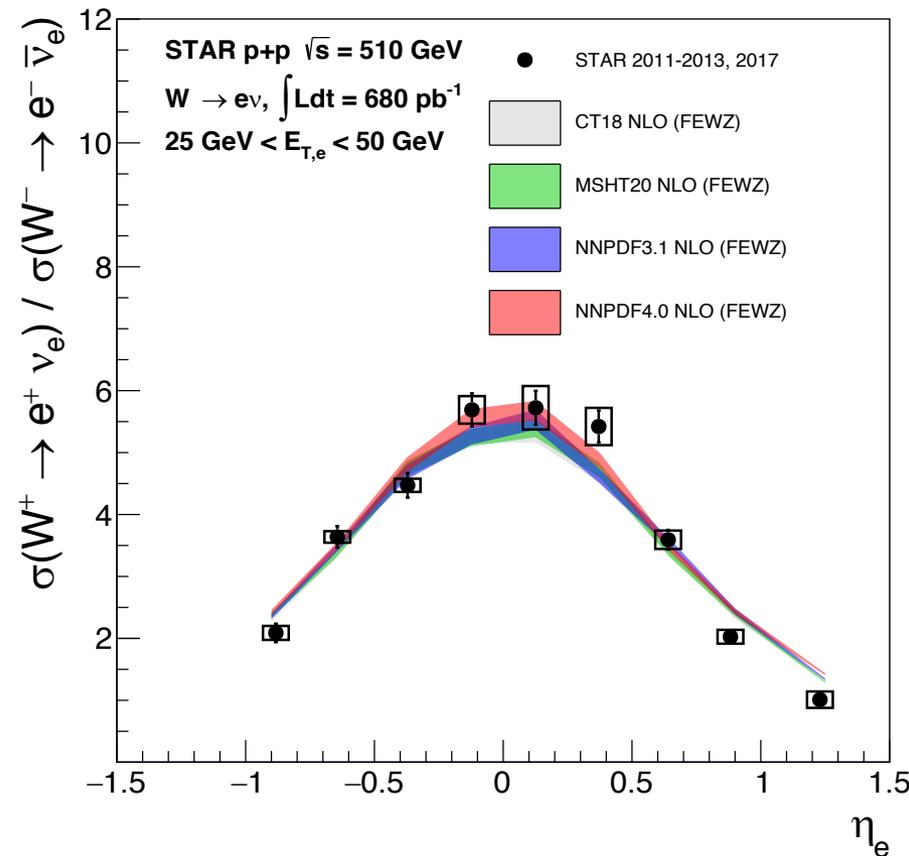
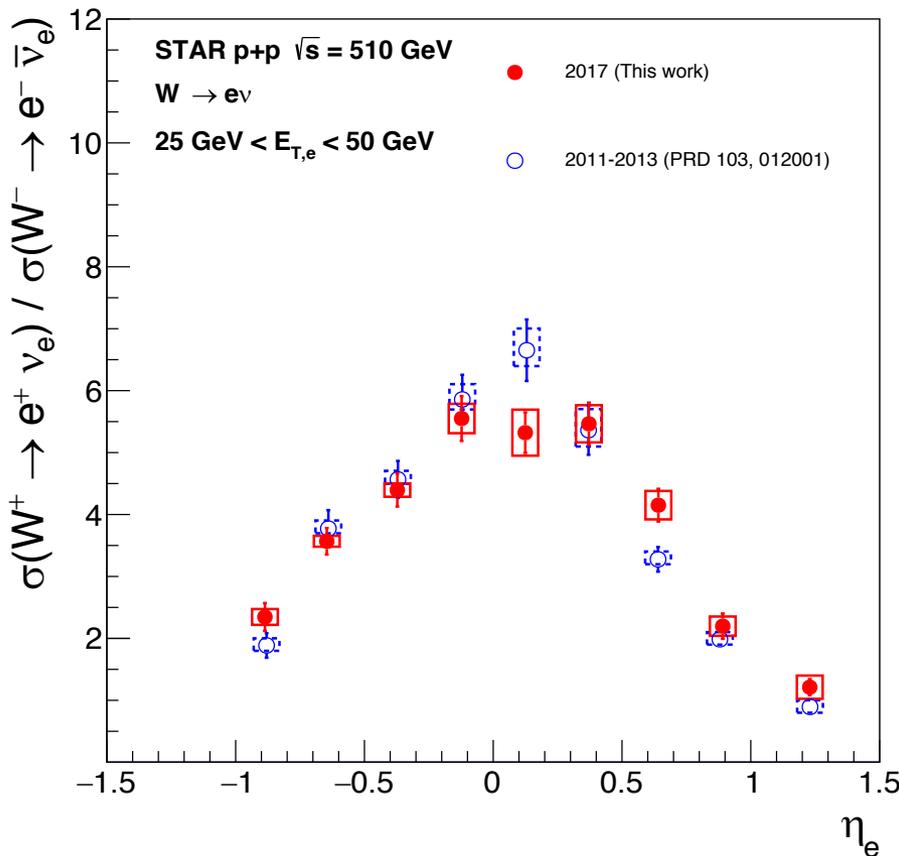


- ~3% shift between Run 17 and previous datasets based on  $\chi^2$  minimization
- Systematic uncertainties with strong bin-wise correlation
  - Luminosity (5%, strongly correlated btw datasets)
  - Tracking (5%, different detector conditions, changes in tracking algo, etc.)

- 30% scale mismatch between STAR and pQCD
  - ~10% uncertainty in data, spread btw. PDF sets
- Currently using FEWZ calculation
  - ResBos2 calculation in agreement with FEWZ
  - PDF dependence WIP



# Fig 10: Cross Section Ratio



- Run 17 result consistent previous STAR publication;  
 $p > 0.05$  based on T test, t value = 1.2199 with NDF = 9 - 1  $\rightarrow$  p value = 0.257242
- Modest improvement (expected) from fitting SeaQuest data  
 $(\chi^2 = 16.1 \rightarrow 14.9$  with NNPDF3.1  $\rightarrow$  4.0 in the barrel region from FEWZ)
- Tension in the endcap region still persists ( $\chi^2 = 7.0$  and 4.5 with NNPDF 4.0 and 3.1)



# 13 Conclusion

The  $W^\pm \rightarrow e\nu$  cross sections and their ratio have been measured with RHIC  $pp$   $\sqrt{s} = 510$  GeV collision data collected at STAR in 2017, corresponding to an integrated luminosity of  $340 \text{ pb}^{-1}$ . The measurements were performed as a function of the leptonic pseudorapidity  $\eta_e$  within the kinematic region defined by leptonic pseudorapidity  $-1.0 < \eta < 1.5$  and transverse energy  $25 \text{ GeV} < E_{T,e} < 50 \text{ GeV}$ . The resulting cross sections and ratio are consistent with previous STAR publications. These measurements provide independent constraints for the non-perturbative  $SU(2)$   $\bar{d}/\bar{u}$  asymmetry at a higher momentum scale ( $\sim M_W$ ), complementary to existing  $pp/pd$  Drell-Yan data, and a low- $\sqrt{s}$  and high- $x$  ( $\sim 0.2$ ) reference for LHC  $W/Z$  programs.

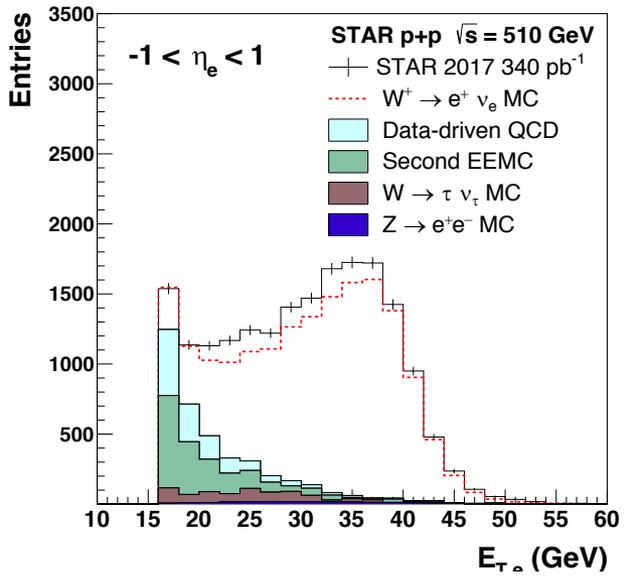
Perturbative calculations at NLO+NLL accuracy with latest PDF sets implemented in the ResBos2 framework [PLACEHOLDER] have been compared to the data. The prediction with the NNPDF4.0 PDF set [11] that includes the latest SeaQuest data [12] in the global fit shows an improvement over the previous NNPDF3.1 PDF set [13] in describing the  $W^+/W^-$  cross-section ratio observed at STAR. However, all calculations overestimate the differential cross sections  $d\sigma(W^\pm \rightarrow e\nu)/d\eta_e$  throughout the studied  $\eta_e$  range.

# Summary

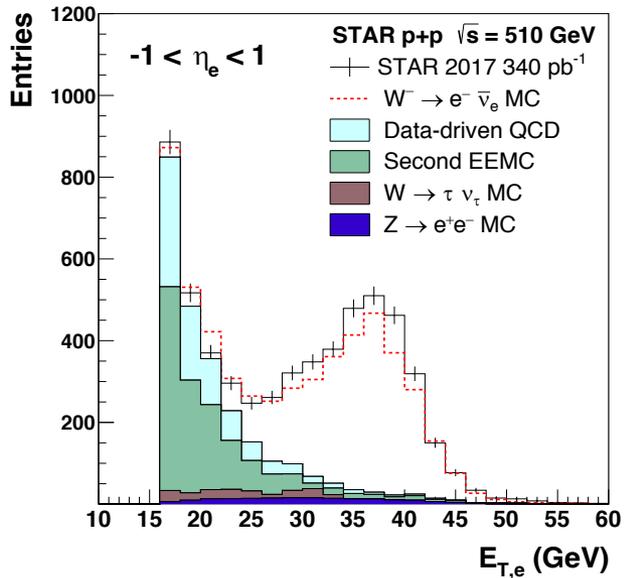
- Data well described by MC + background estimation
- Further corrections well understood
- Consistency between STAR datasets
- pQCD predictions overestimate STAR cross section by  $\sim 2\sigma$
- Modest improvement (expected) in description of STAR cross-section ratio by fitting SeaQuest (NNPDF3.1  $\rightarrow$  4.0)
- Tension in the Endcap region

# MC Normalization

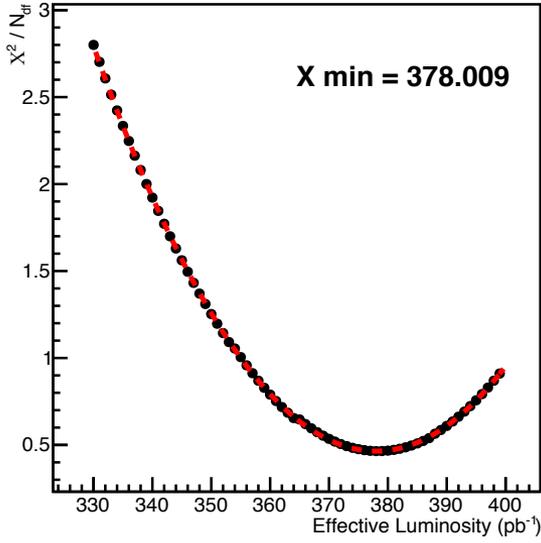
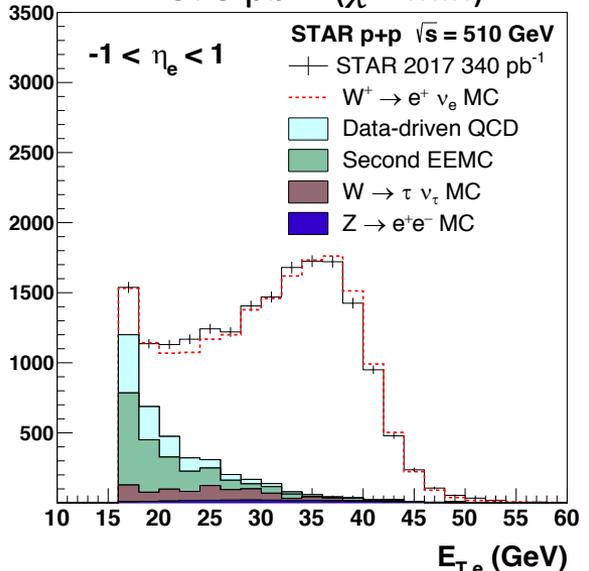
L = 343 pb<sup>-1</sup> (352 pb – 2.6%)



W<sup>+</sup> + W<sup>-</sup>



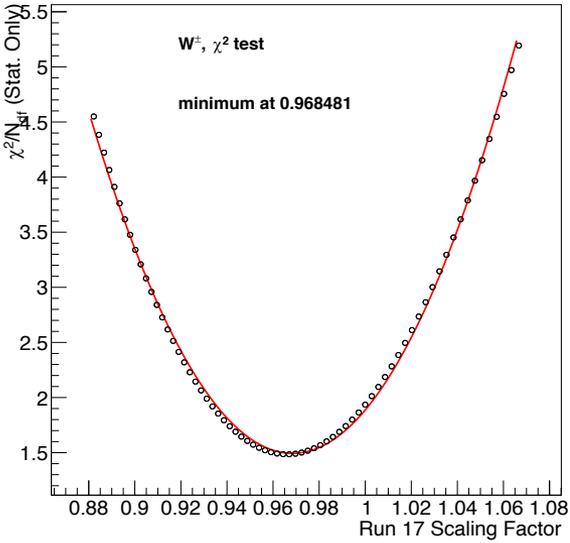
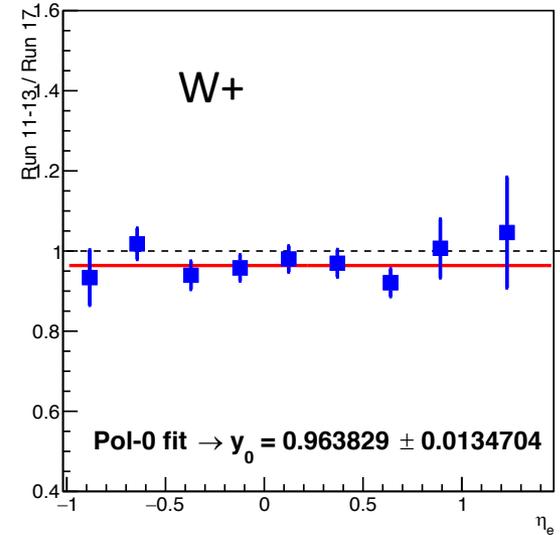
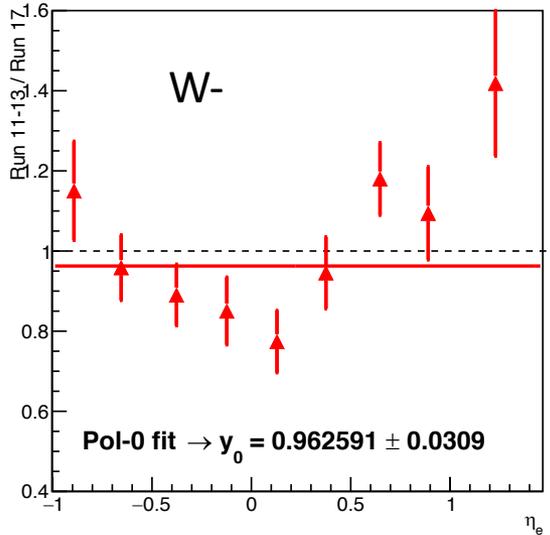
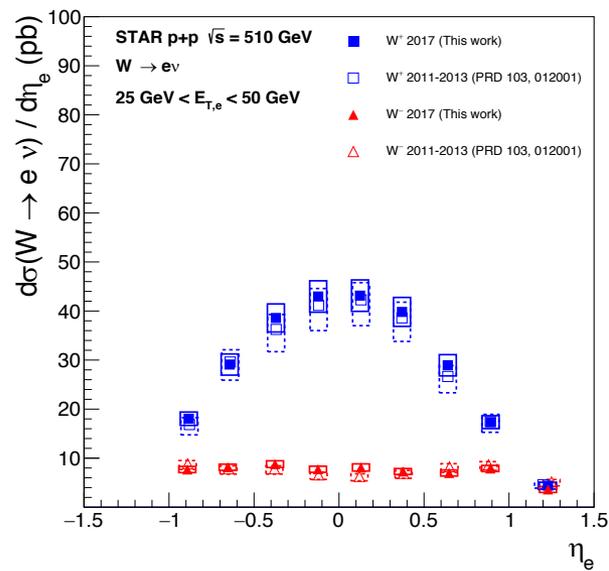
L = 378 pb<sup>-1</sup> ( $\chi^2$ -min)



- Luminosity based on Jamie's table 352 pb<sup>-1</sup>
- Systematic shift between MC and data
- Lumi correction + Changes in analysis details (MC gain smearing, etc.)
- $\chi^2$  test minimizing effective luminosity = 378 pb<sup>-1</sup>
- Well within (lumi + track) 352 ( $\pm 9\% \pm 5\%$ ) pb<sup>-1</sup>
- Around 1.4 $\sigma$  within 343 ( $\pm 5\% \pm 5\%$ ) pb<sup>-1</sup>
- Dropping claims about "MC normalized to data luminosity"
- Instead, "MC normalization based on  $\chi^2$  minimization"



# Systematic shift between Run 17 and 11-13



- 2.6% luminosity correction included for both Run 17 and 11-13
- $> 2\sigma$  shift based on pol-0 fit of  $W^+$
- $\chi^2$  scan (based on statistical uncertainty) suggests there is 3.2% systematic shift between Run 11-13 and Run 17



# Backup



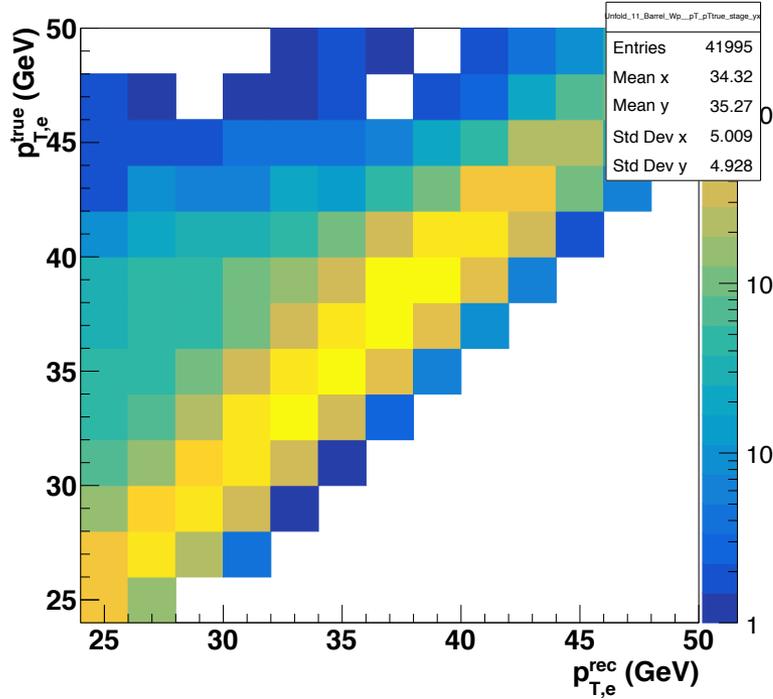
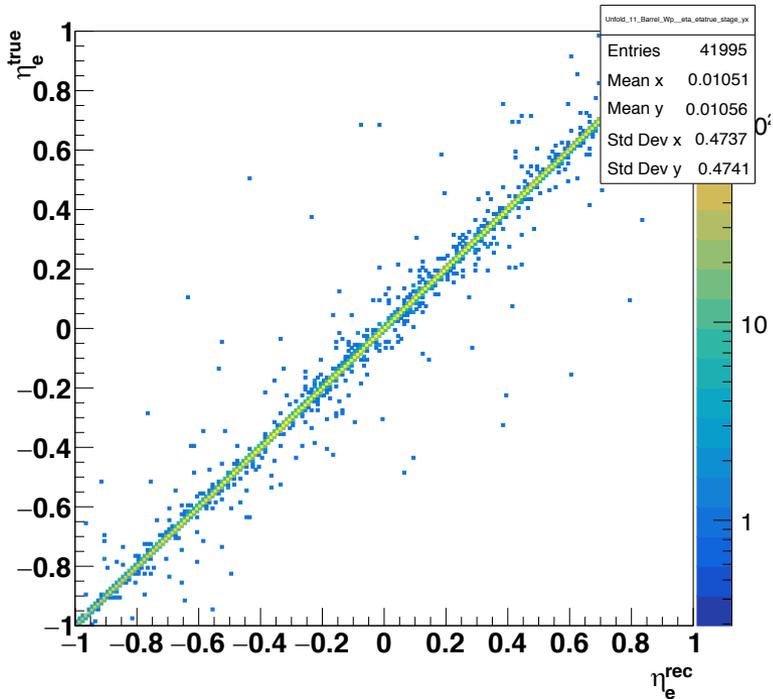
6/25/25

Jae D. Nam

25



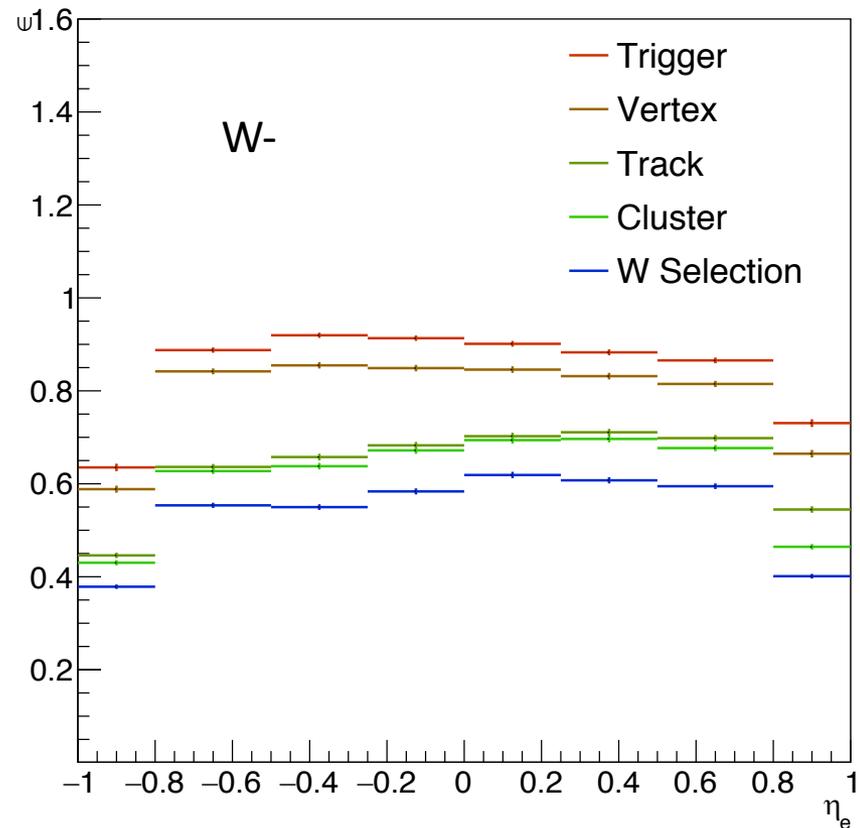
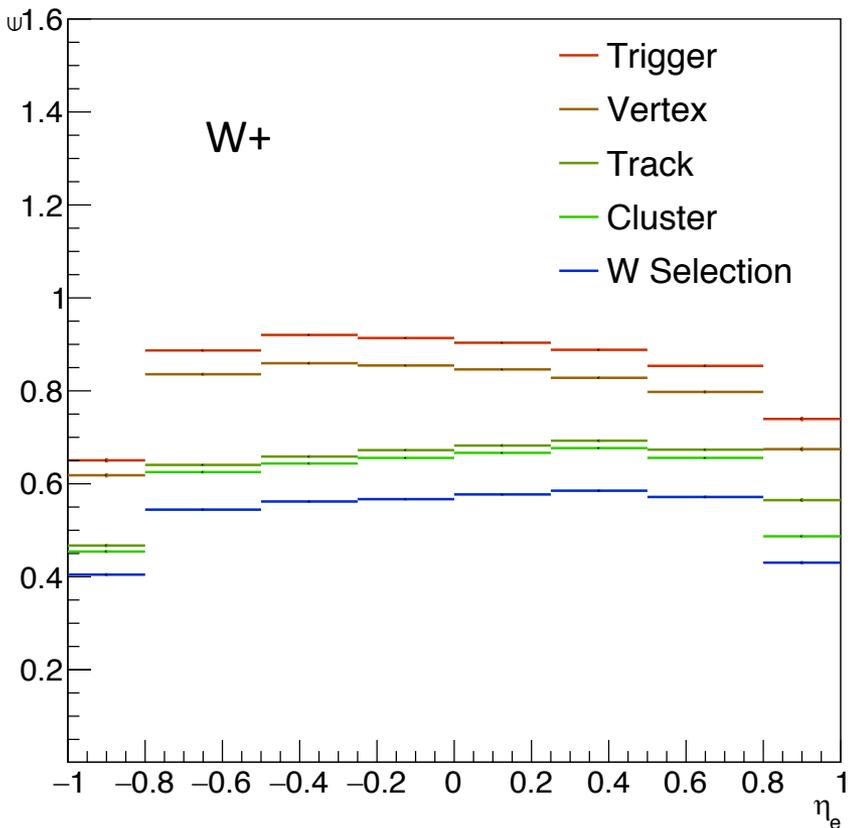
# Bin migration correction?



- Good reconstruction of  $\eta$  and  $E_T$
- Binwise correction factor  $\sim 0.5\%$
- Ignored; not included in the calculation, and was not considered as systematics

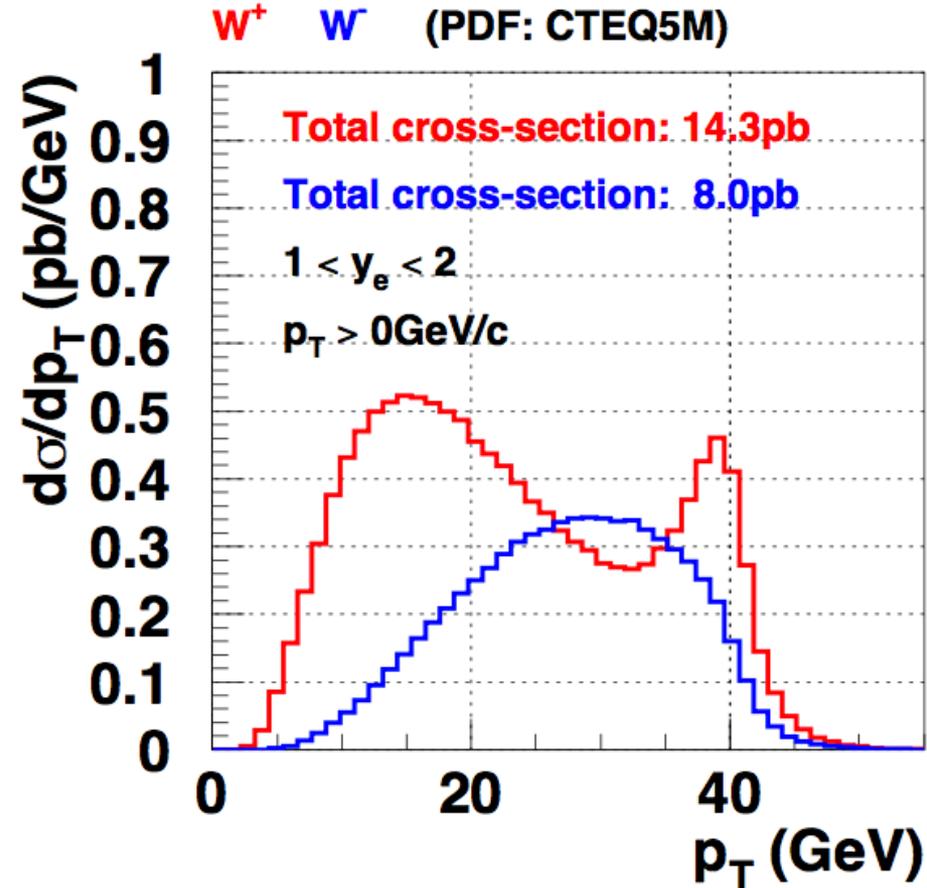
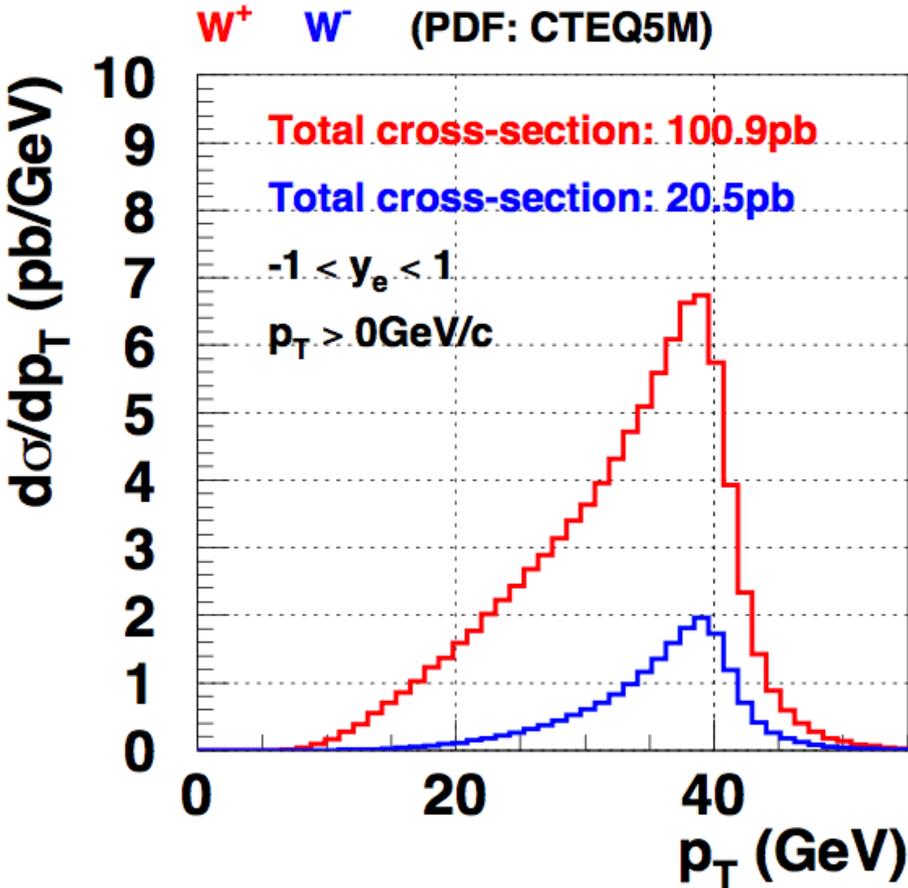


# Efficiency breakdown



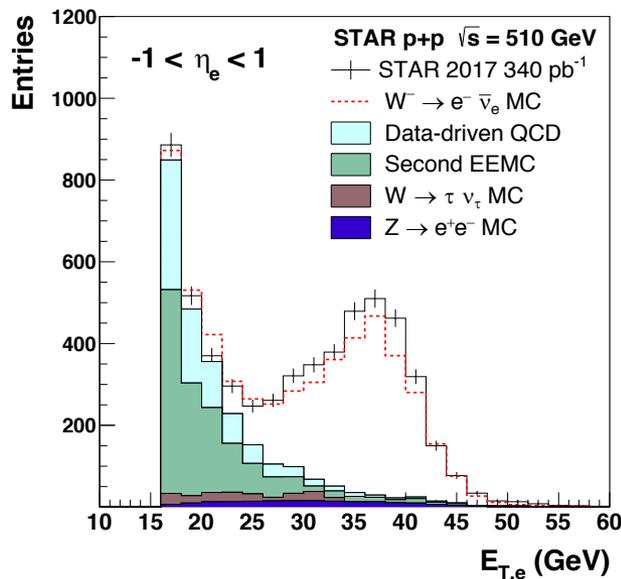
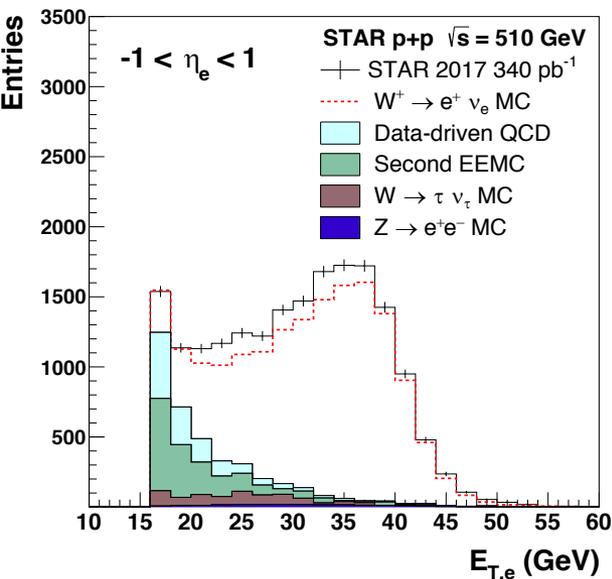
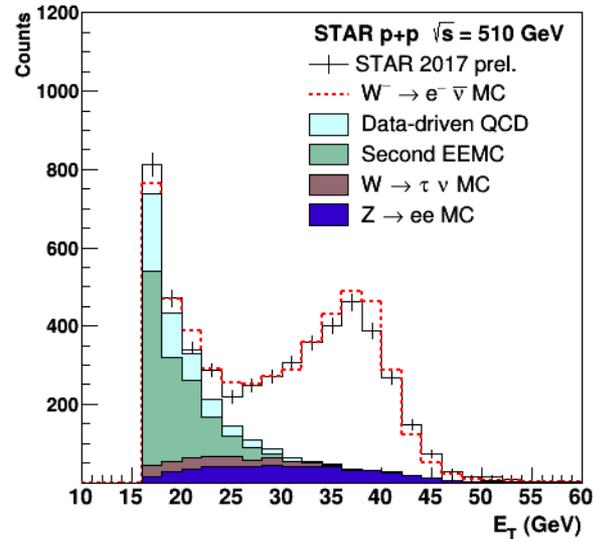
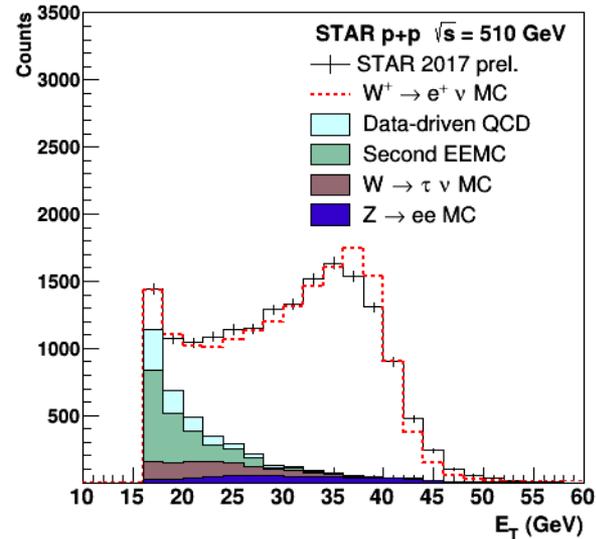
# Why not one value for certain syst.?

P. M. Nadolsky and C.P. Yuan,  
Nucl.Phys. B666 (2003) 31.



- The shape of Jacobian peak, and thus EMC/Efficiency systematics, should depend on  $\eta_e$

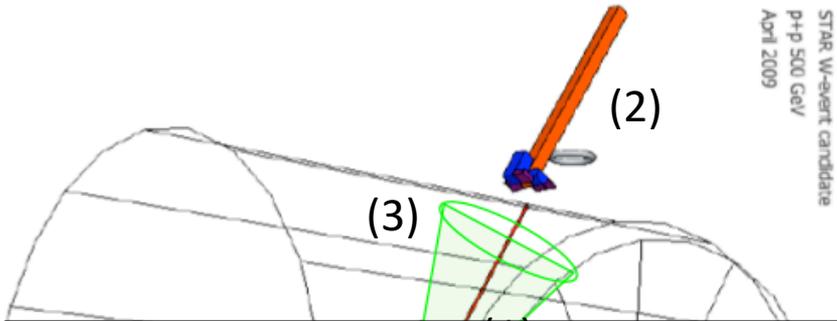
# Comparison to Preliminary



- Top: Preliminary
  - MC normalized to  $352 \text{ pb}^{-1}$
- Bottom: Publication
  - MC normalized to  $343 \text{ pb}^{-1}$
  - Includes MC gain smearing of 5%
  - Software-level inconsistency corrections (analysis note)

# Review of Event Reconstruction

Event Display of typical W event



$E_T^{\Delta R < 0.7}$  and  $E_T^{\Delta \phi \sim \pi}$  are formed with track momentum and EMC tower energy ( $p$  of particles reconstructed in detector)

(4)  $p_{T,bal}$  is reconstructed with objects (jets) reconstructed with anti-kT clustering algorithm (E-scheme,  $R = 0.6$ ,  $\min-p_T = 3.5$  GeV)

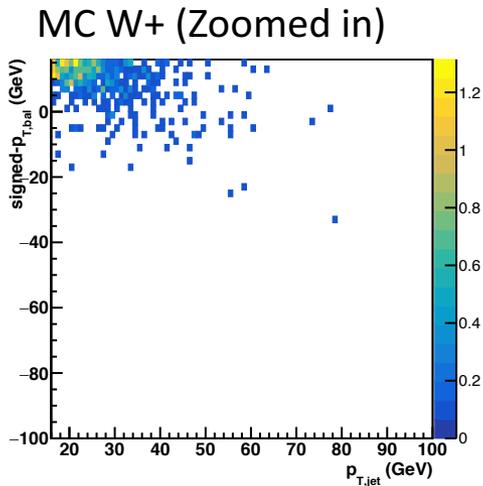
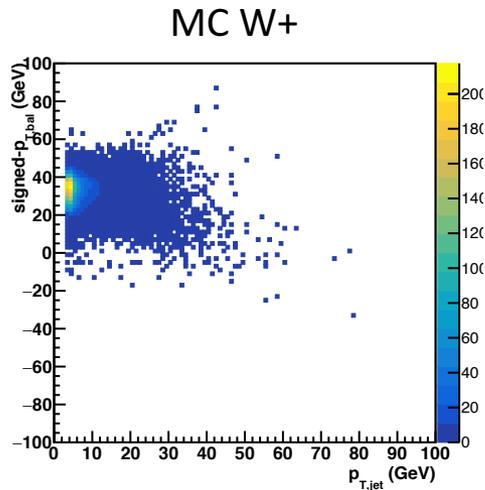
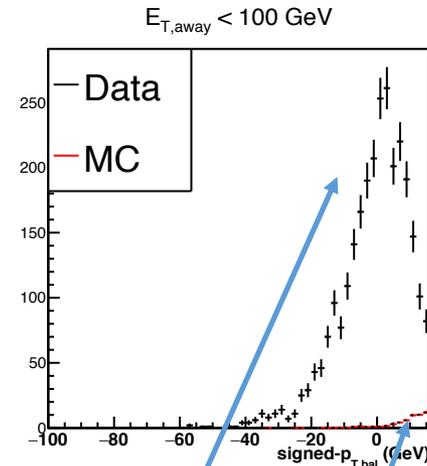
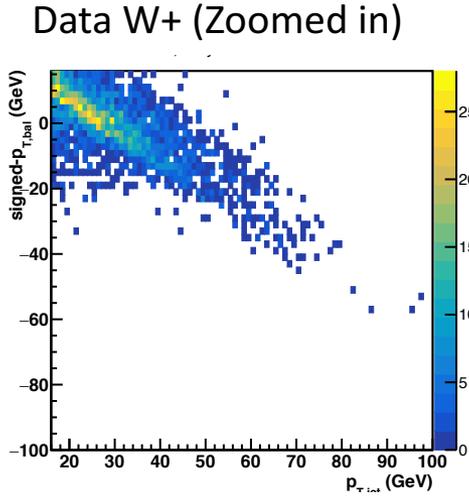
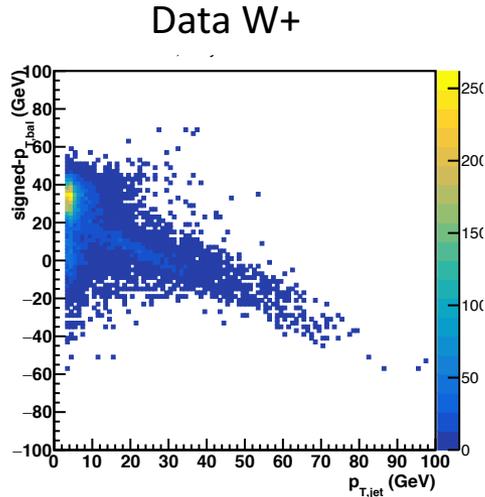
Relevant studies in:

- [https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam\\_20250326\(1\).pdf](https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam_20250326(1).pdf)
- [https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam\\_20250409.pdf](https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam_20250409.pdf)
- [https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam\\_20250423\(1\).pdf](https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam_20250423(1).pdf)
- [https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam\\_20250604.pdf](https://drupal.star.bnl.gov/STAR/system/files/userfiles/6368/Nam_20250604.pdf)

- 1) A high momentum track is identified ( $p > \text{threshold}$ )
- 2) Energy cluster is formed (2x2 towers, each covering  $0.05 \times 0.05$  in  $\eta \times \phi$ ; cluster:  $0.1 \times 0.1$ )
- 3) Isolation requirement ( $E_T^{2 \times 2} / E_T^{\Delta R < 0.7} \sim 1$ ,  $E_T^{2 \times 2} / E_T^{4 \times 4} \sim 1$ )
- 4) Backward (neutrino direction) energy flow requirement ( $E_T^{\Delta \phi \sim \pi} < \text{threshold}$ )
- 5) Energy imbalance ( $p_{T,bal}$ ) ( $\hat{p}_{T,e} \cdot \sum [\vec{E}_T \text{ and } \vec{p}_T] > \text{threshold}$ )



# Signed- $p_{T,bal}$ and Jet $p_T$ (W+)

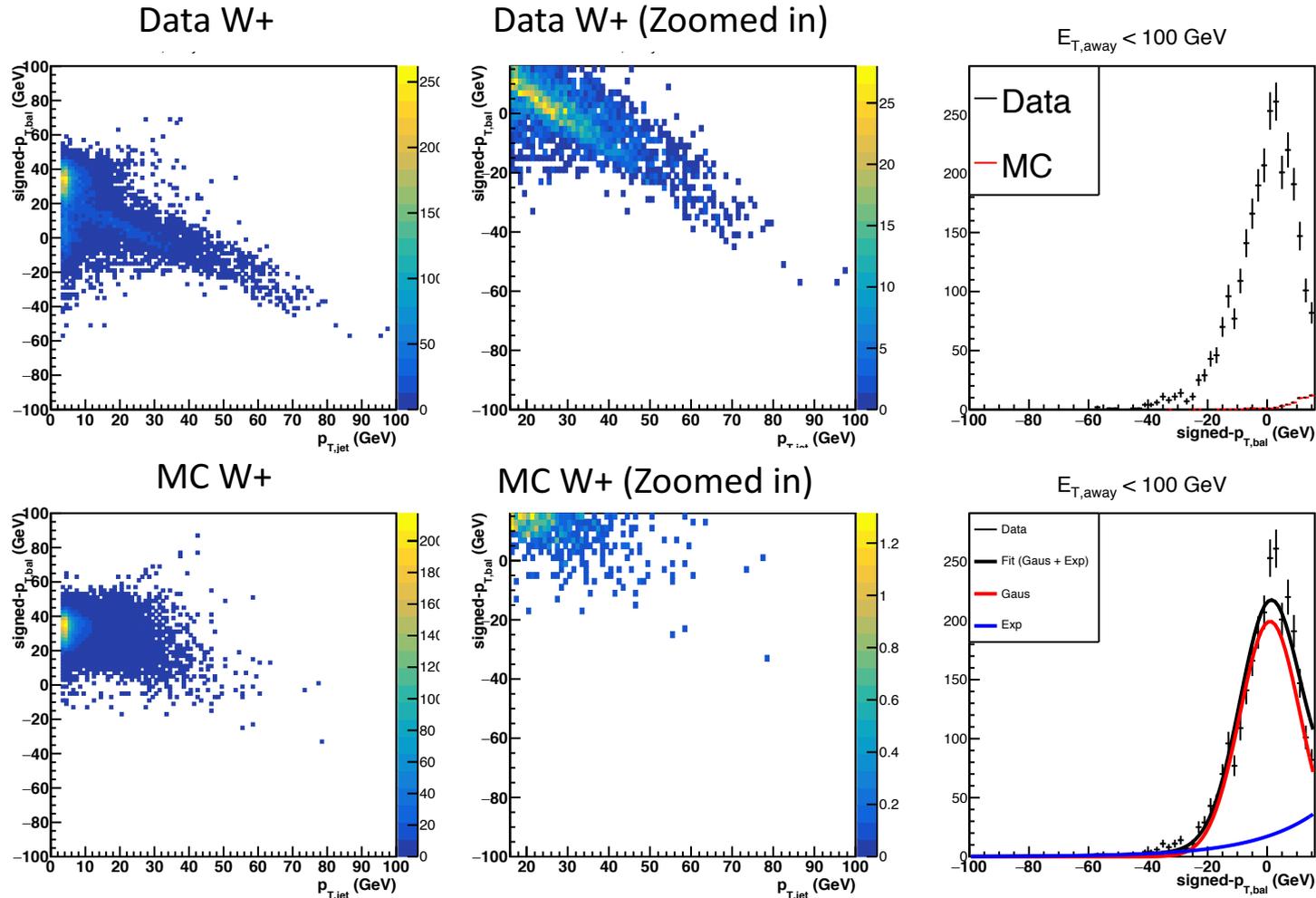


Asymmetric low- $p_{T,bal}$  tail of W

Symmetric contribution dominated by Z, QCD-dijet

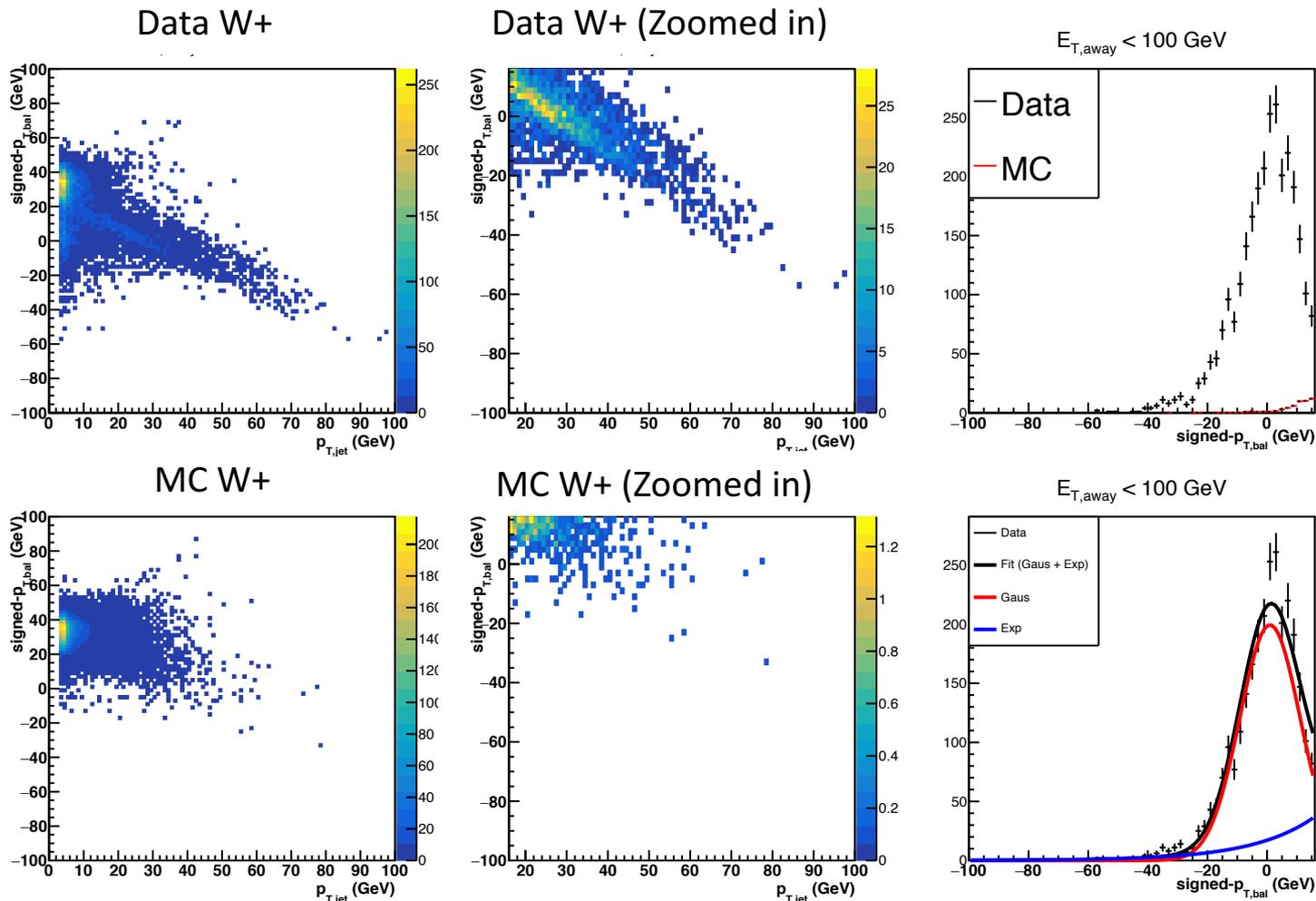
- Looking at signed- $p_{T,bal} < 16$  GeV,  $p_{T,jet} > 16$  GeV

# Signed- $p_{T,bal}$ and Jet $p_T$ ( $W+$ )



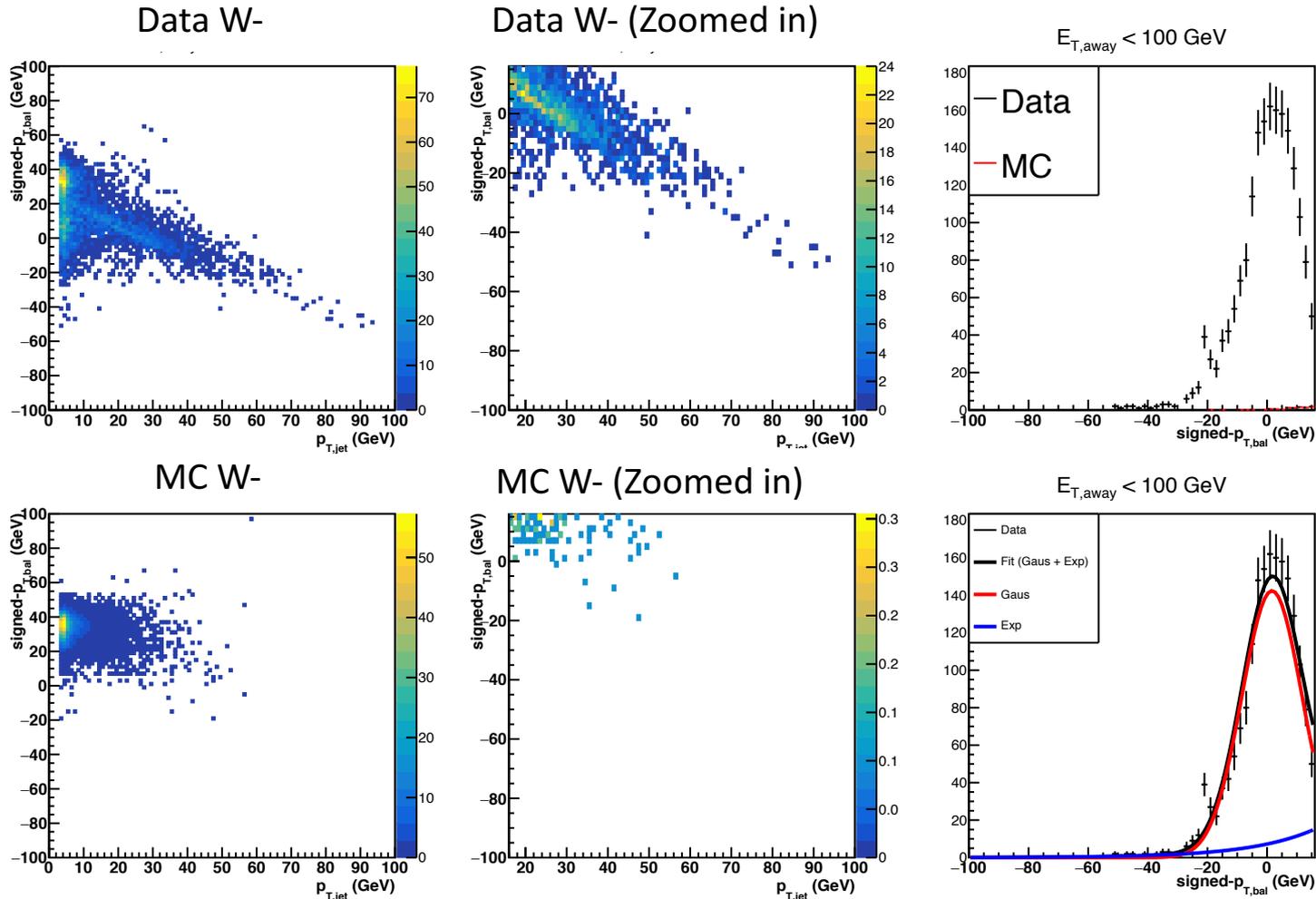
- Symmetric (QCD/Z) and asymmetric (possible  $W \rightarrow e$  or  $W \rightarrow \tau$ ) fits
- No justification behind choice of fit function (Gaus, Exp)

# Signed- $p_{T,bal}$ and Jet $p_T$ (W+)



- Asymmetric contribution can be as large as  $\sim 5\%$  of signal

# Signed- $p_{T,bal}$ and Jet $p_T$ (W-)



- Similar observation with W-  
(Asymmetric contribution can be as large as  $\sim 5\%$  of signal)

# Summary of Model Uncertainty

- Mock  $W$  study with new QCD jet definition suggests **no significant** mis-estimation of signed- $p_{T,bal}$  cut
  - Similar study was repeated with  $W$  data and found that the mis-estimation effect can **only be as large as**  $\sim 5\%$
- No significant contribution from signed- $p_{T,bal}$  and  $E_{T,away}$
- The following lines will be added In the text,
    - Emission of soft and hard gluons simulated with LO+PS models, such as Pythia, is subject to uncertainties originating from the underlying assumptions behind parton showering algorithms, and cannot be trusted entirely.
    - The resulting misdescription of soft and hard QCD radiation may appear as an incorrect estimation of the reconstruction efficiency in equation (X).
    - The inefficiency of selection criteria based on the activity within a large area of the detector, such as  $E_T^{\Delta R < 0.7}$  and  $E_T^{\Delta\phi \sim \pi}$ , and the description of reconstructed jets, such as **signed- $p_{T,bal}$** , have been re-evaluated using reconstructed  $W$  and  $Z$  samples from data without these requirements.
    - The impact of the misestimation of the reconstruction efficiency to the overall cross section was found to be less than 5% and taken as a **systematic uncertainty**.